



Welcome to The Natural World N190: Stars and Galaxies

Dale E. Mais

Pharmacologist by trade-Biotechnology

-Women's health care

-Astronomy is my hobby

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Welcome to Astronomy

AST203:

Stars and Galaxies

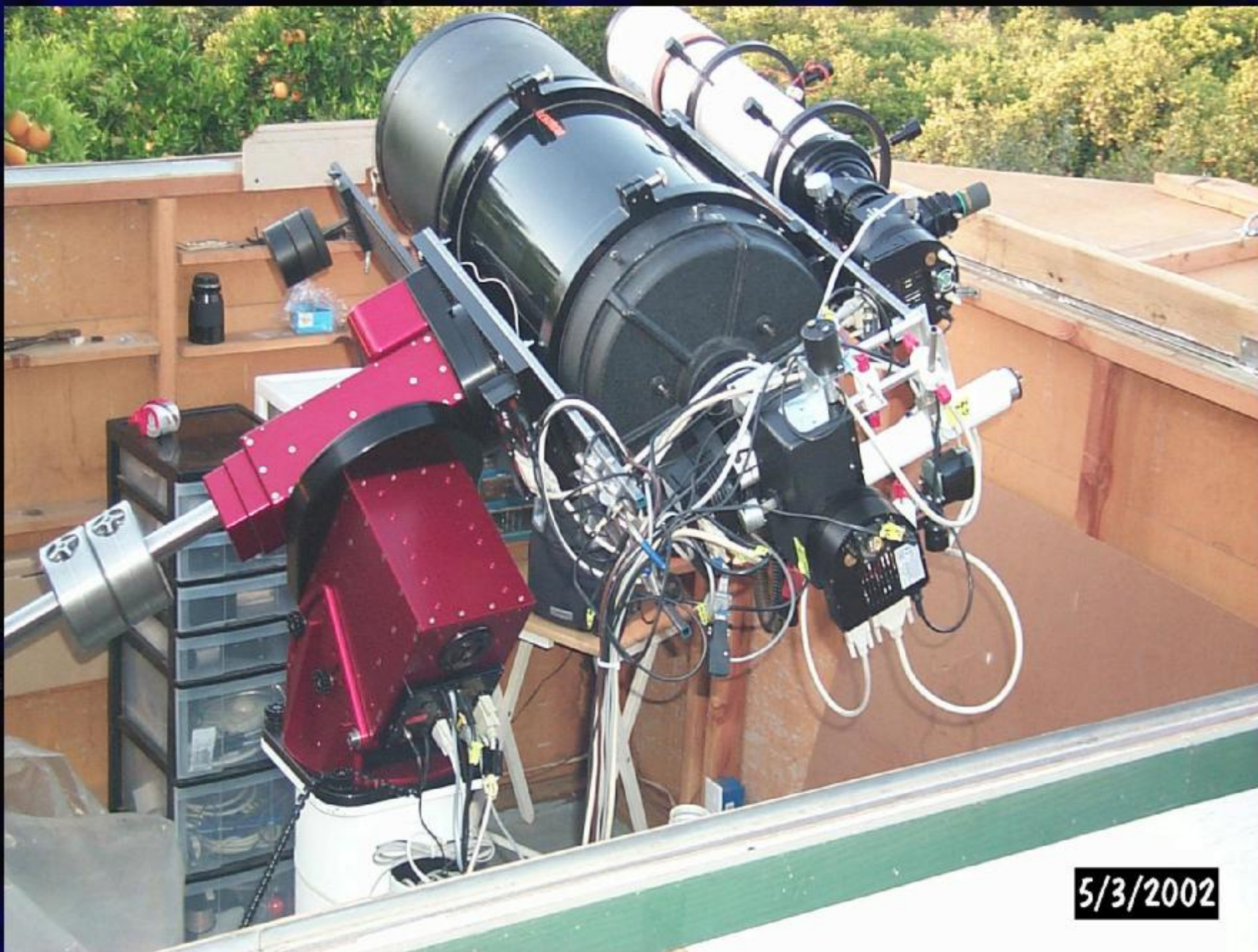
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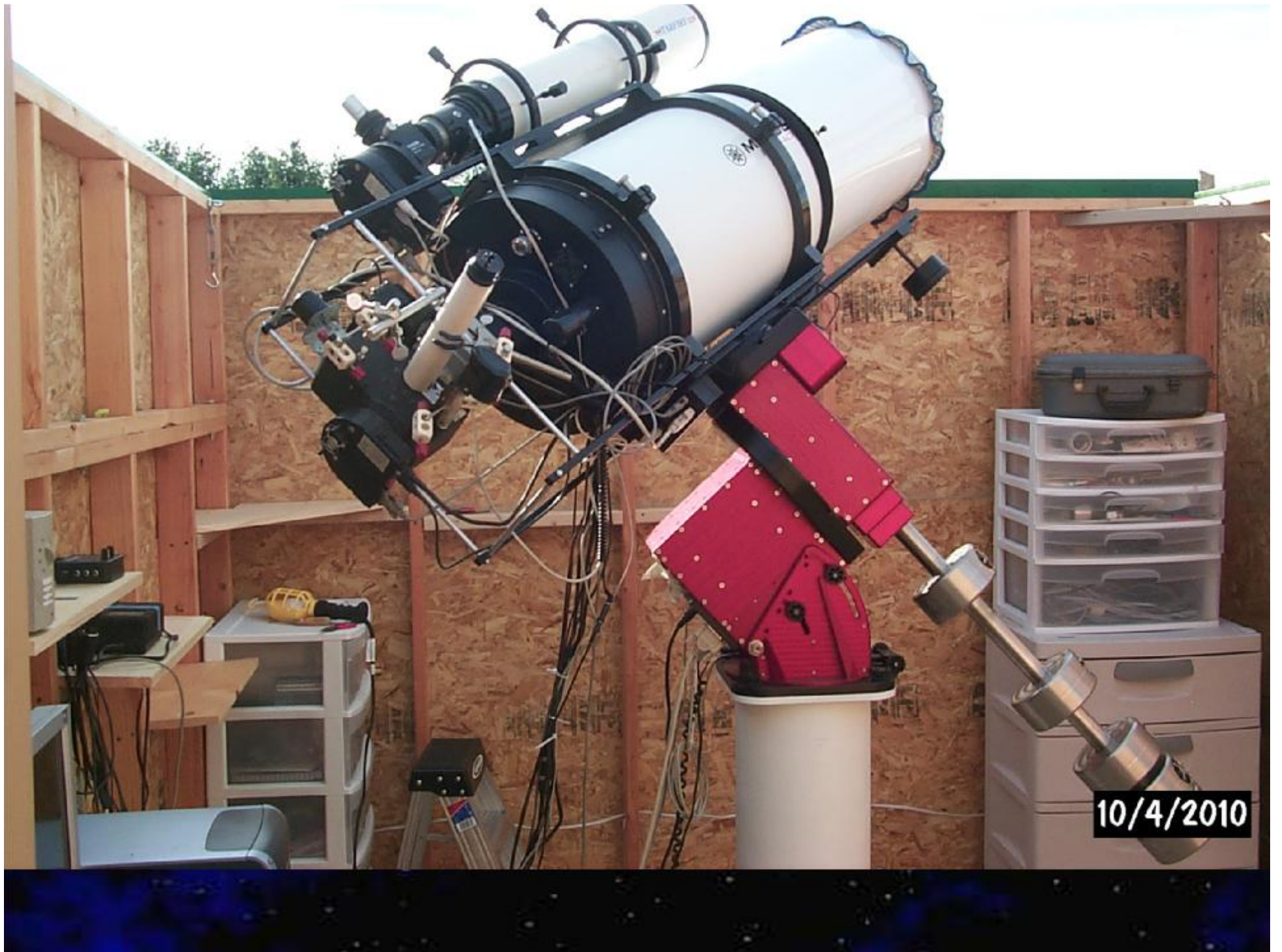
-Women's health care

-Astronomy is my hobby

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5/3/2002



The Natural World N190 COURSE OUTLINE

Spring 2012

1. Organization of the Sky

The Sky at Night	
Seasons / Lunar Phases / Eclipses	
Geocentric/Heliocentric Cosmology	

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2. Geocentric to Heliocentric transition The rise of the Scientific Process

Copernicus – Brahe	
Kepler's Laws	
Galileo	
Newton's Laws	

3. How we gather information: The tools of the trade

The Nature of Light – Radiation	
The Nature of Light – Spectroscopy	
The Tools of Astronomy	



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4. The building blocks of galaxies: Stars, how they live and die

The Nature of Stars	
The Interstellar Medium	
Stellar Formation	
Stellar Evolution	
Stellar Death	
Neutron Stars/Black Holes	

5. The building blocks of Universe: Galaxies

The Milky Way Galaxy	
Normal Galaxies	
Types of galaxies, A universe of galaxies	

All Powerpoint lecture slides are posted at my website
as pdf files

Download and print if you wish
Take notes on
Can use on exams/quizes

Mais-ccd-spectroscopy.com

Go to links for Palomar College

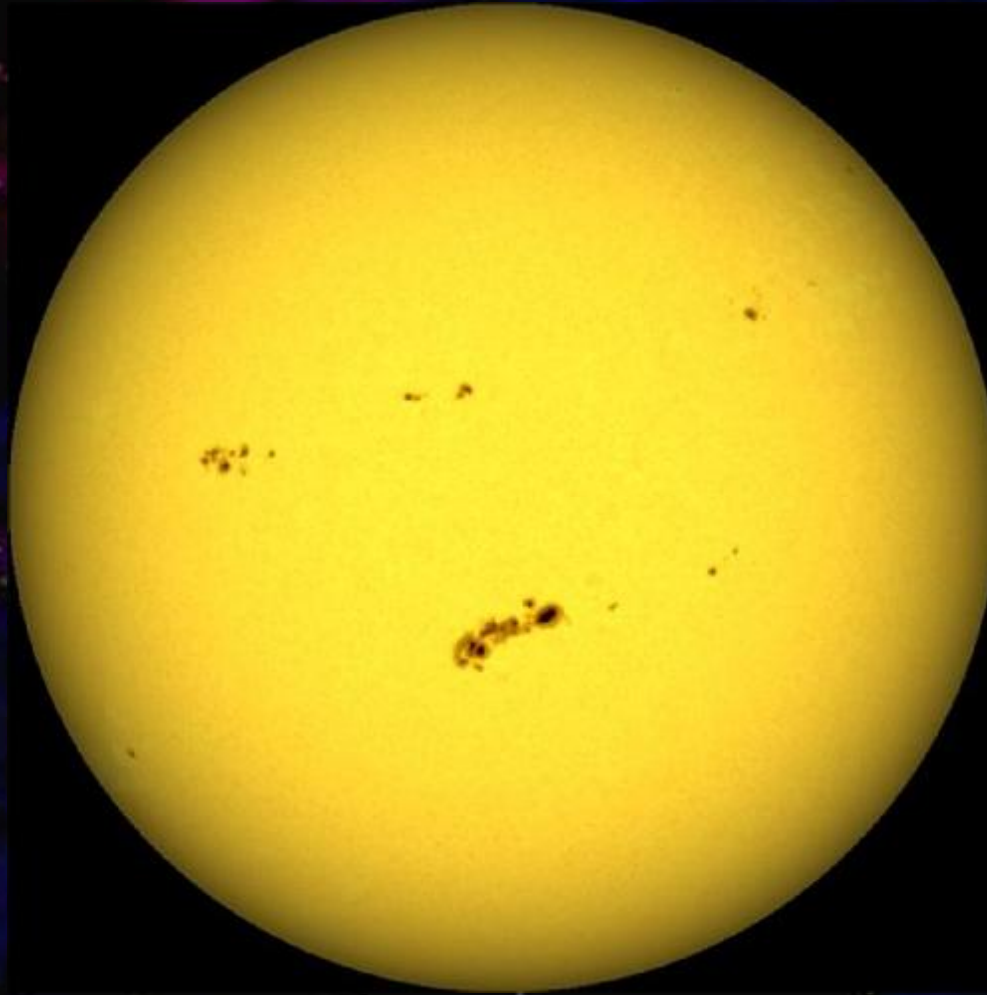
dale.mais@mpiresearch.com

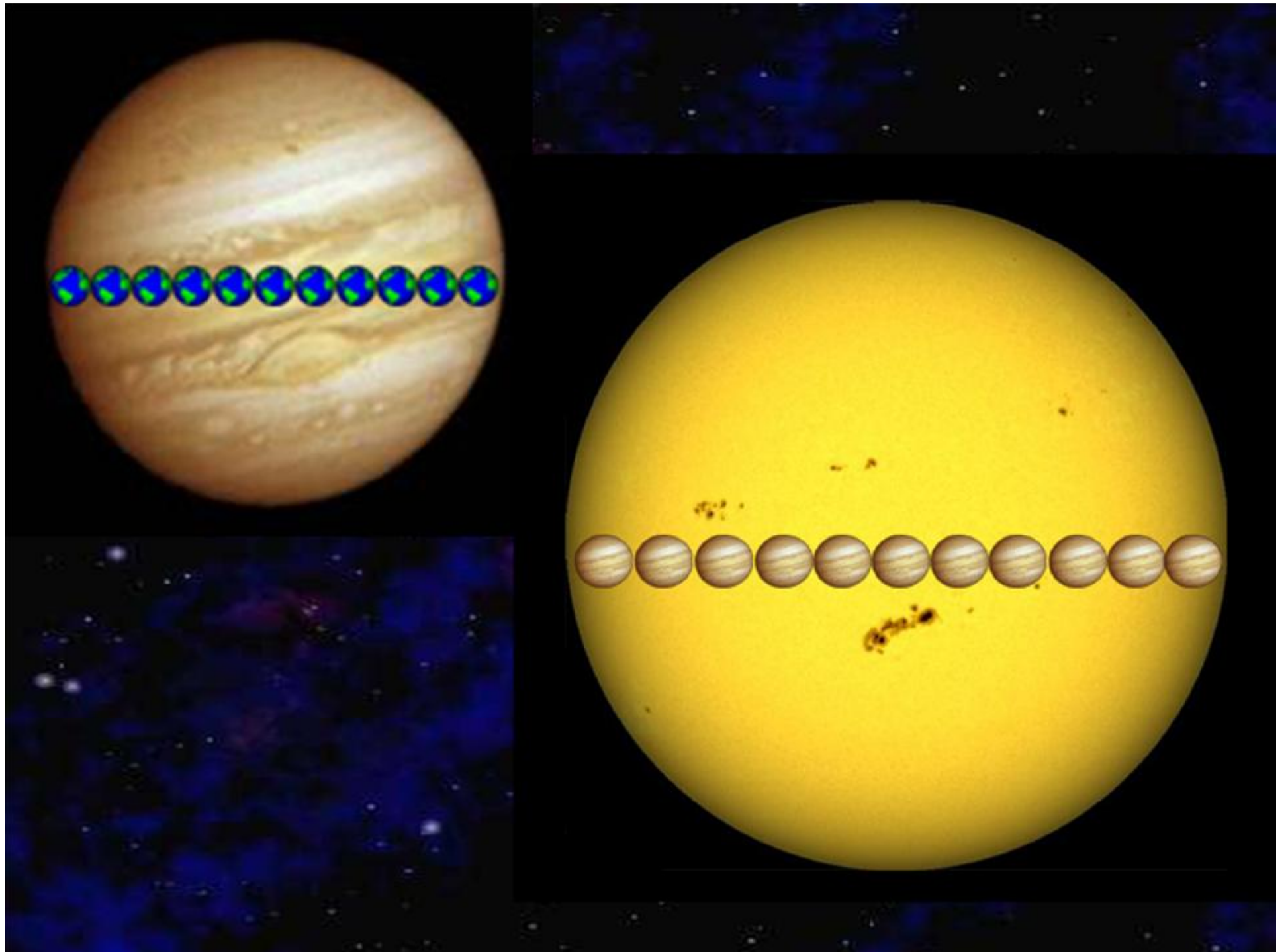
"The COSMOS is all that is...
or ever was... or ever will be"



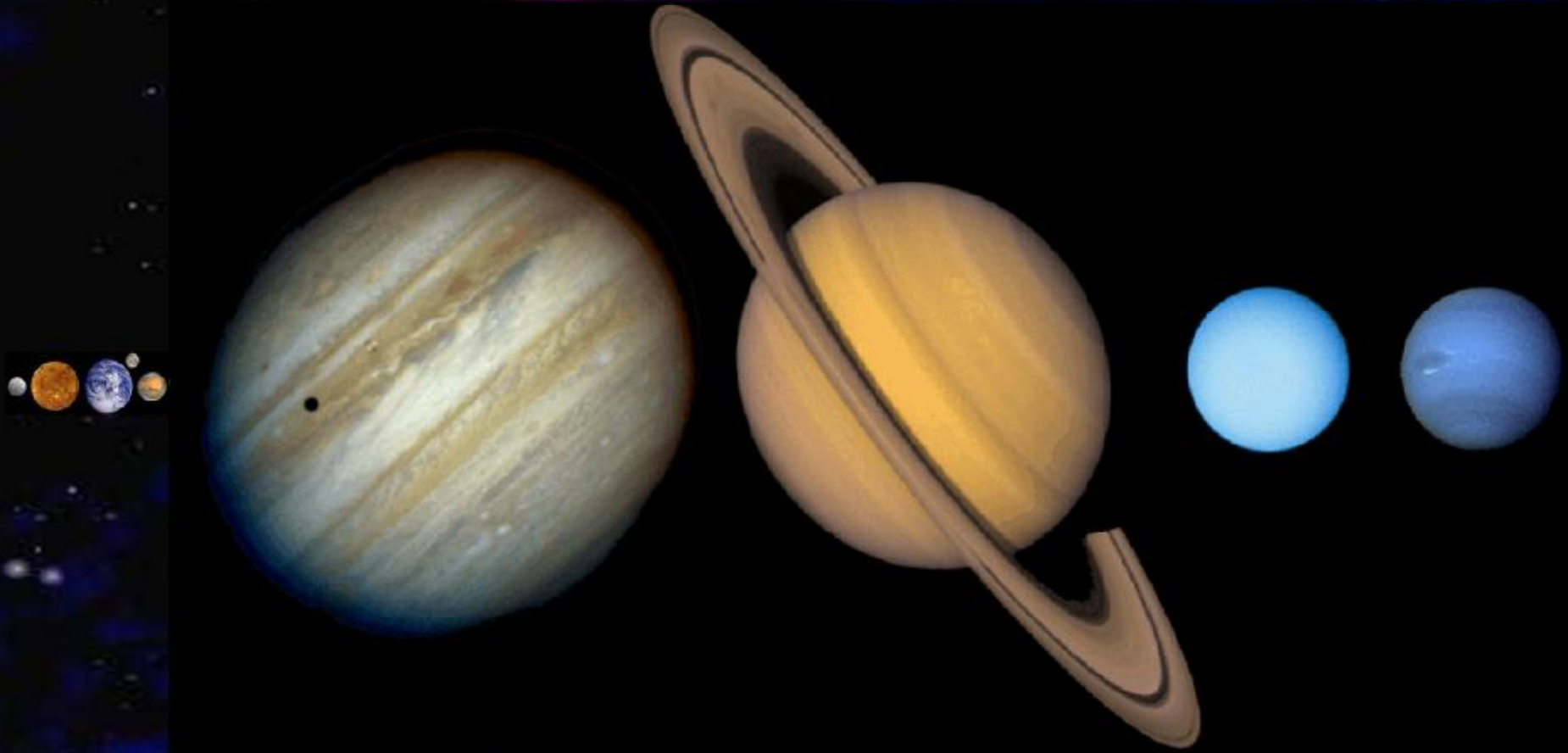
- Carl Sagan
COSMOS

Solar System Inventory: 1 Star





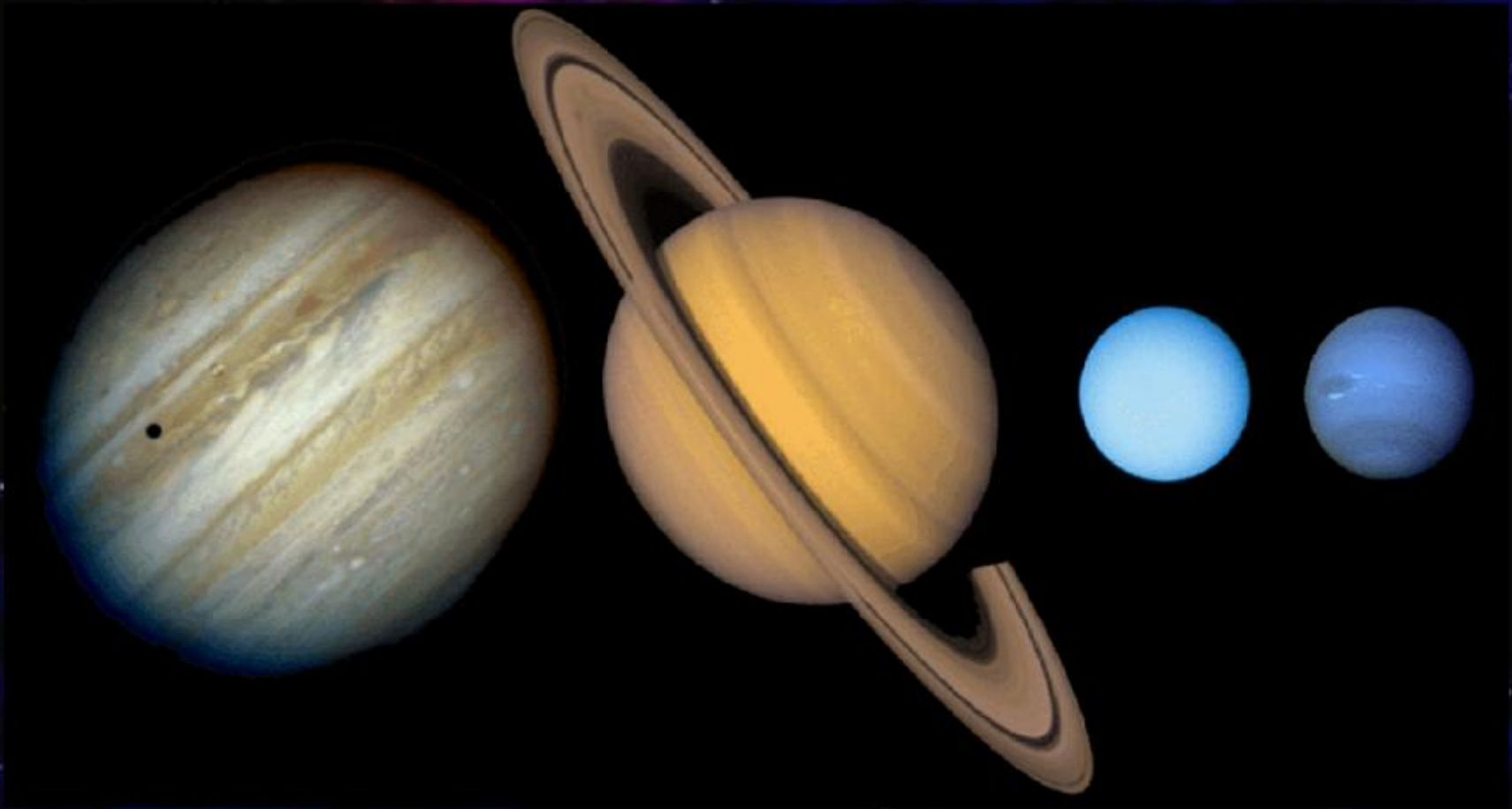
Solar System Inventory: 8 Planets



The Terrestrial Planets



The Jovian Planets

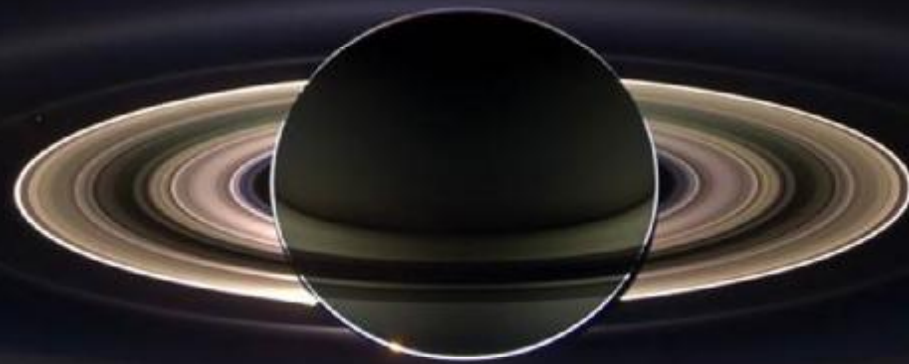


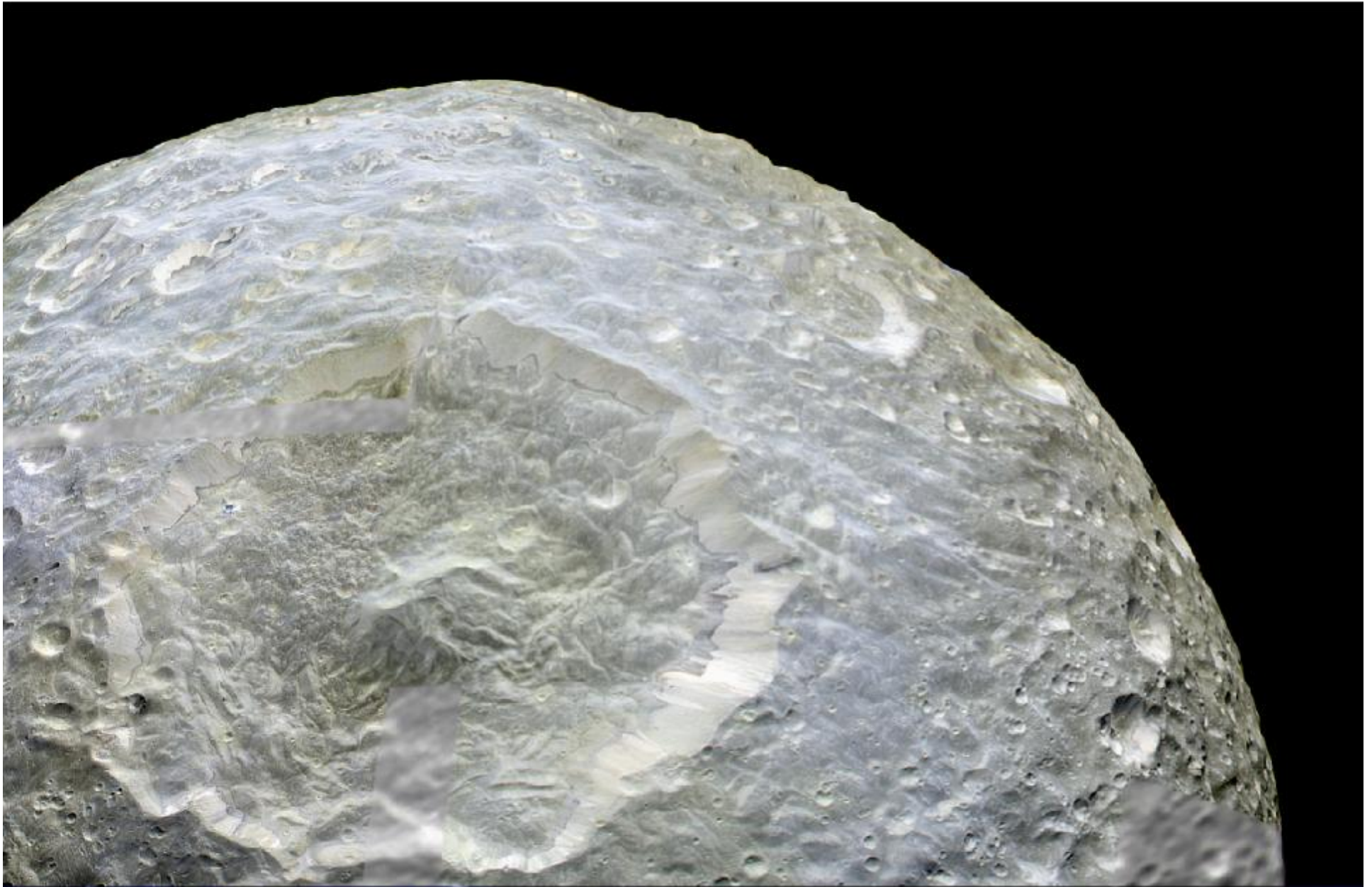
The Jovian Planets



tia
be found at
.net

The Jovian Planets





Mimas, from Cassini fly by, 2010

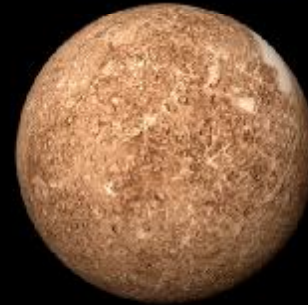
Solar System Inventory: 91 + moons



Ganymede
5262 km



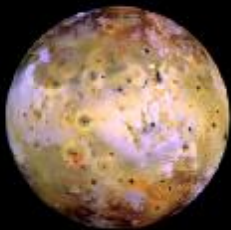
Titan
5150 km



Mercury
4880 km



Callisto
4806 km



Io
3642 km



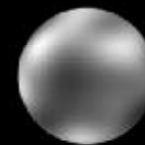
Moon
3476 km



Europa
3138 km



Triton
2706 km



Pluto
2300 km



Titania
1580 km

Solar System Inventory: Countless numbers of Asteroids, Comets, Meteors, Dust

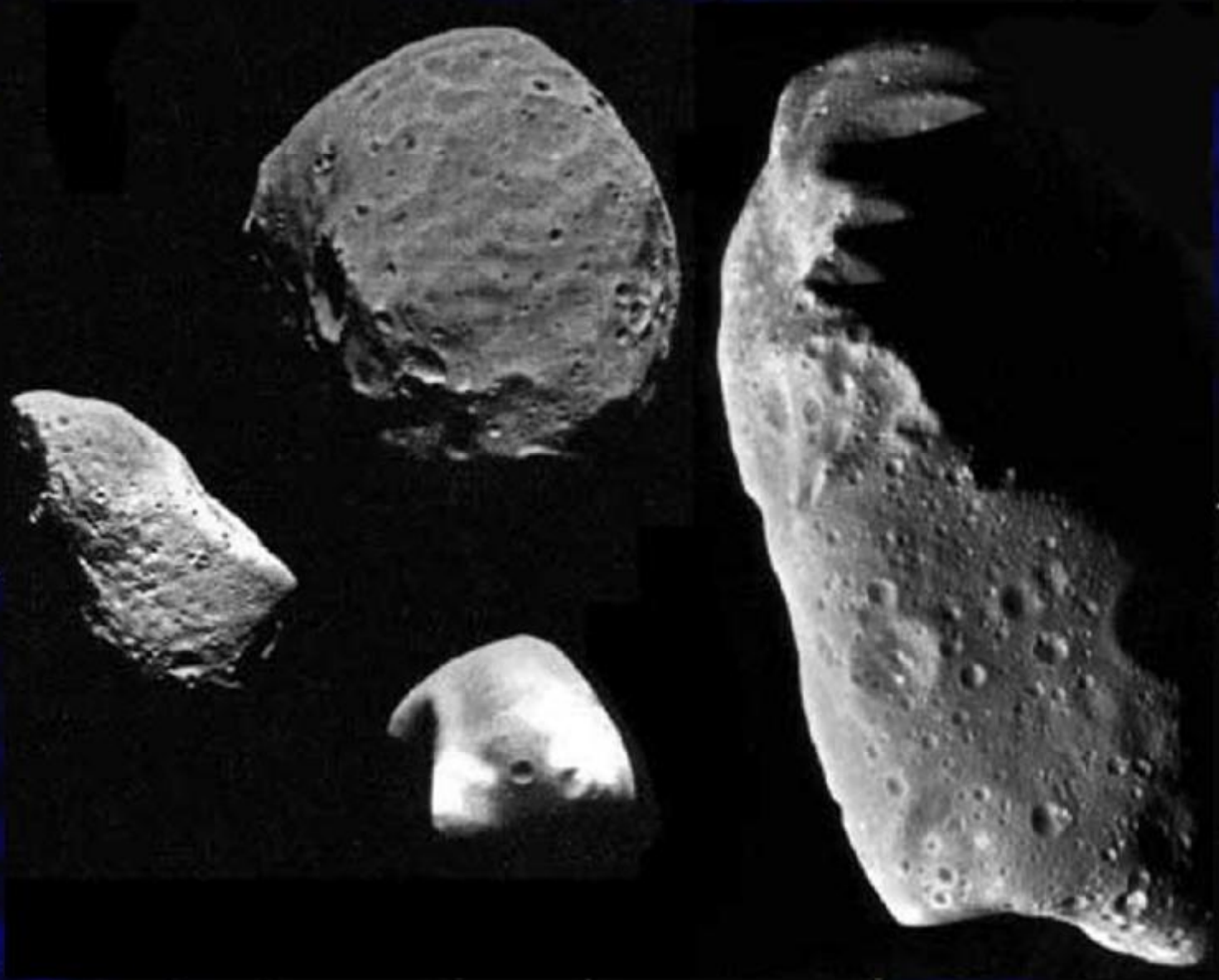


SHENMAN Miller-Ledger



"I DON'T SEE HOW ANY PLANET THAT ALLOWS PARIS HILTON TO RECORD A CD HAS ANY RIGHT PASSING JUDGMENT ON US."

Asteroids (some are dwarf planets now)



Largest known trans-Neptunian objects (TNOs)



Eris



Pluto



Makemake



Haumea



Sedna



Orcus

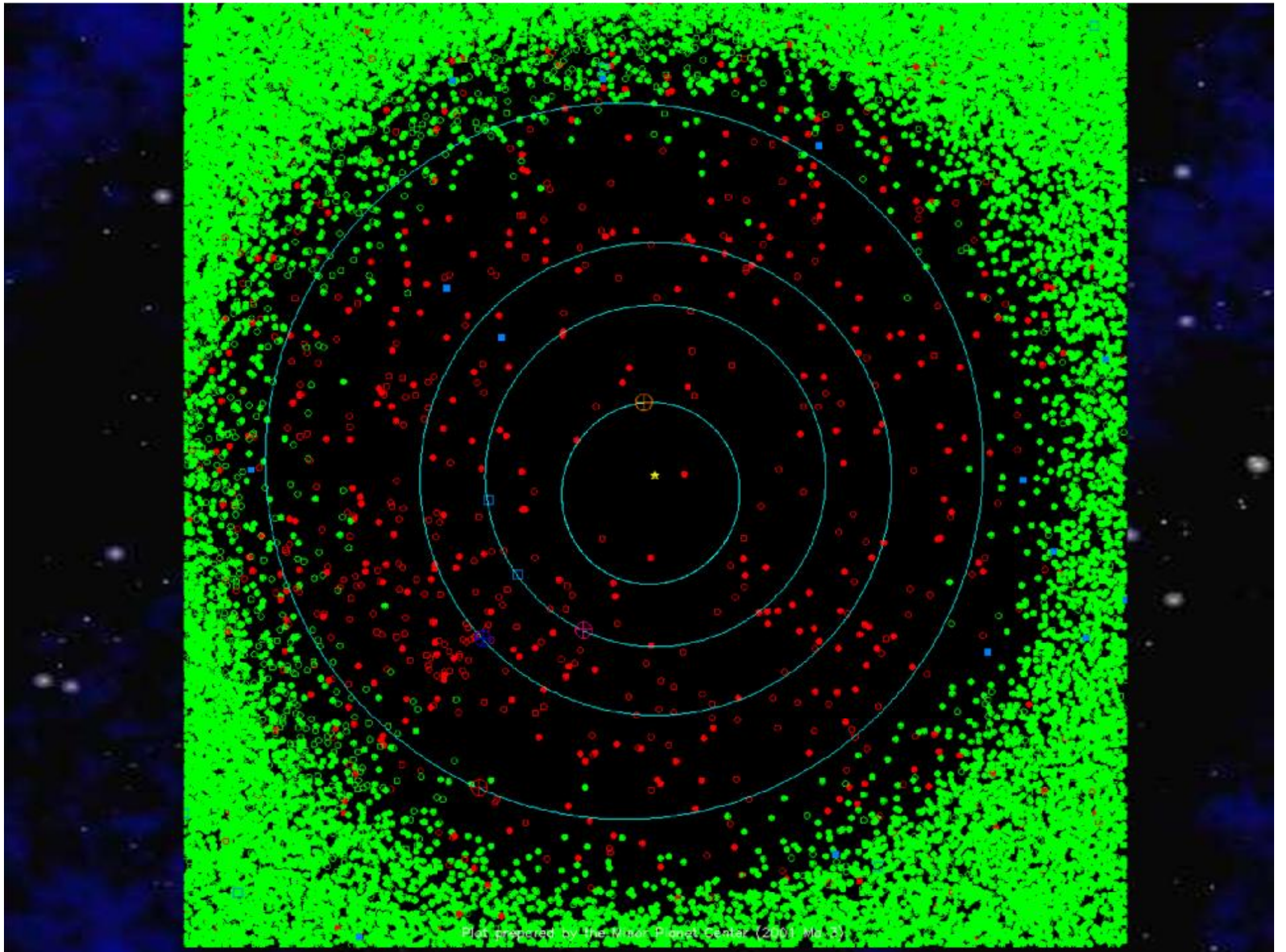


2007 OR₁₀

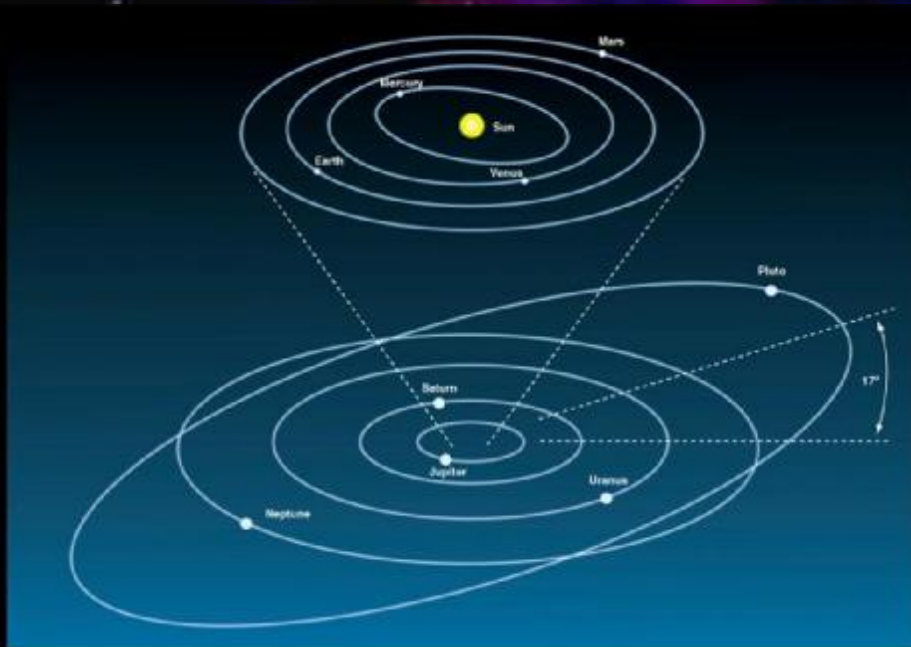


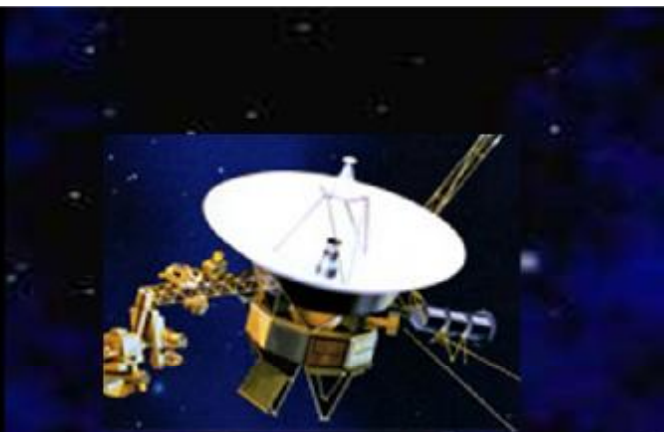
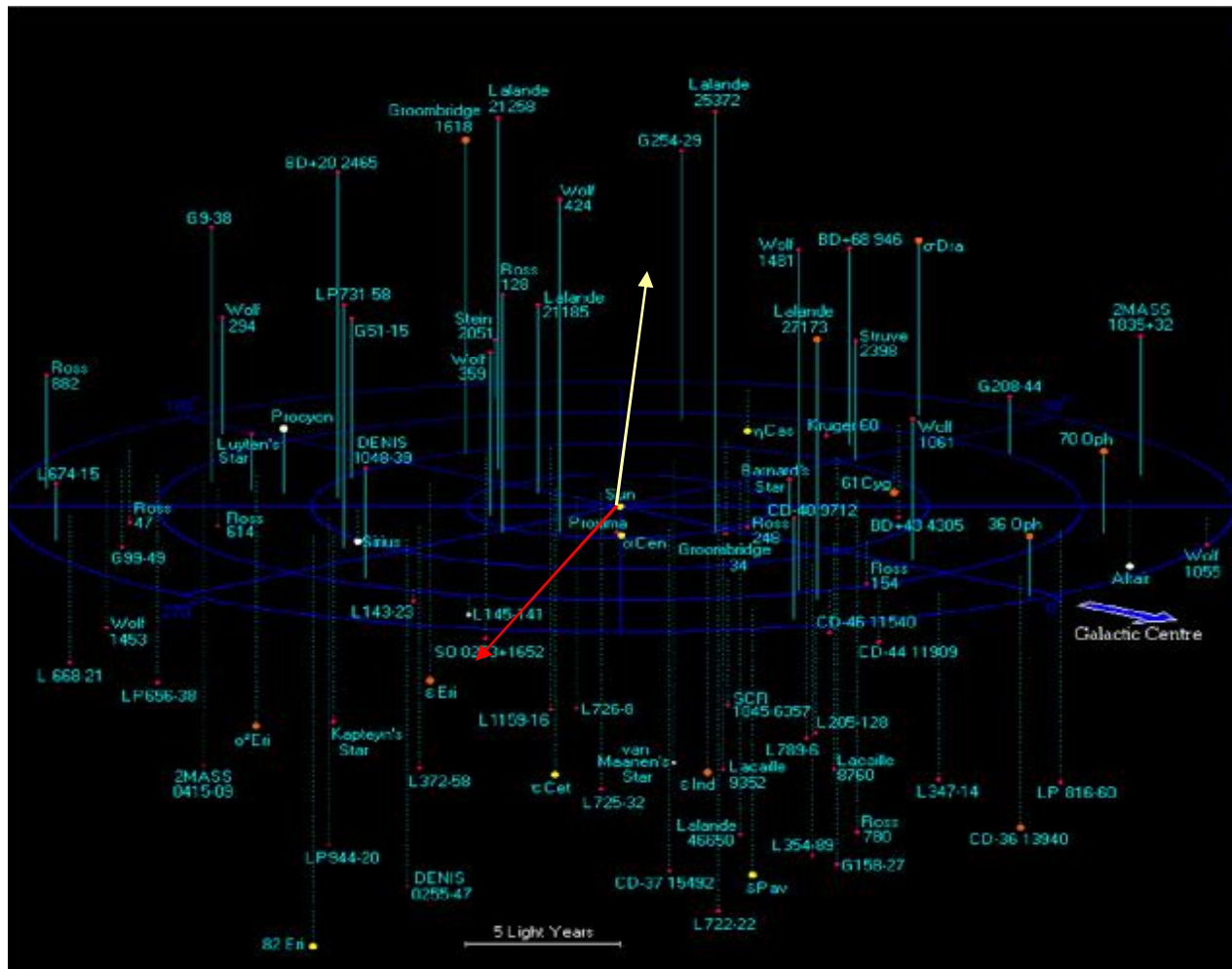
Quaoar



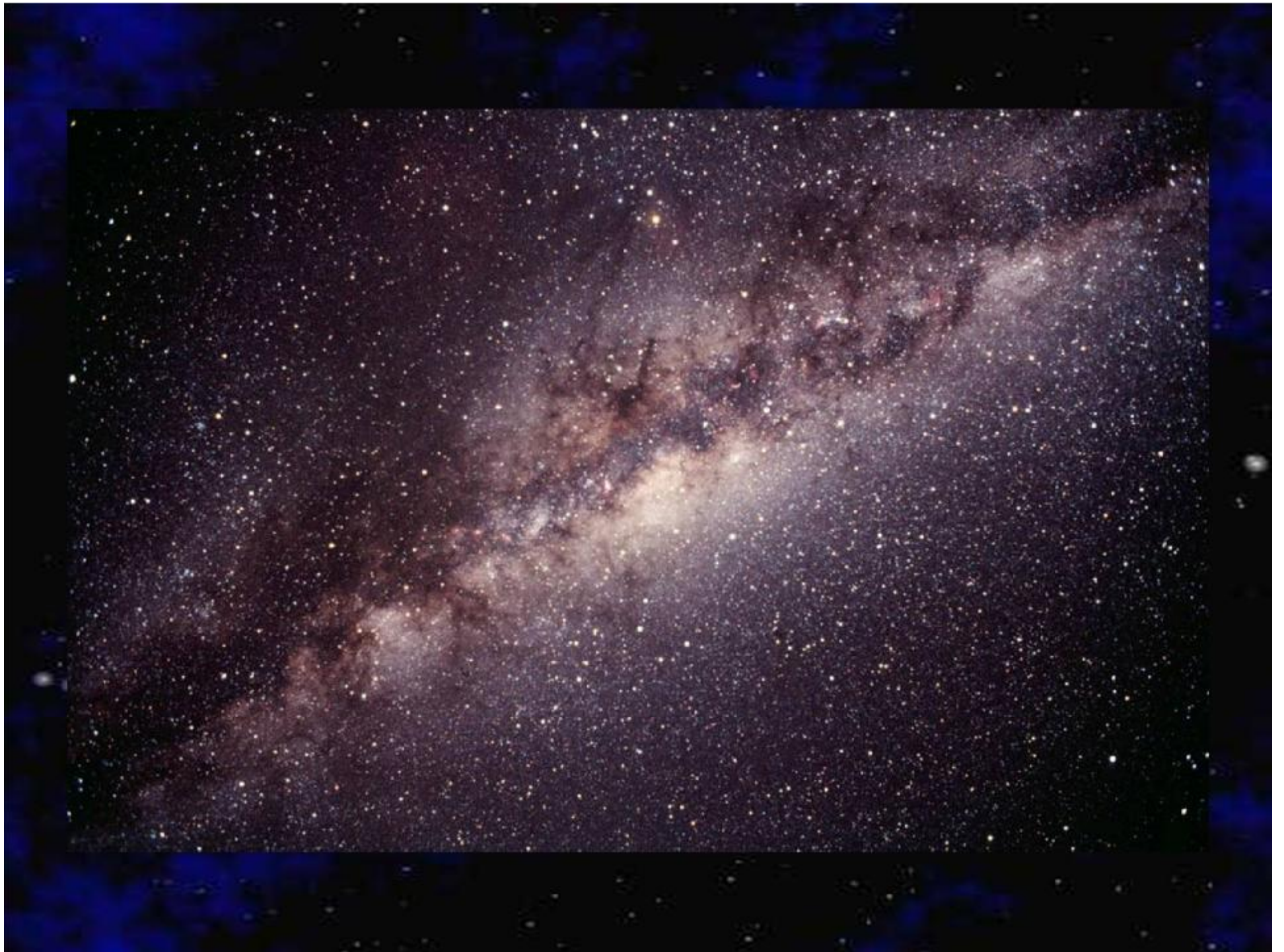


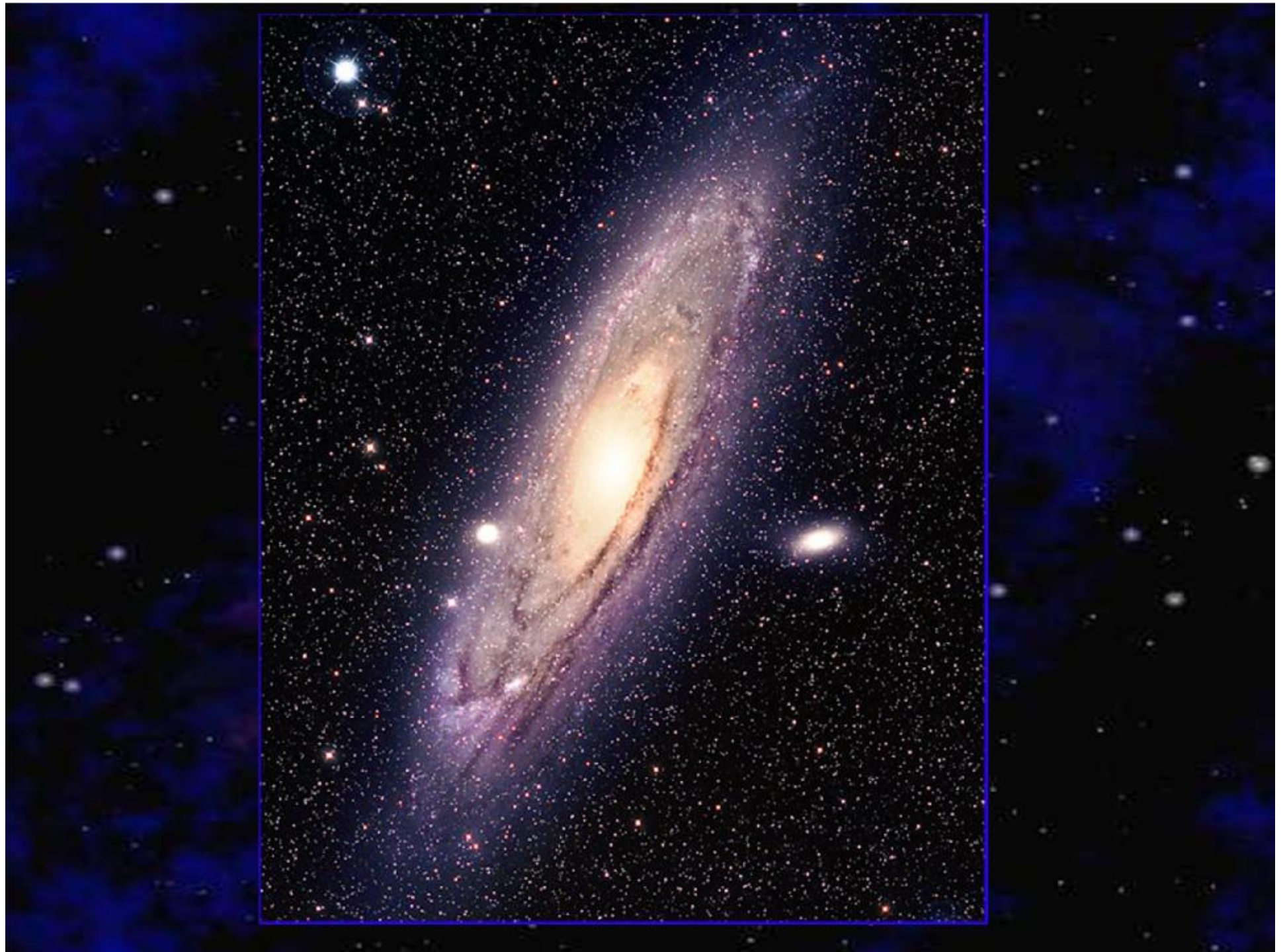
What is our PLACE in the COSMOS?





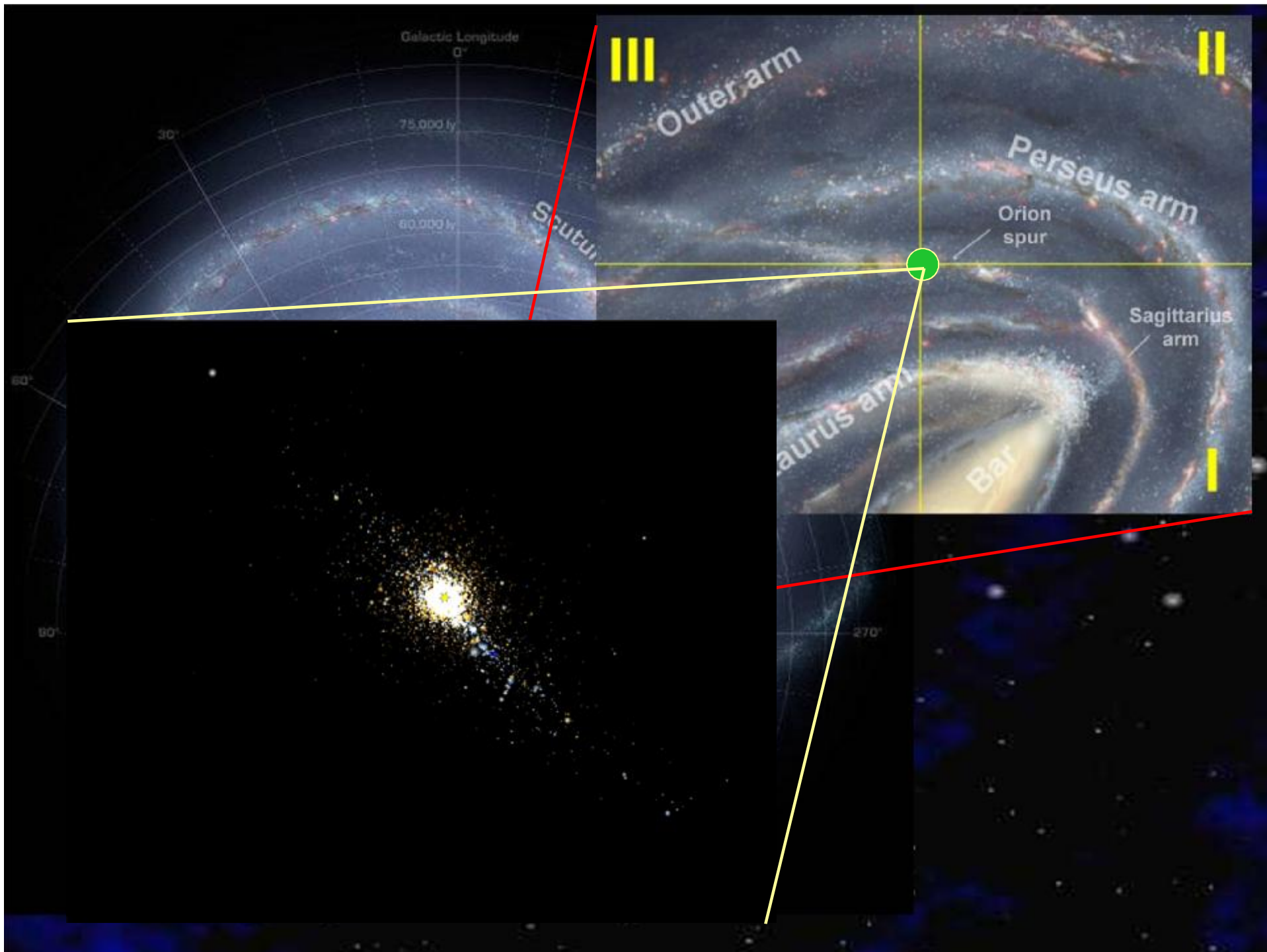
Star	Sun	α^1 Cen	α^2 Cen	Barnard's	BD +36° 2147	Lalande 21185	Sirius	ϵ Eri	CoD -36° 15693	61 Cyg A	Procyon	61 Cyg B	ϵ Ind	τ Cet
s.d.	1.0	1.20	0.90	0.06	0.14	0.15	1.43	0.77	0.20	0.52	1.77	0.36	0.69	0.85
Ly.	0.0	4.39	4.39	5.94	8.20	8.31	8.60	10.49	10.73	11.35	11.40	11.42	11.82	11.89











A New Scale of Things

Sun....the size of a volley ball

Earth....half the size of a BB, 100 yards away

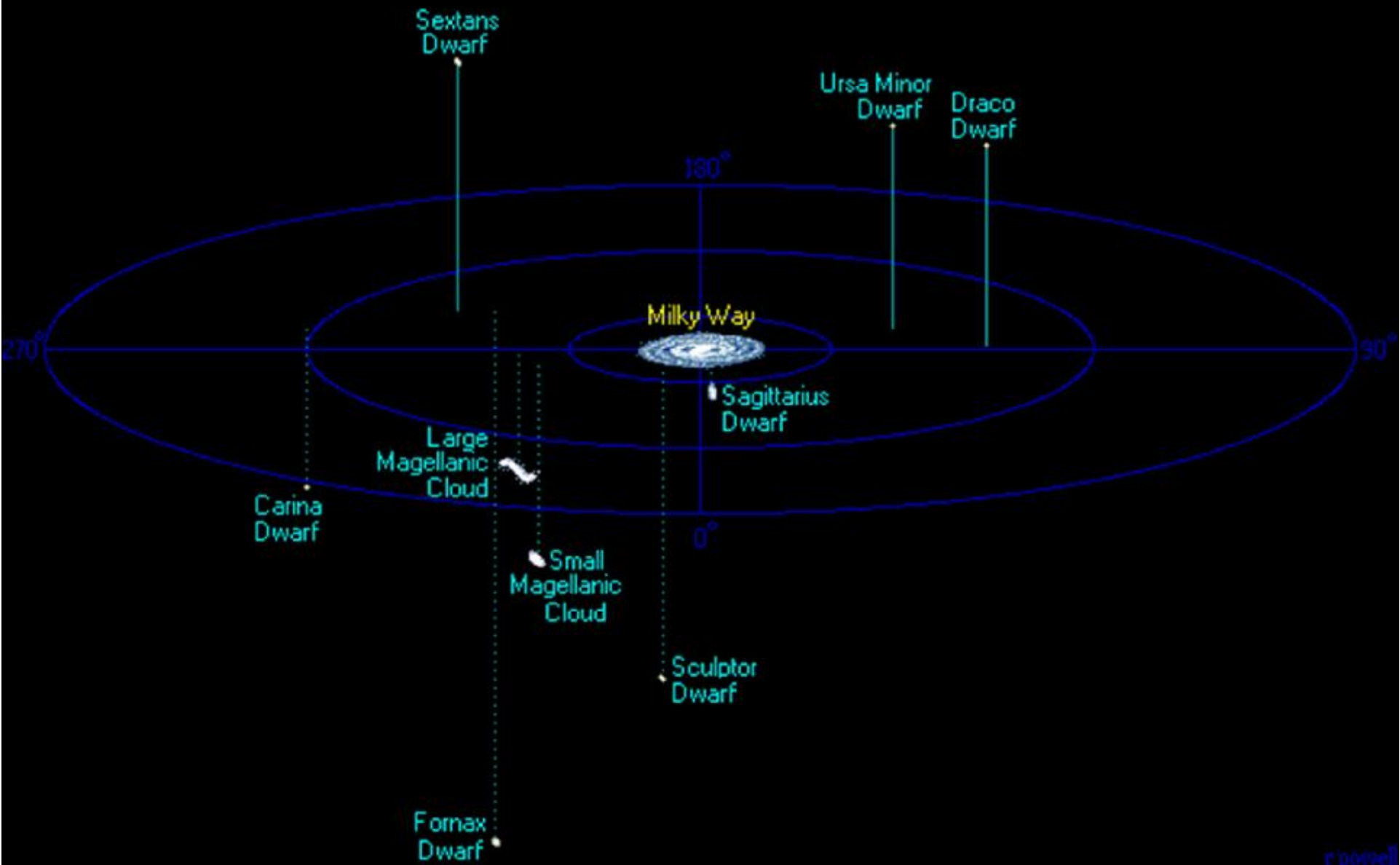
Solar System....ends at Notre Dame campus

Pioneer Space Craft....approaching Niles

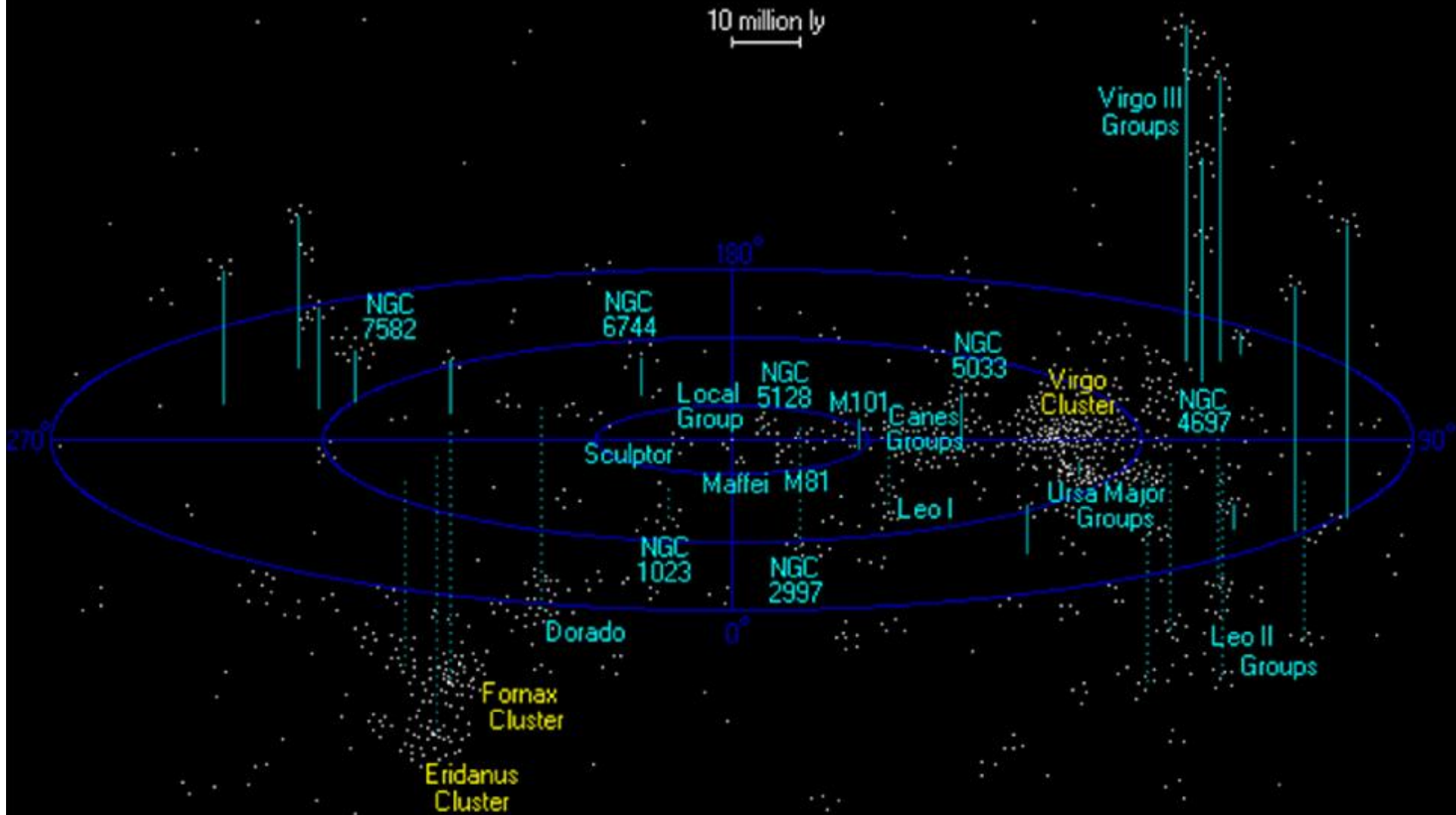
Nearest star....State of New Mexico

Center of Milky Way Galaxy....40x farther then Earth-Moon

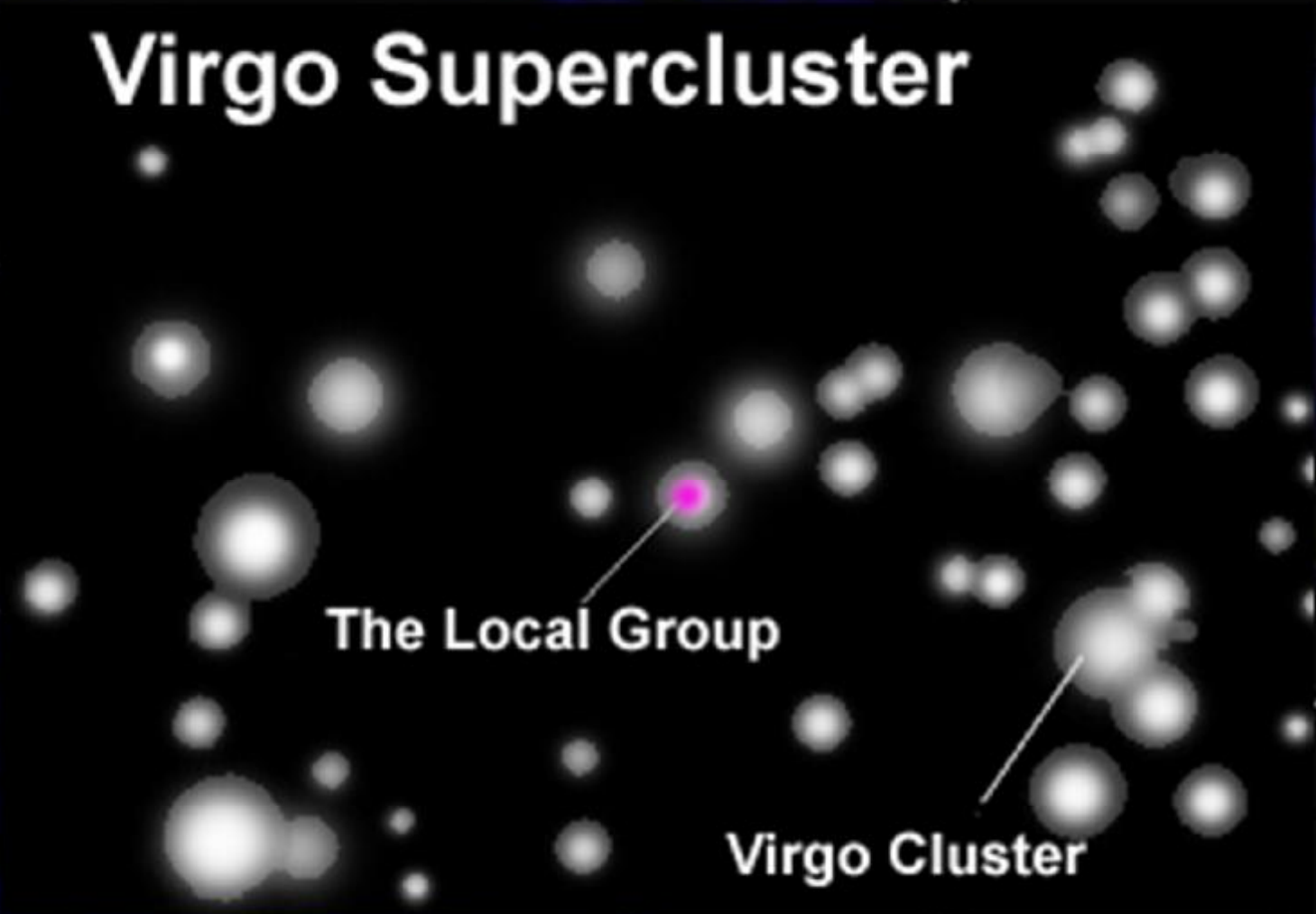
100 000 ly



10 million ly

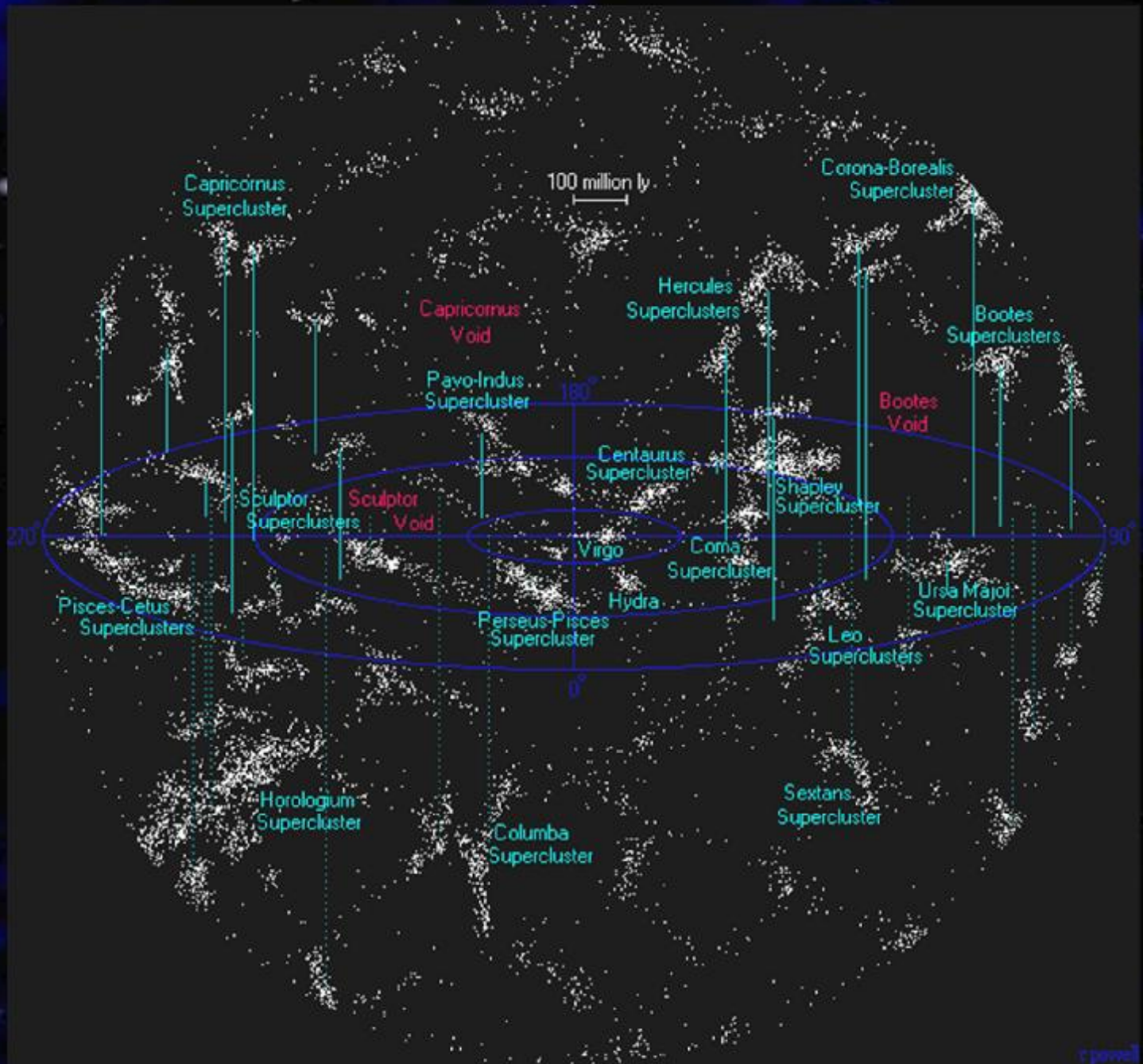


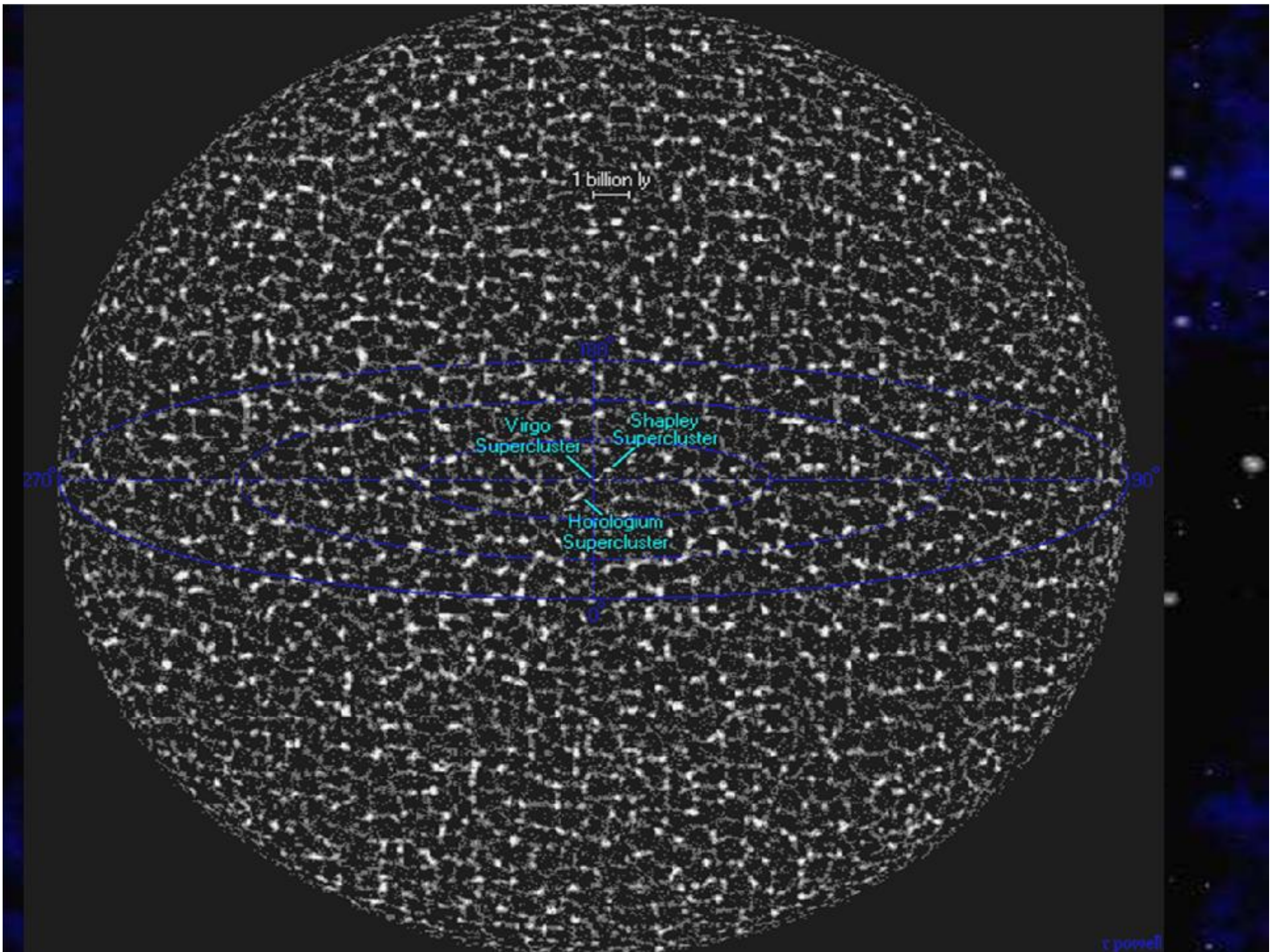
Virgo Supercluster

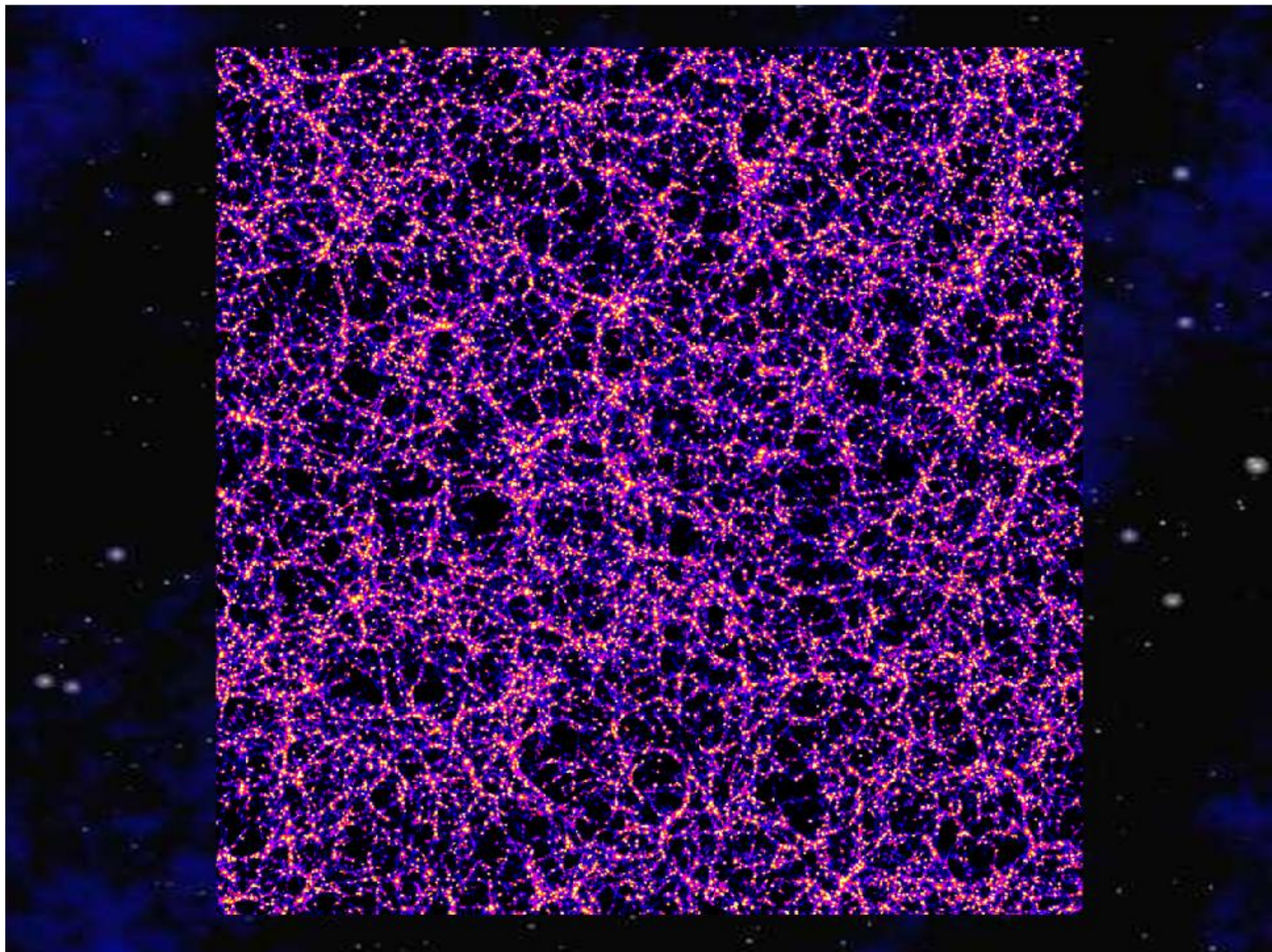
A diagram of the Virgo Supercluster. The background is a dark blue field of stars. A central black rectangle contains a field of white circles representing galaxy clusters. The circles vary in size and brightness. A pink dot is located in the center-left area, with a white line pointing to it from the text 'The Local Group'. To the right and slightly below the pink dot is a larger, more complex cluster of white circles, with a white line pointing to it from the text 'Virgo Cluster'. The title 'Virgo Supercluster' is written in white at the top left of the black rectangle.

The Local Group

Virgo Cluster



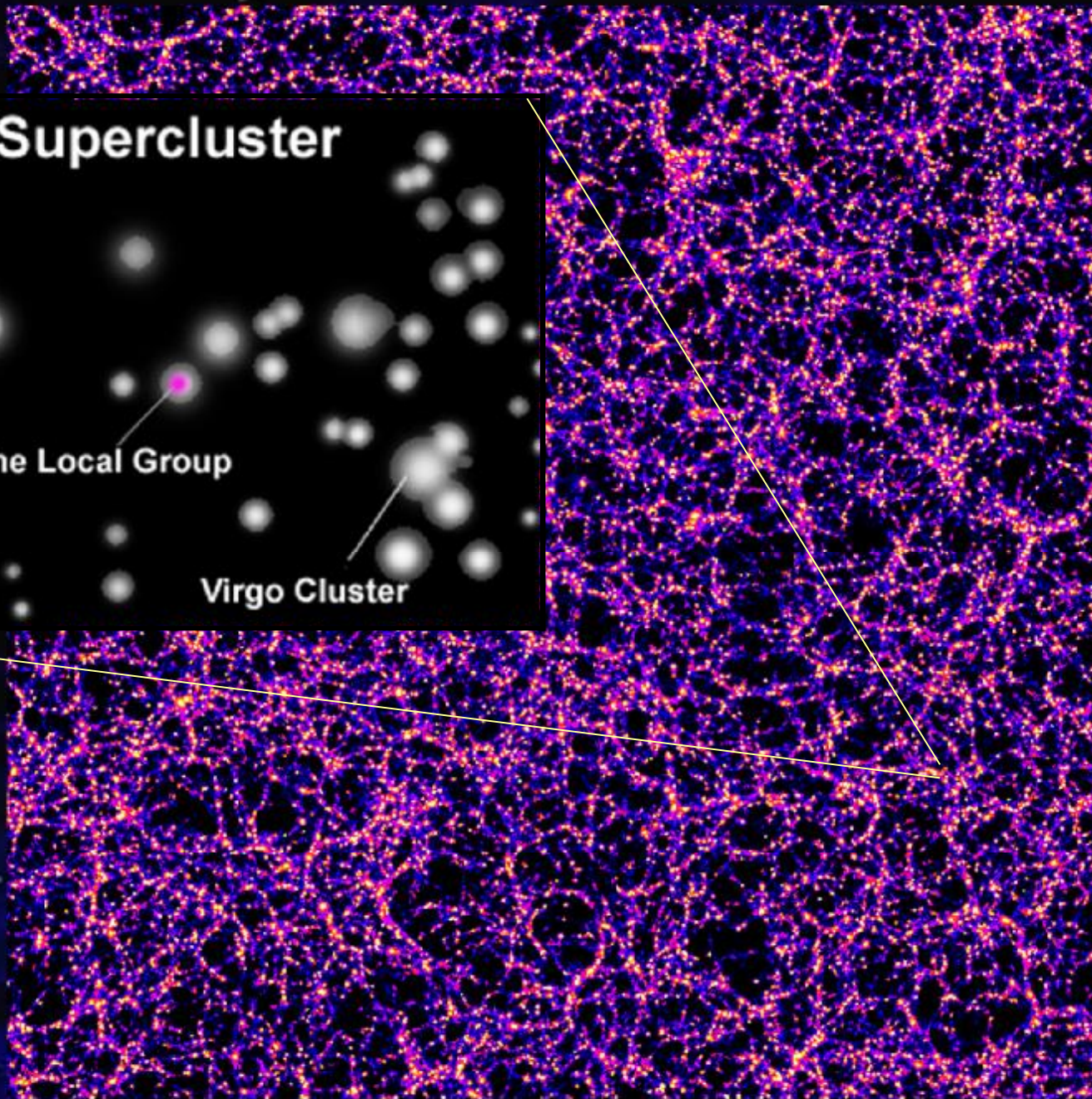




Virgo Supercluster

The Local Group

Virgo Cluster



Deepest Image EVER





Organization of the Sky

Three Main Ideas:

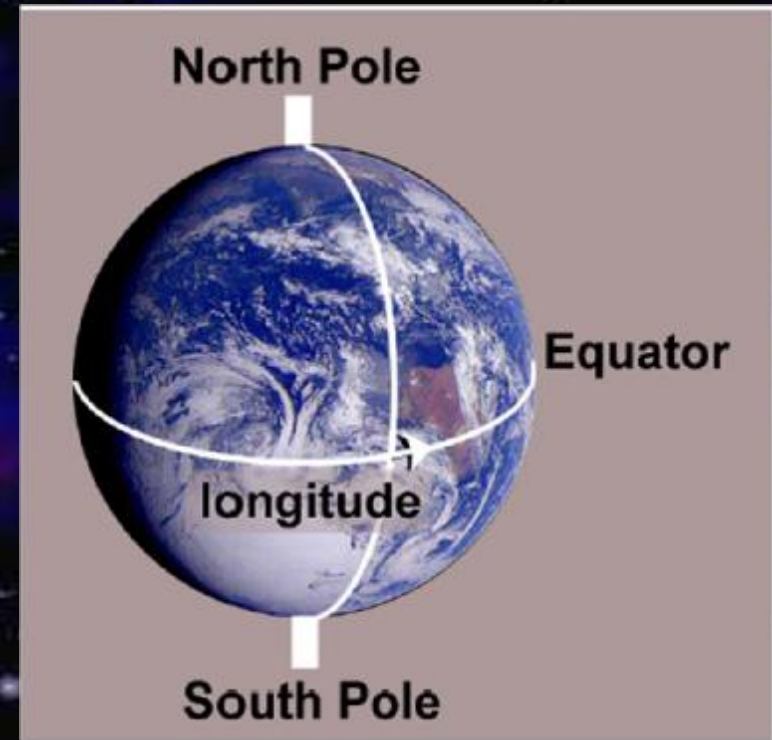
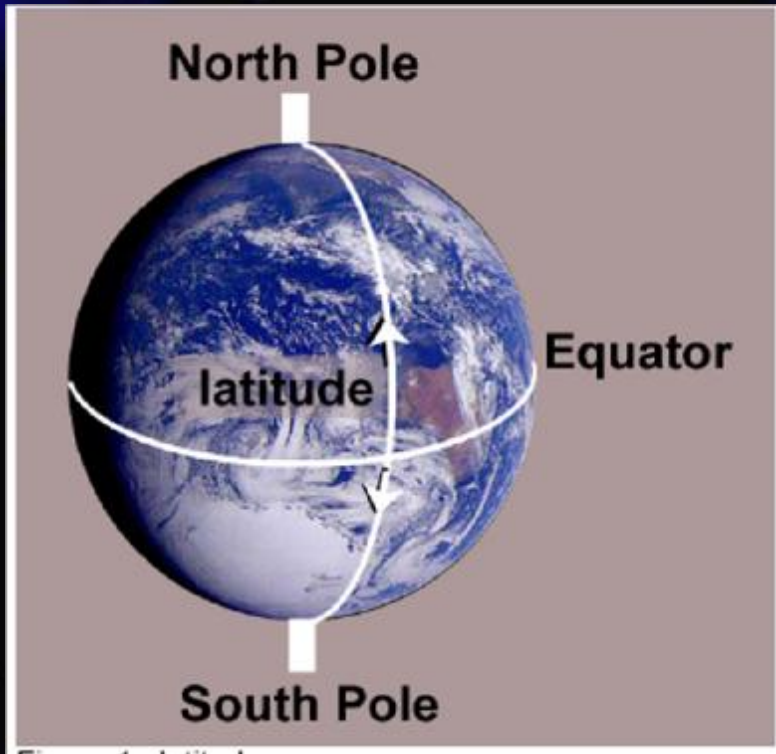
§ What do we see in the sky?

§ Is there ORDER or CHAOS?

§ How are the motions related?

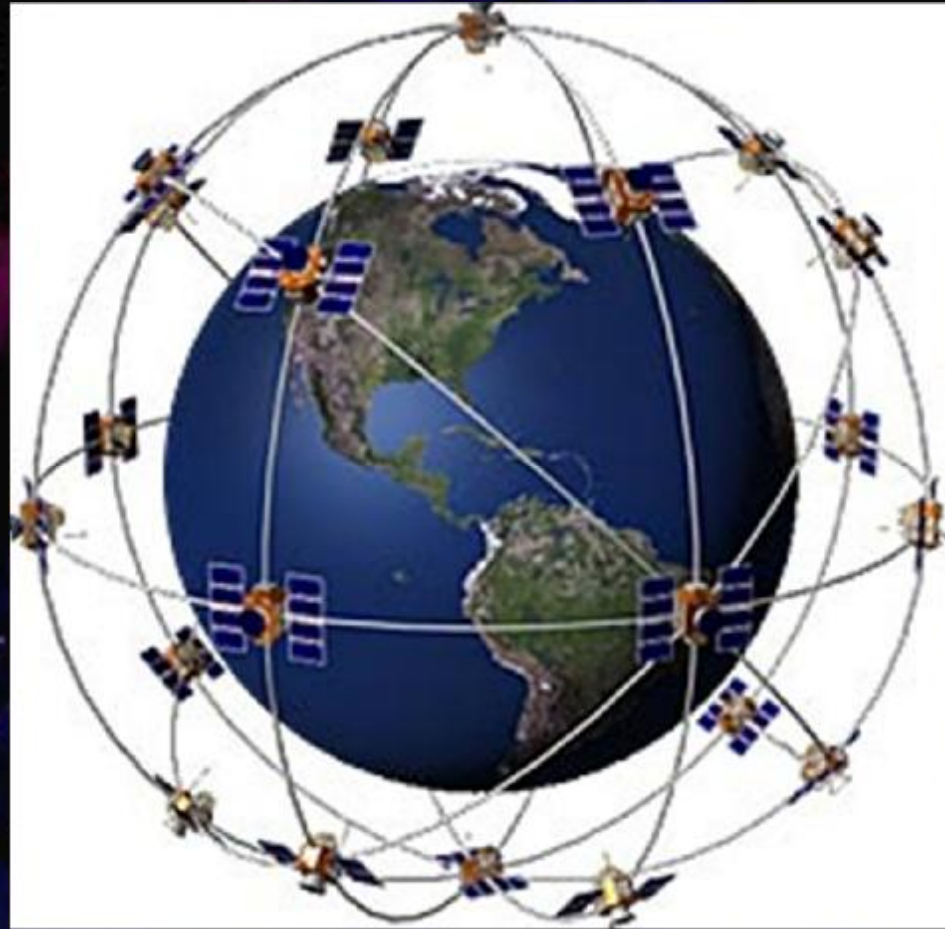
1. The Sky at Night
2. Seasons, Lunar phases, Eclipses
3. Geocentric → Heliocentric Cosmology

Navigating the Night Sky -The Earth

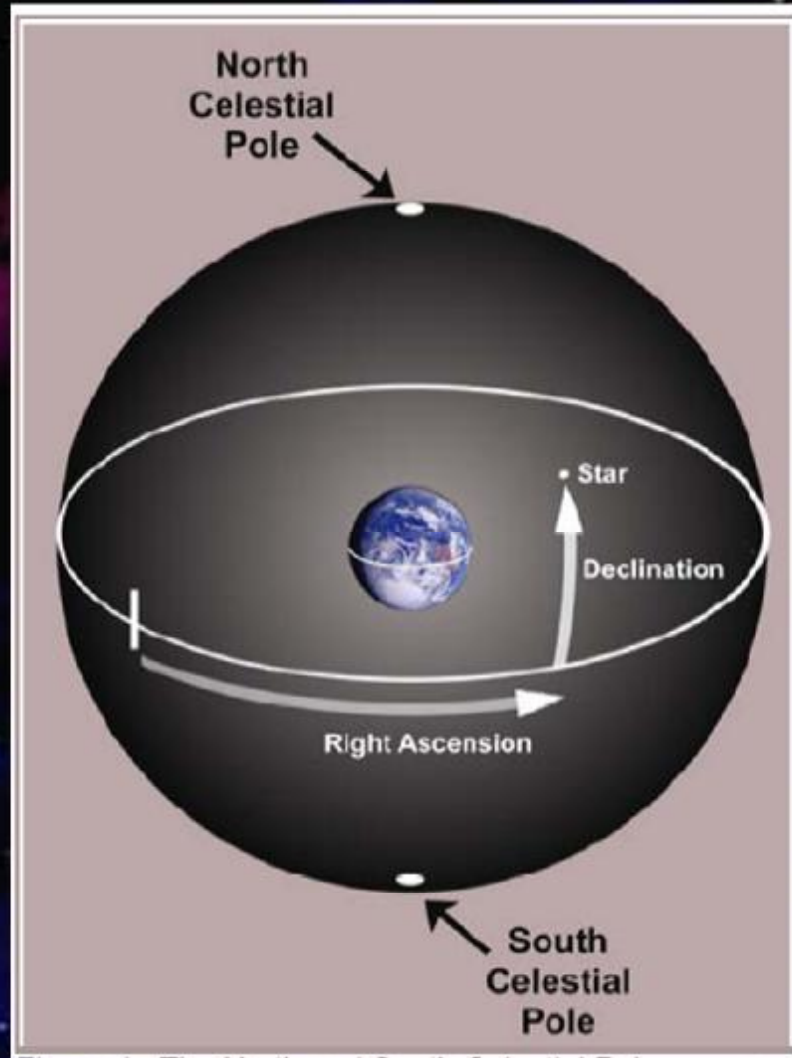


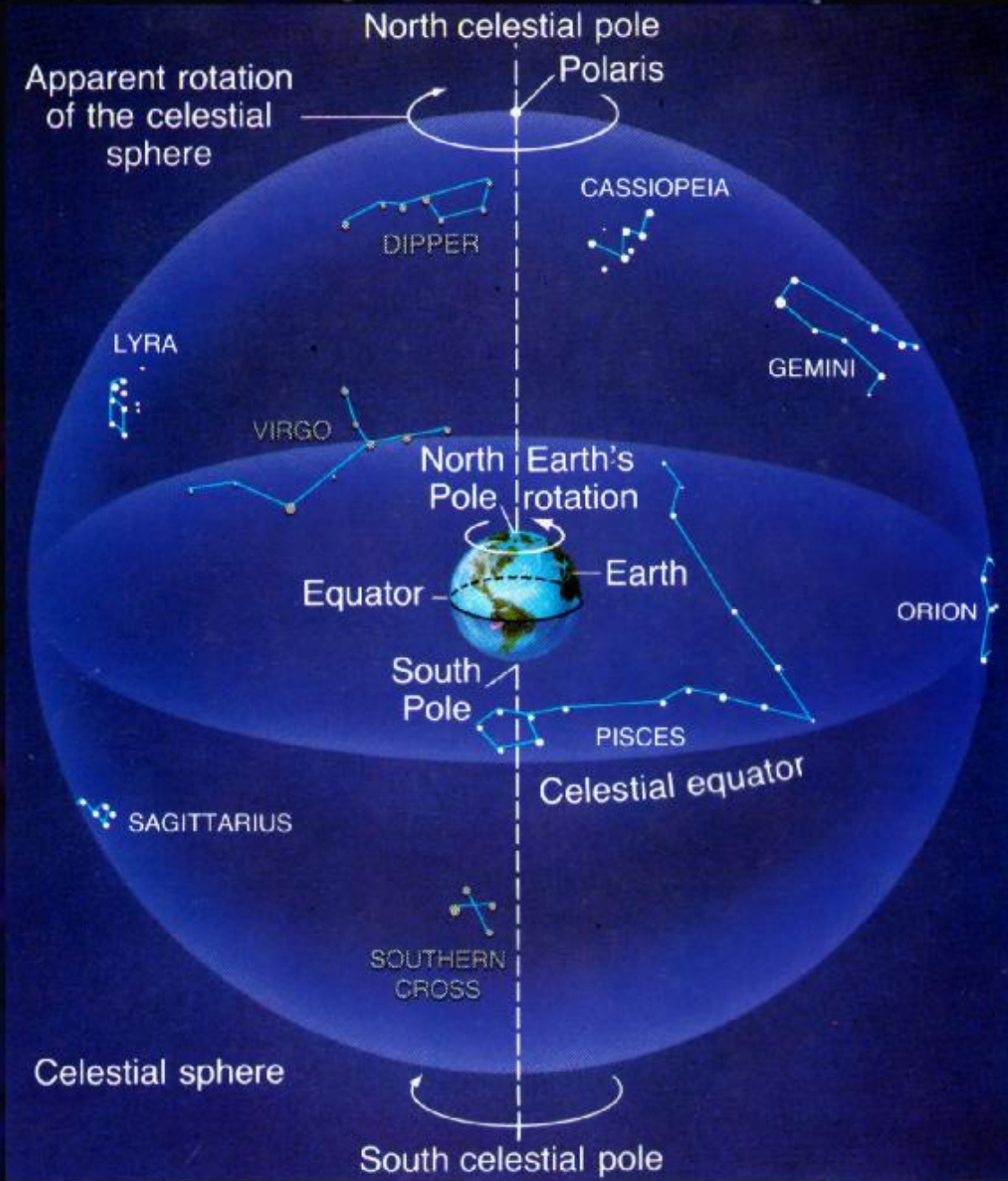
Latitude and Longitude

Navigating the Night Sky - The Earth



Navigating the Night Sky - The Sky





Star Trails:



Star Trails:



Star Trails:

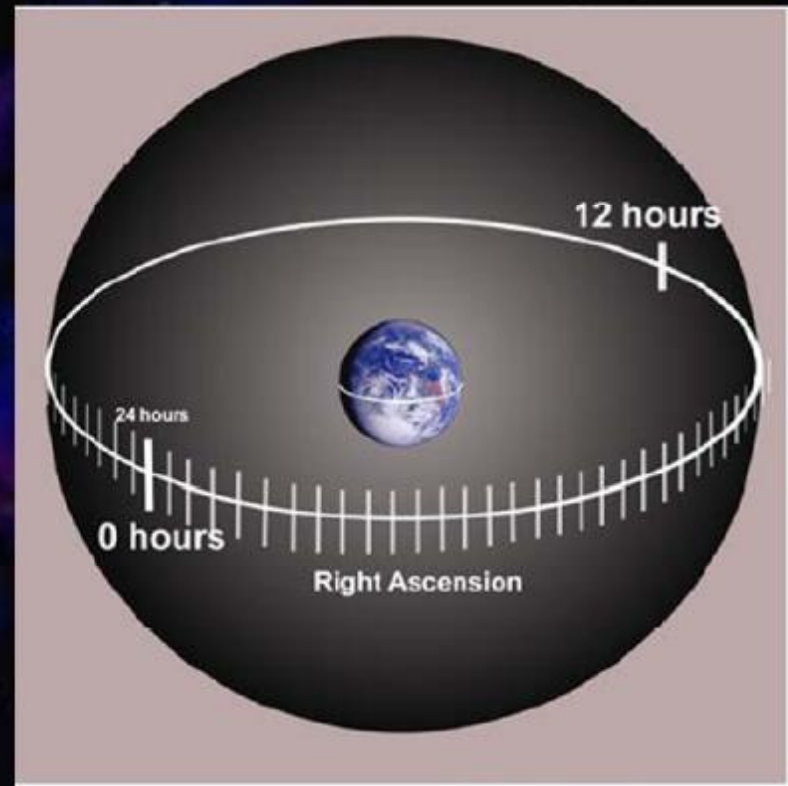
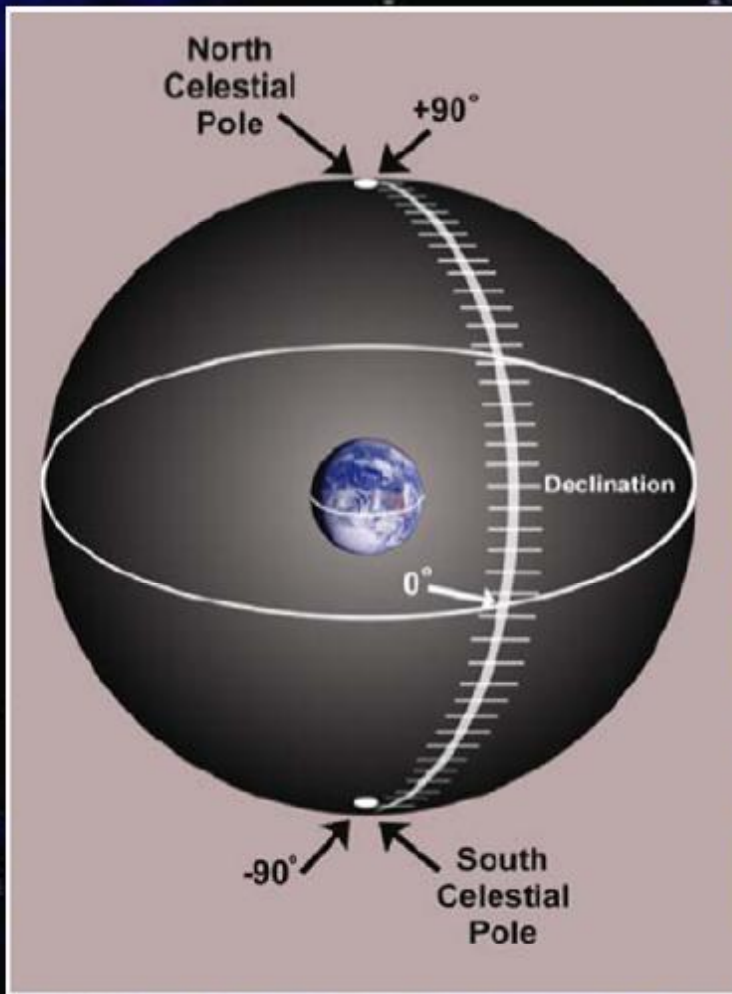


Star Trails:





Navigating the Night Sky - The Sky



Declination and Right Ascension

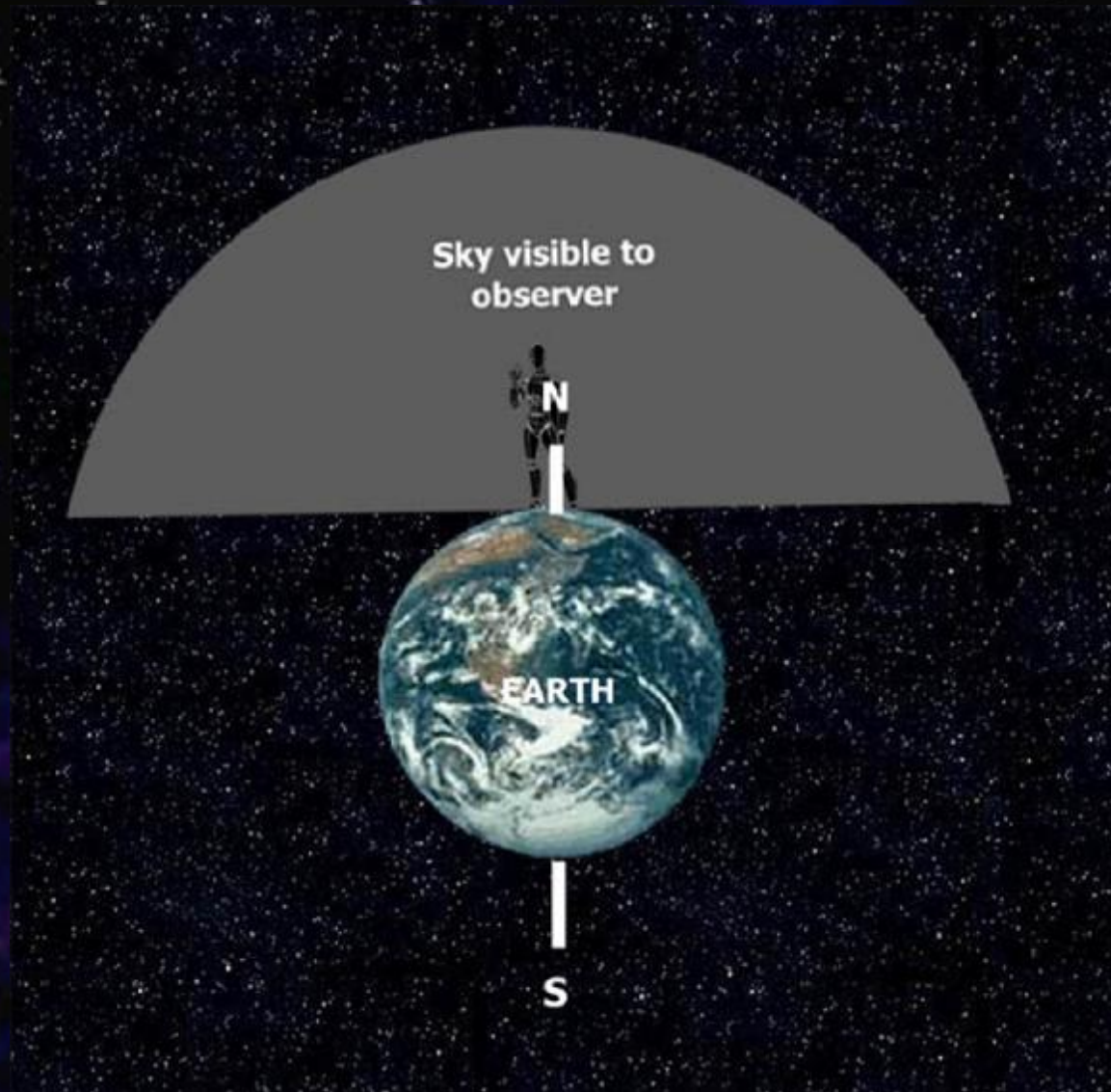
Latitude and Longitude = Declination and Right Ascension



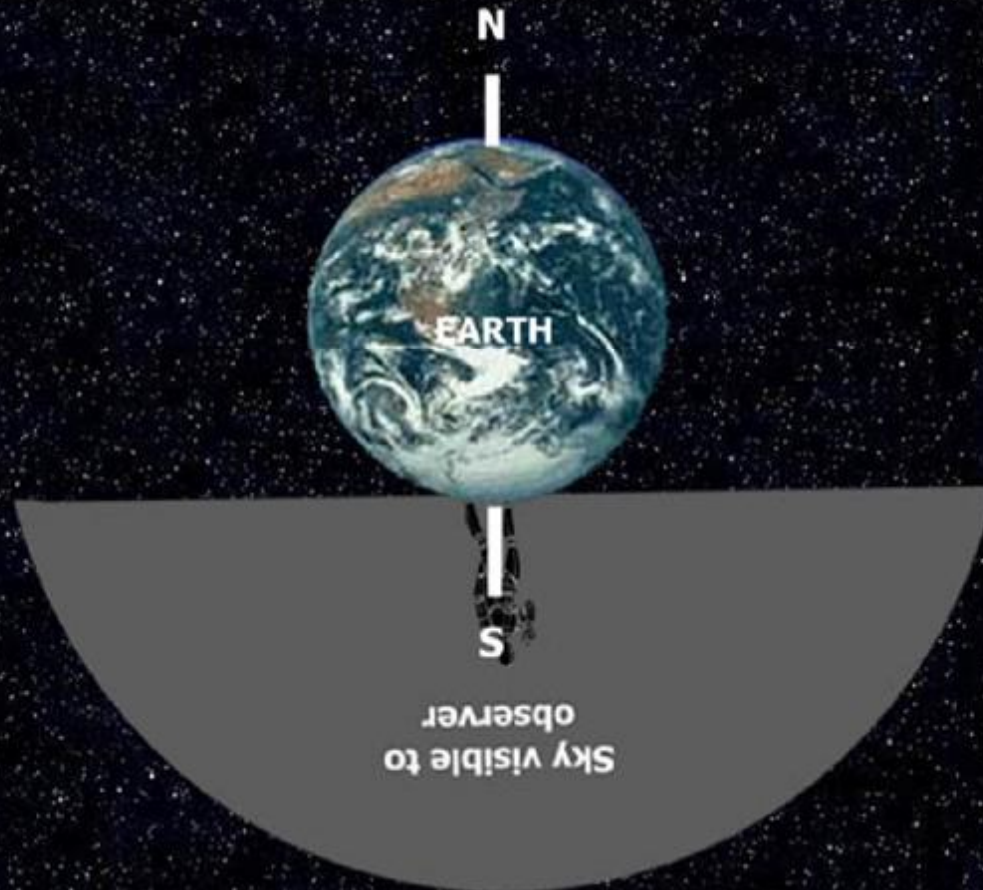
CELESTIAL SPHERE

It is often convenient to imagine the Earth is at the center of a great "sphere of the sky"

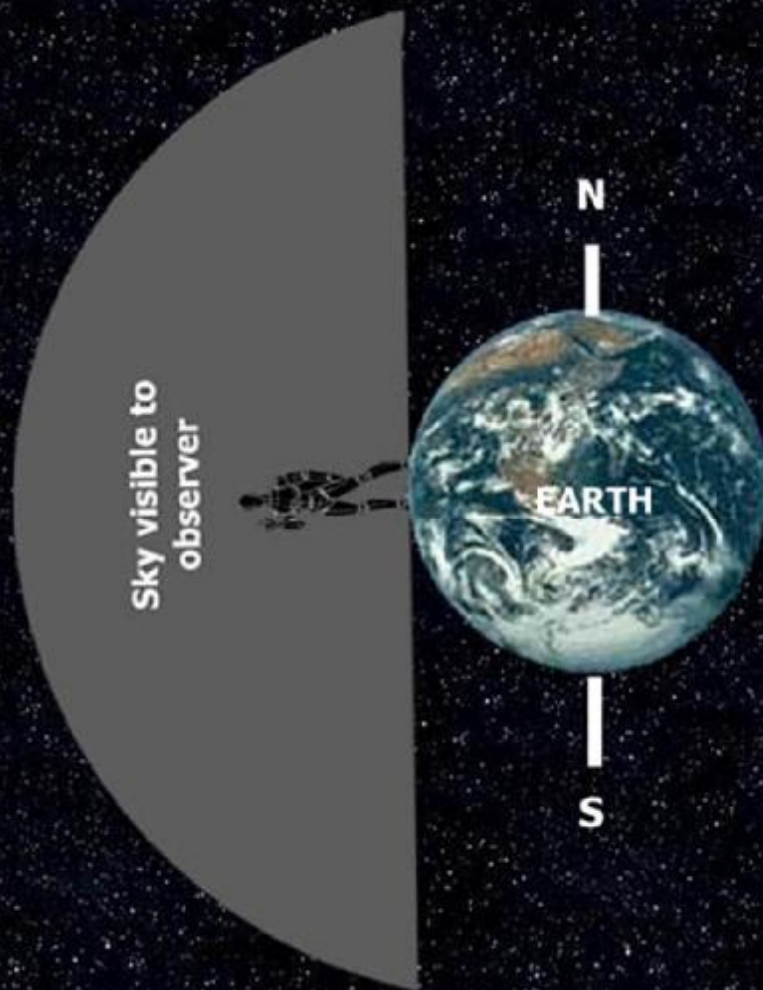
Observer at North Pole



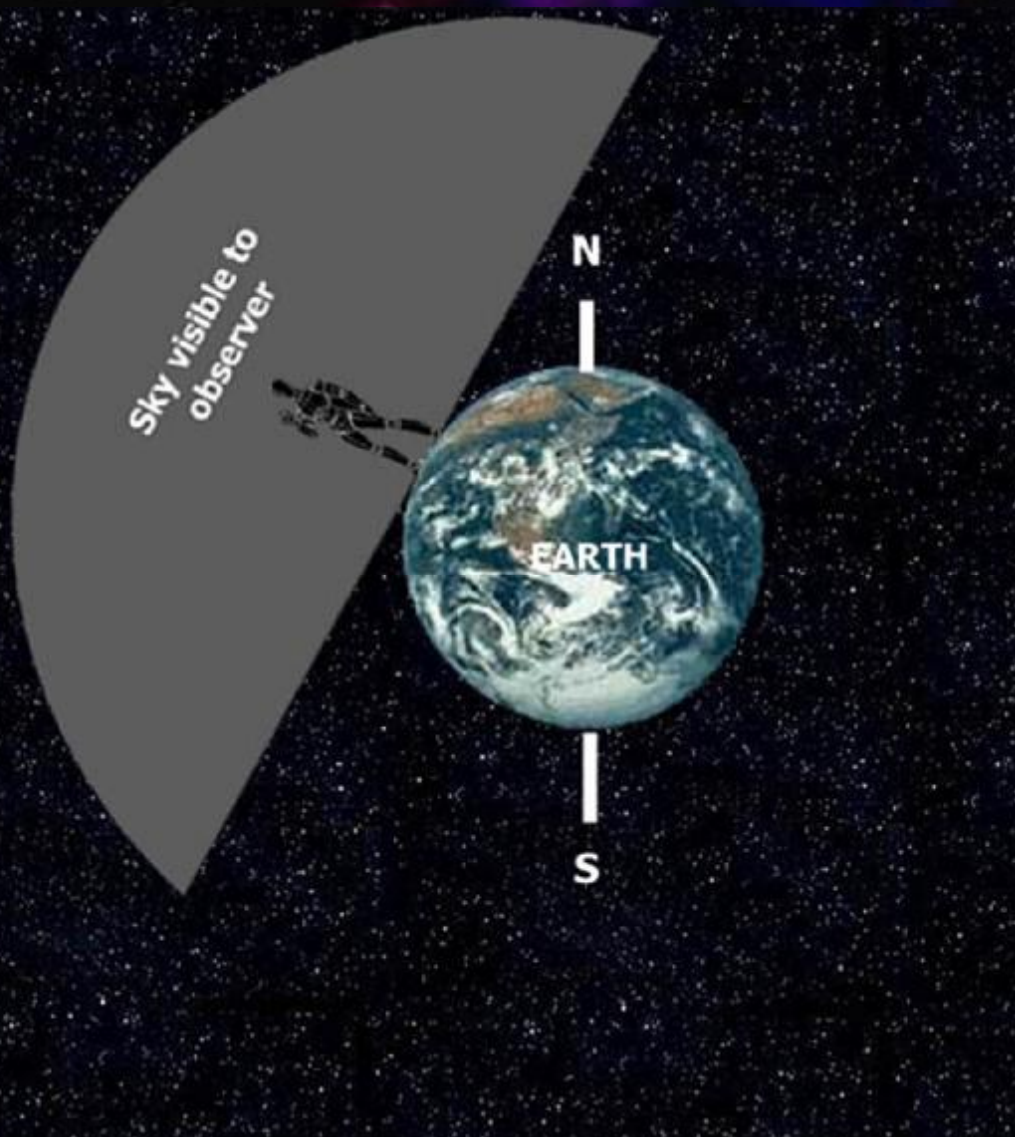
Observer at South Pole



Observer at Equator



Observer at San Diego

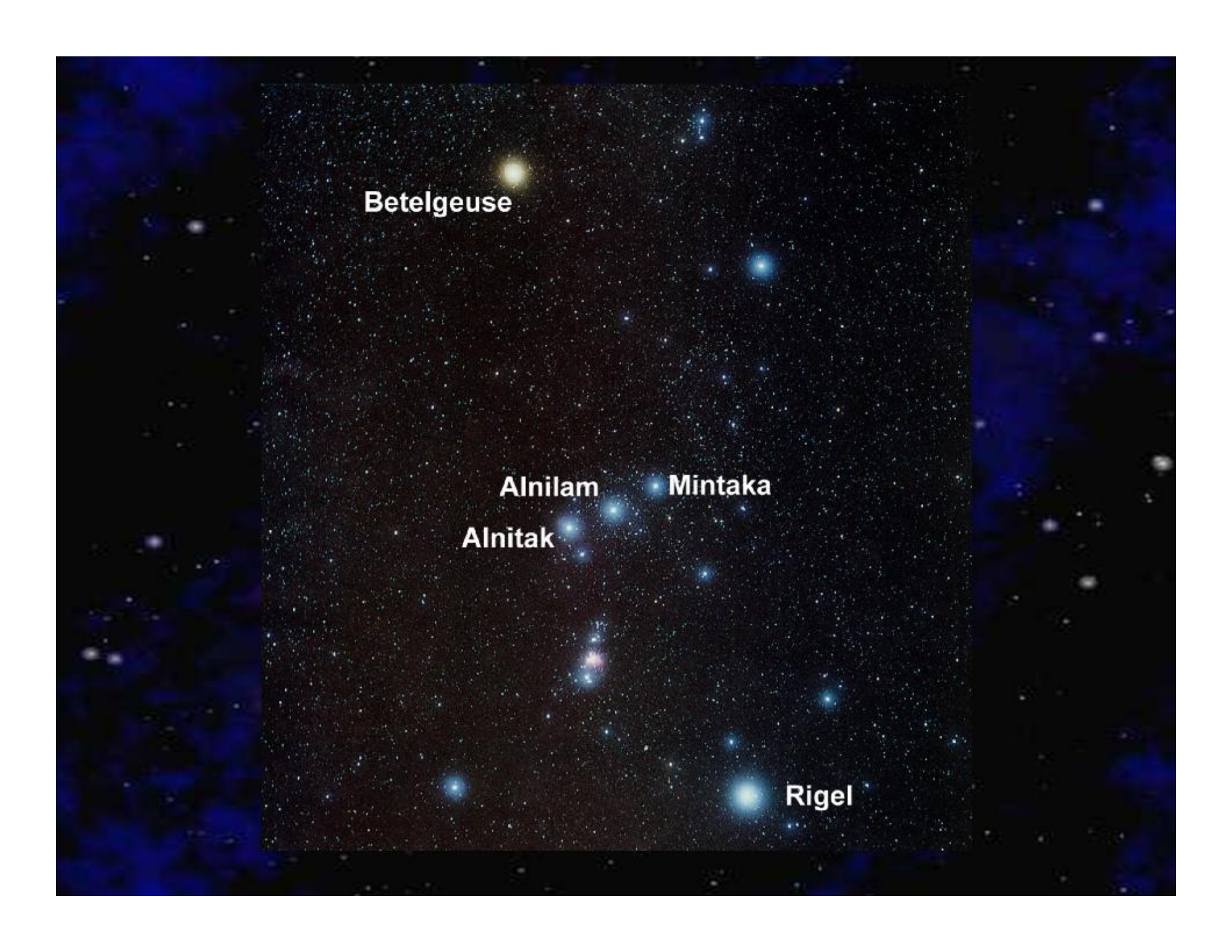


CONSTELLATIONS

A pattern of stars named after mythological animals, "characters" or objects.

A constellation name or pattern usually has no physical significance!



A star map of the Orion constellation. The background is a dark blue field filled with numerous small white stars. A central rectangular area is highlighted in black, containing the main stars of the constellation. The stars are labeled with white text: Betelgeuse (top left), Anilam (middle left), Mintaka (middle right), Alnitak (bottom left), and Rigel (bottom right).

Betelgeuse

Anilam

Mintaka

Alnitak

Rigel

Bayer System (Johann Bayer 1603)

Developed the modern system of star designations

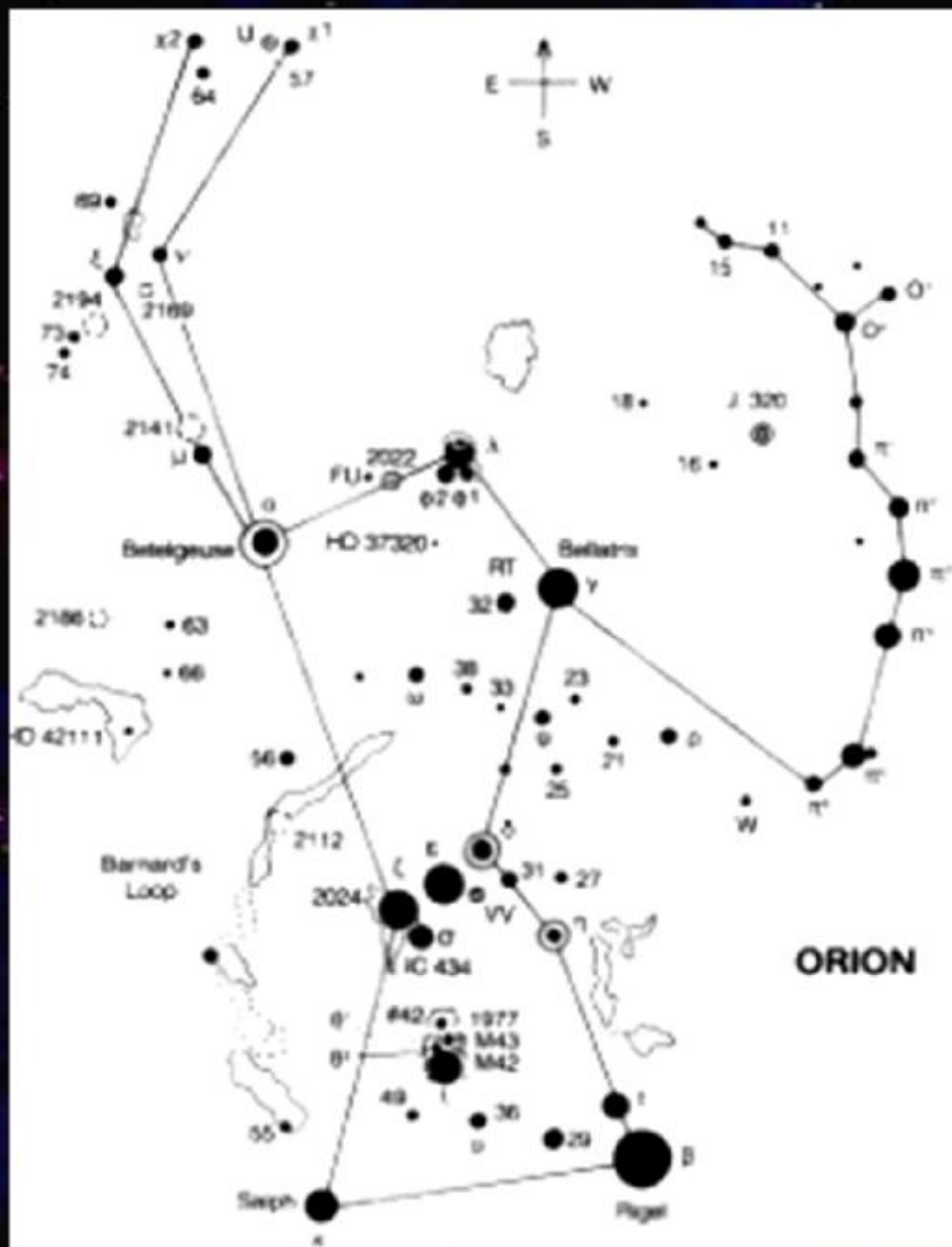
Greek Alphabet + Possessive ending:

α – brightest β – 2nd brightest γ – 3rd brightest

α Centauri

β Cygni

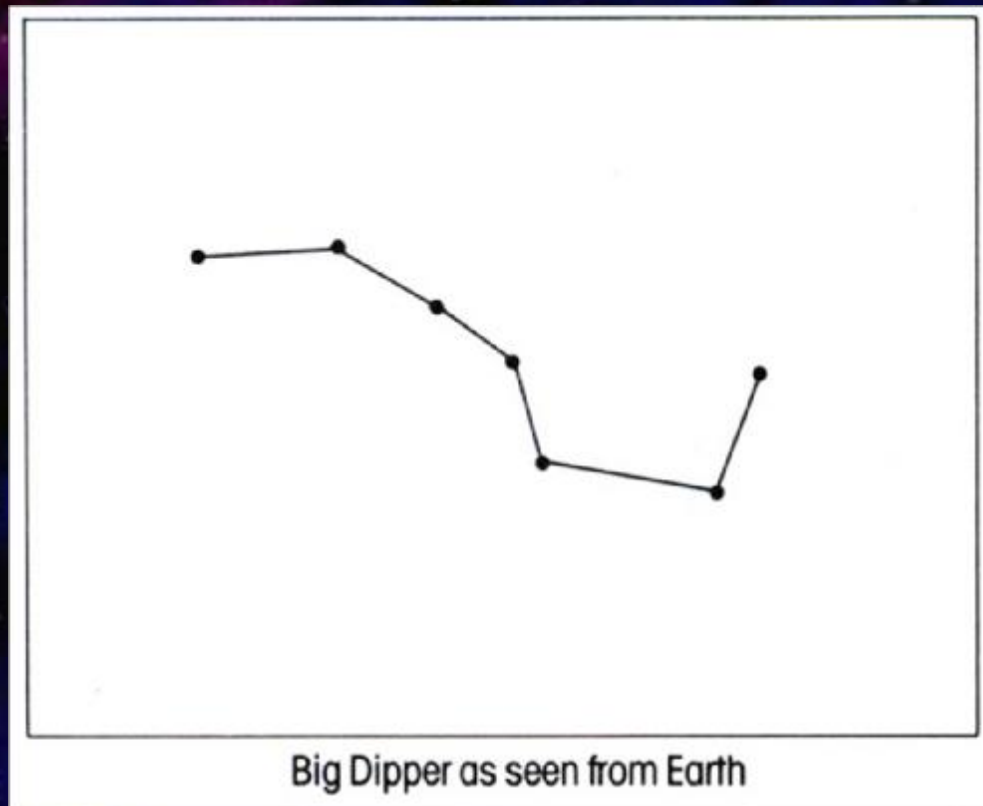
δ Cephei



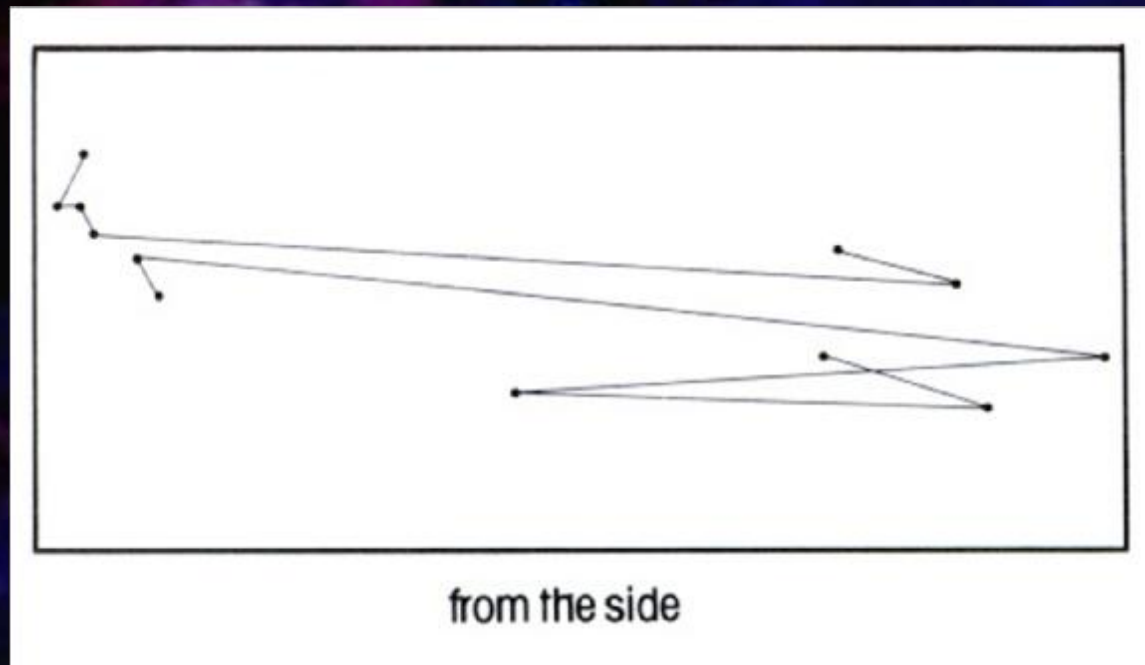
ORION



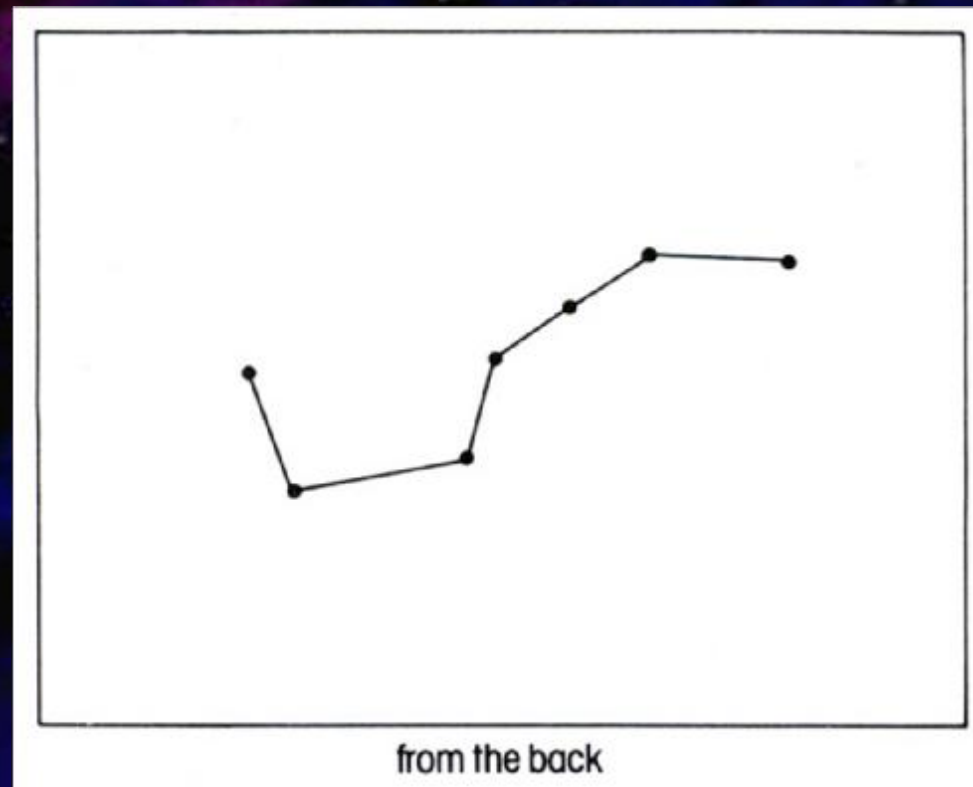
The illusion of constellations

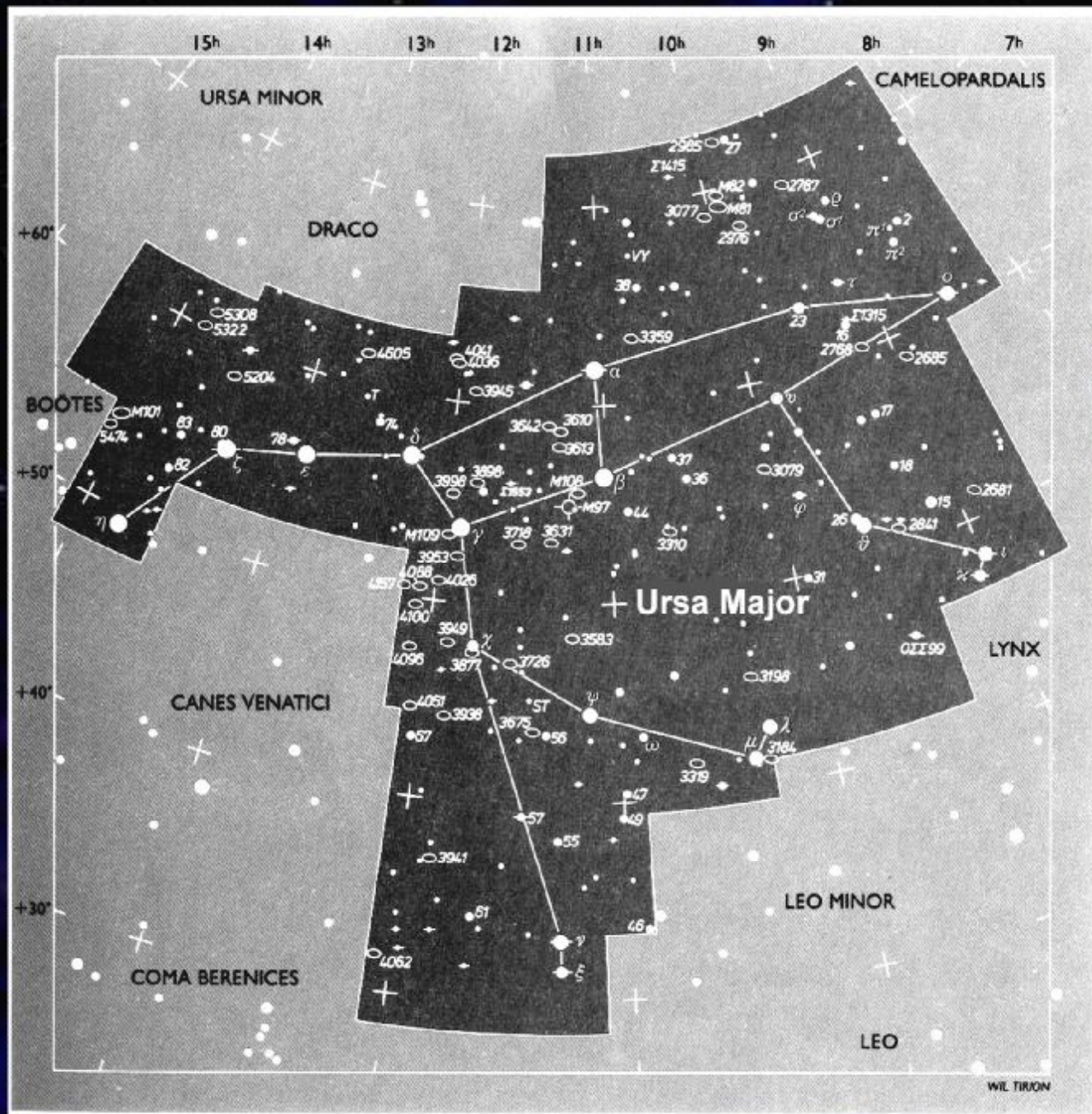


The illusion of constellations



The illusion of constellations





A dark blue night sky filled with stars and a prominent purple nebula. The nebula is a diffuse, glowing cloud of gas and dust, primarily in shades of purple and magenta, with some darker blue and black regions. It is centered in the upper half of the frame. The background is a deep blue, almost black, with numerous small, bright white and yellow stars scattered throughout. The overall appearance is that of a deep space or astronomical image.

MECHANICS OF THE SKY:

APPARENT MOTIONS

SOLAR DAY:

The time it takes the Sun to complete two successive crossings of the meridian.

24 hours

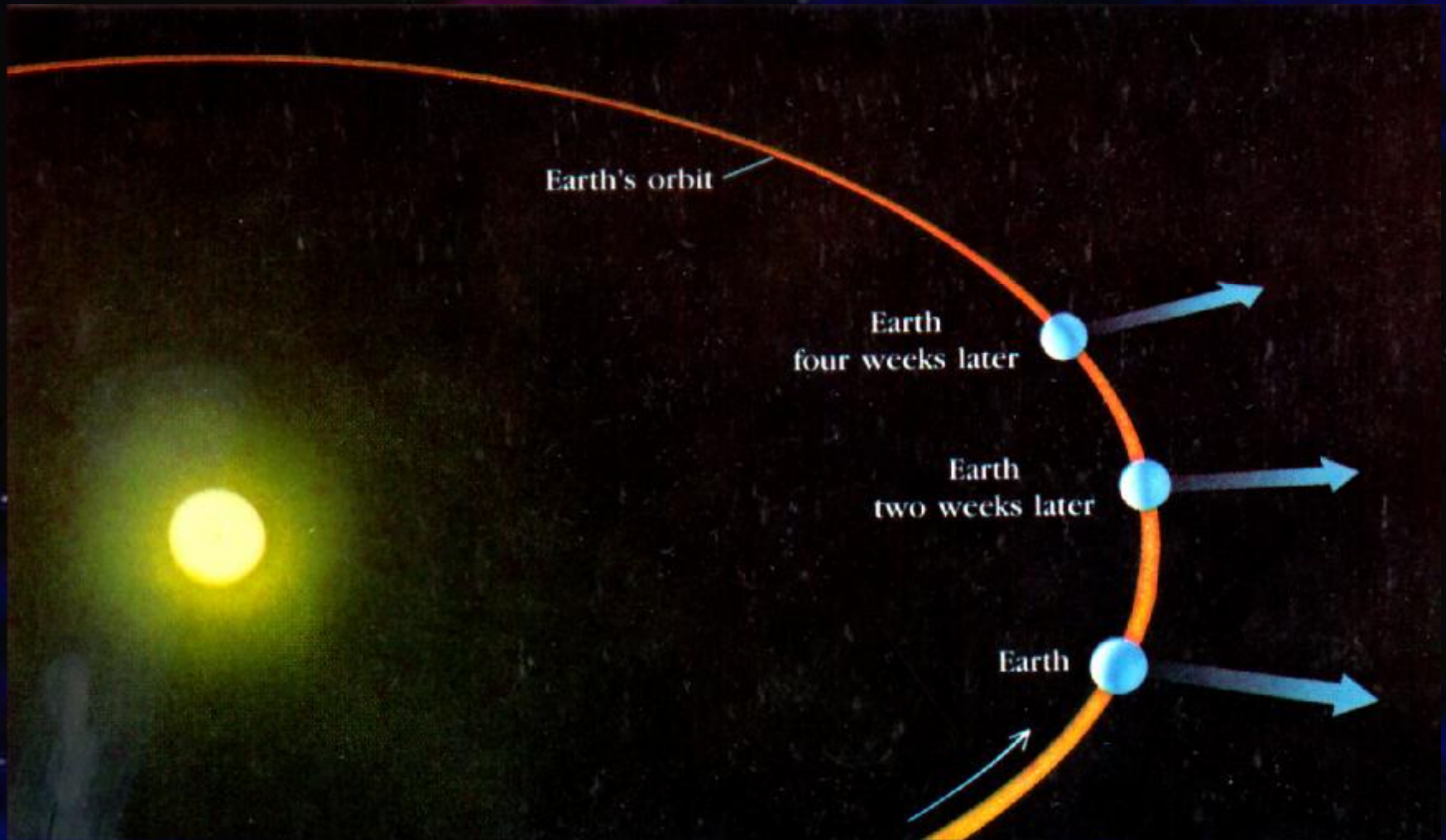
Caused by the ROTATION of the Earth on its axis

SIDEREAL DAY:

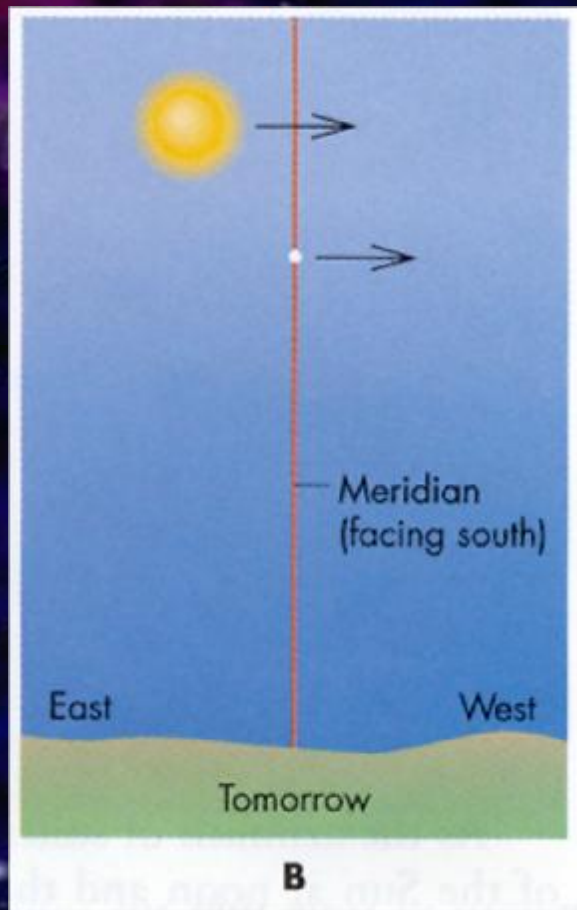
The time it takes for two successive crossings of a celestial object (other than the Sun, Moon or planets) across the meridian.

23 hours 56 minutes 4.091 seconds

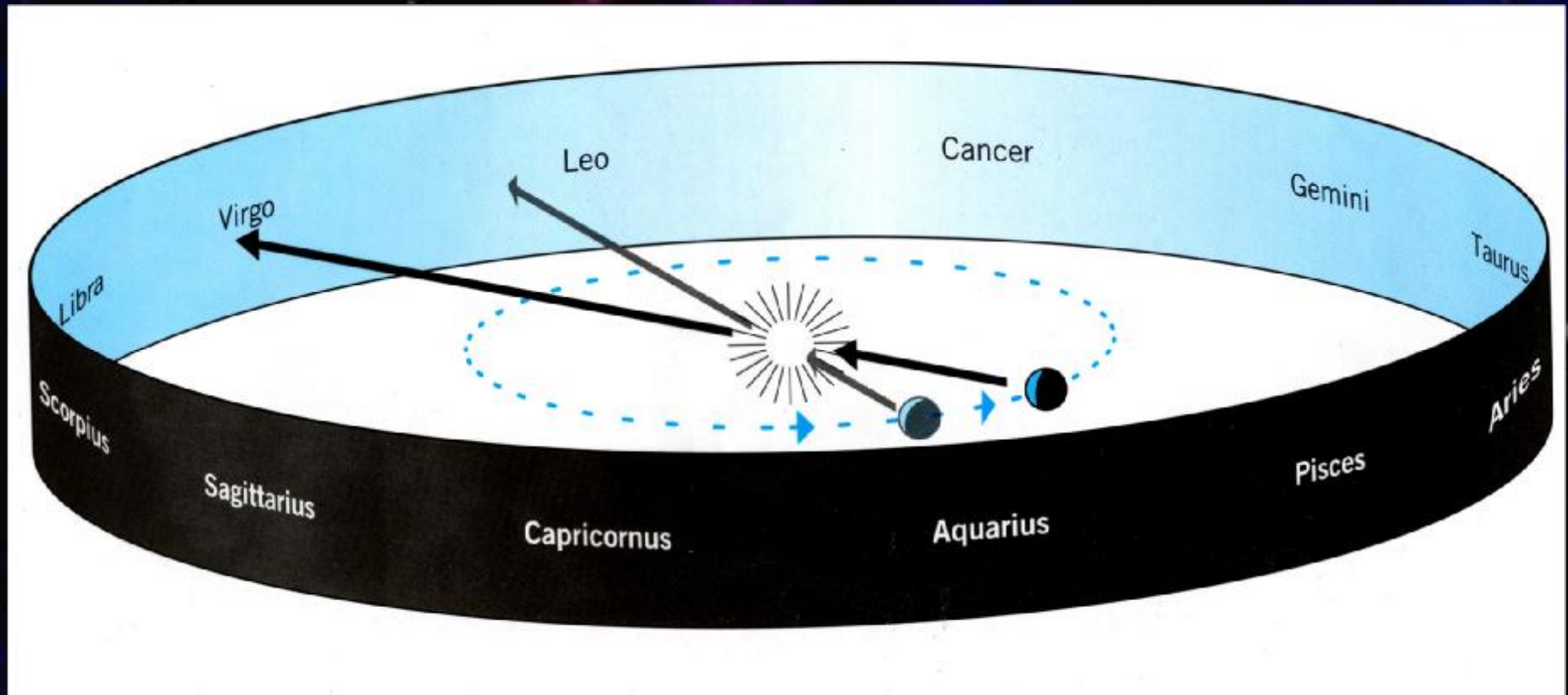
Solar Day vs. Sidereal Day



Solar Day vs. Sidereal Day



ANNUAL MOTION OF THE SUN



§ *ECLIPTIC* – the apparent path the Sun travels across the sky.

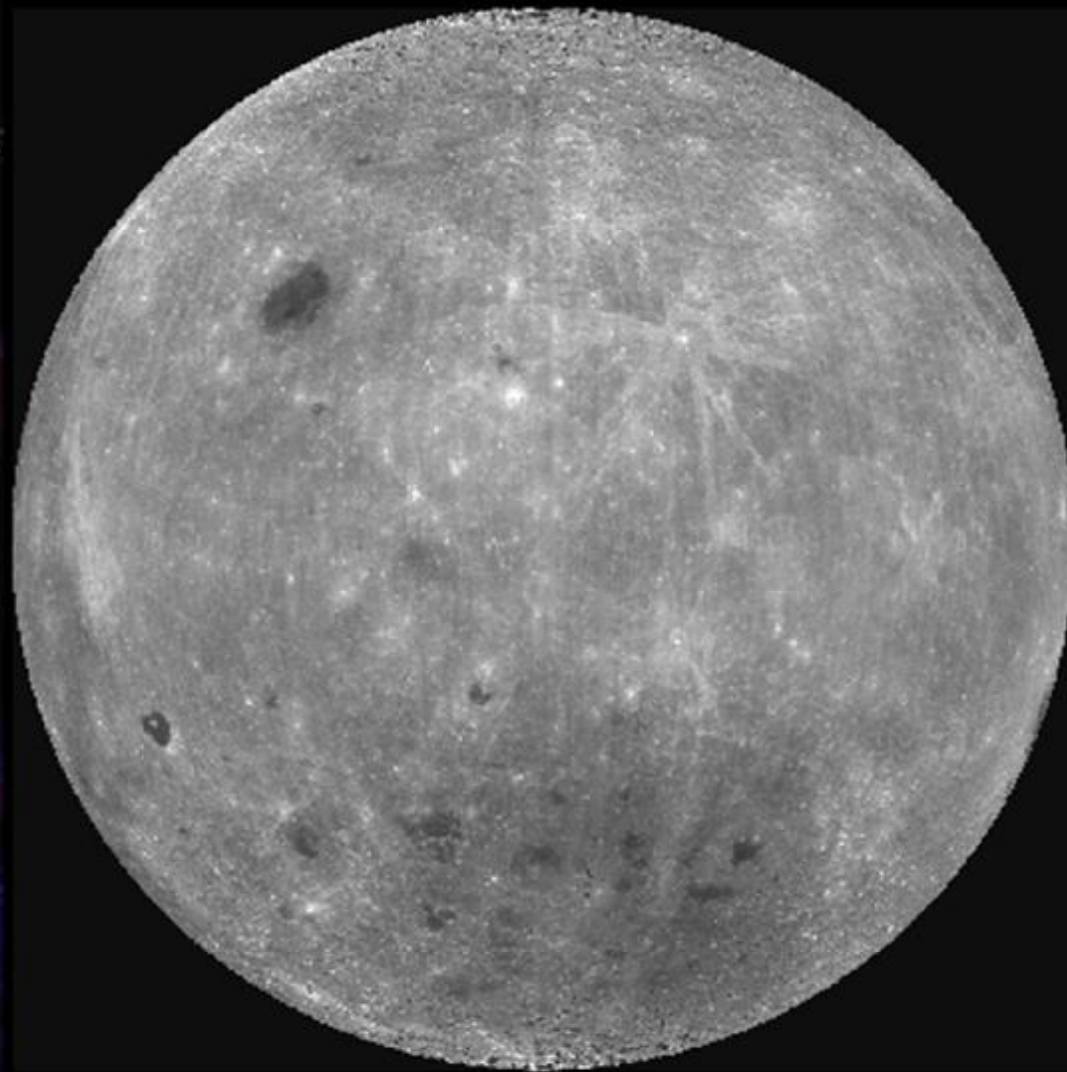








Near Side



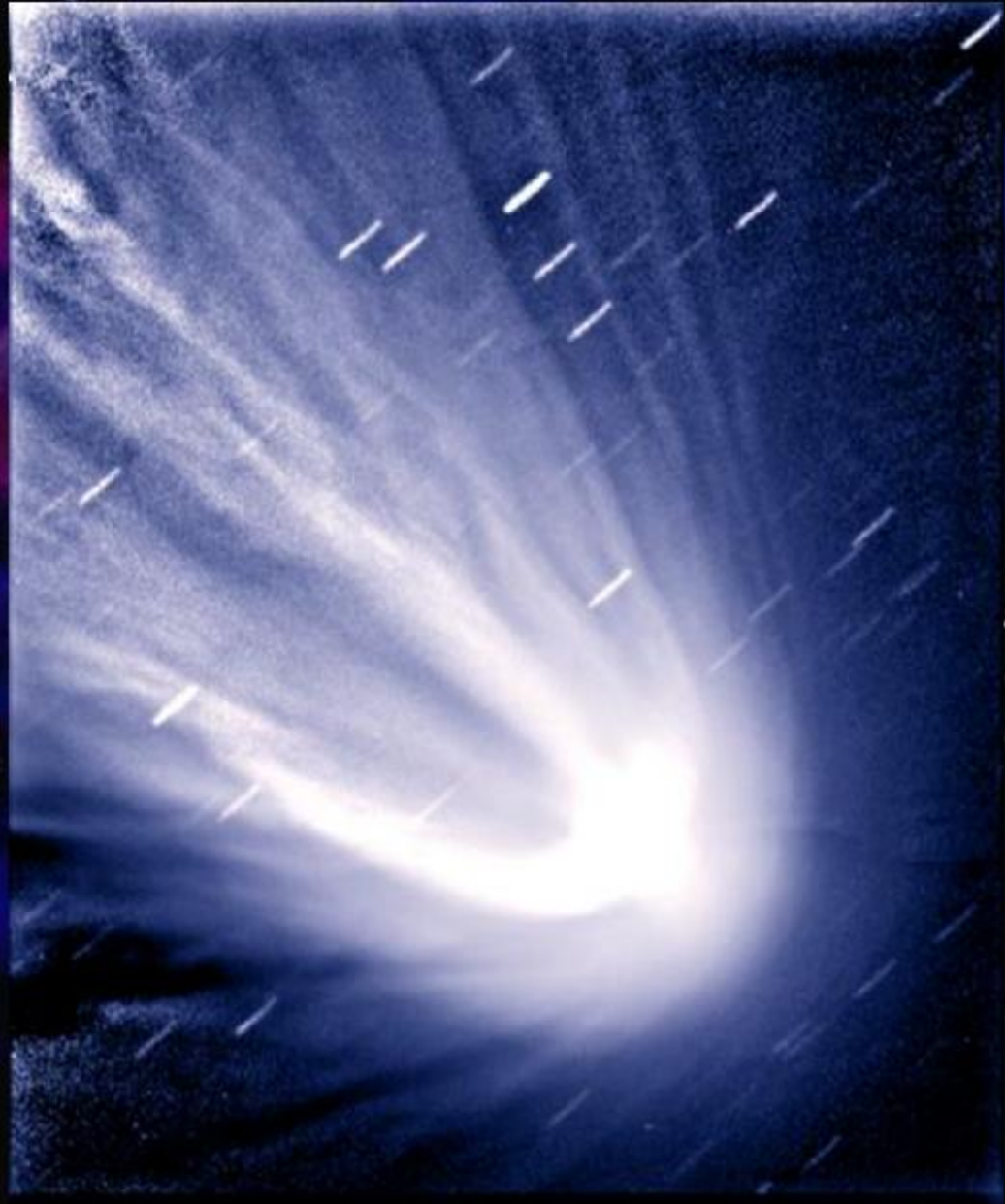
Far Side

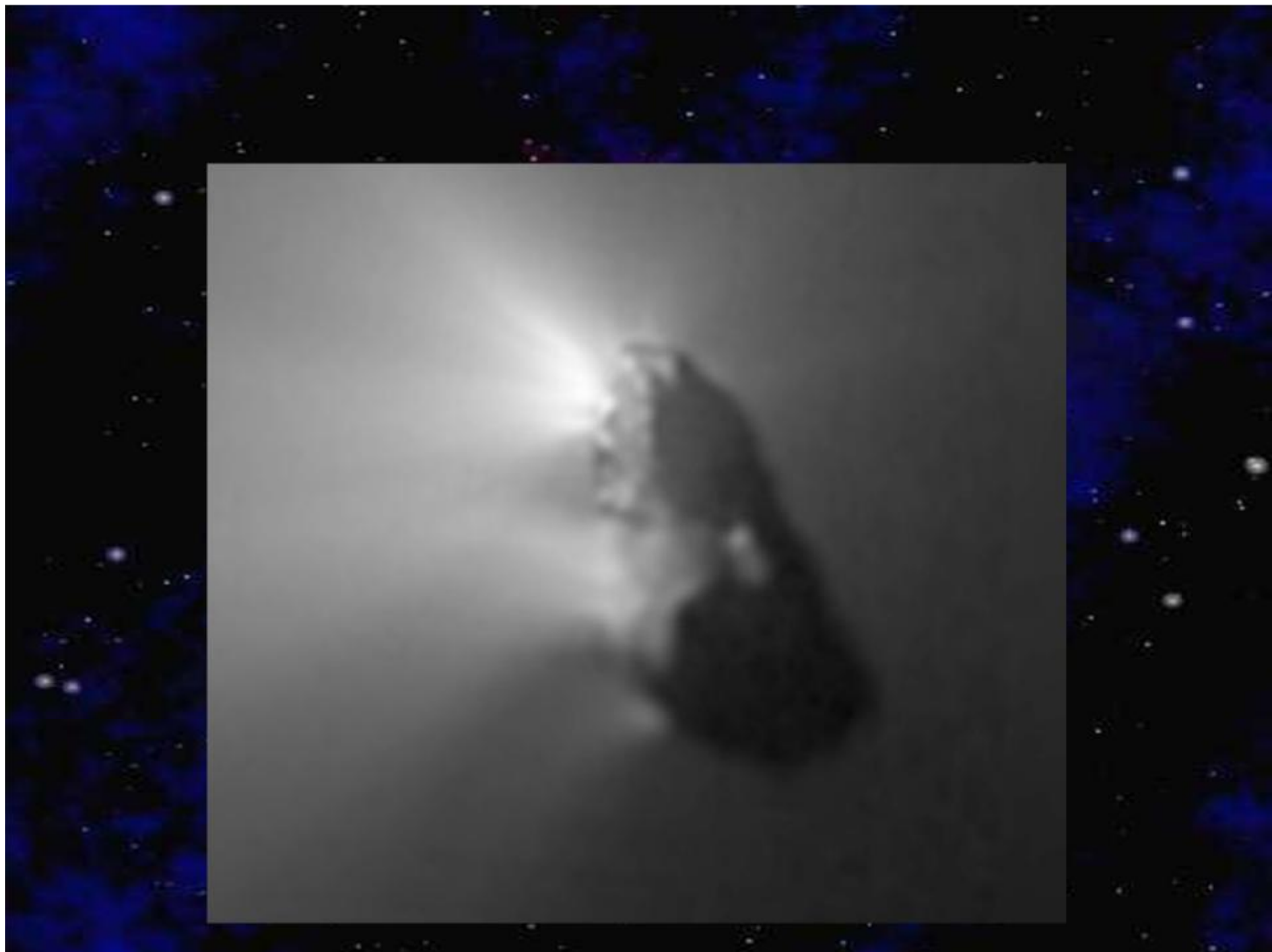
Lunar Craters



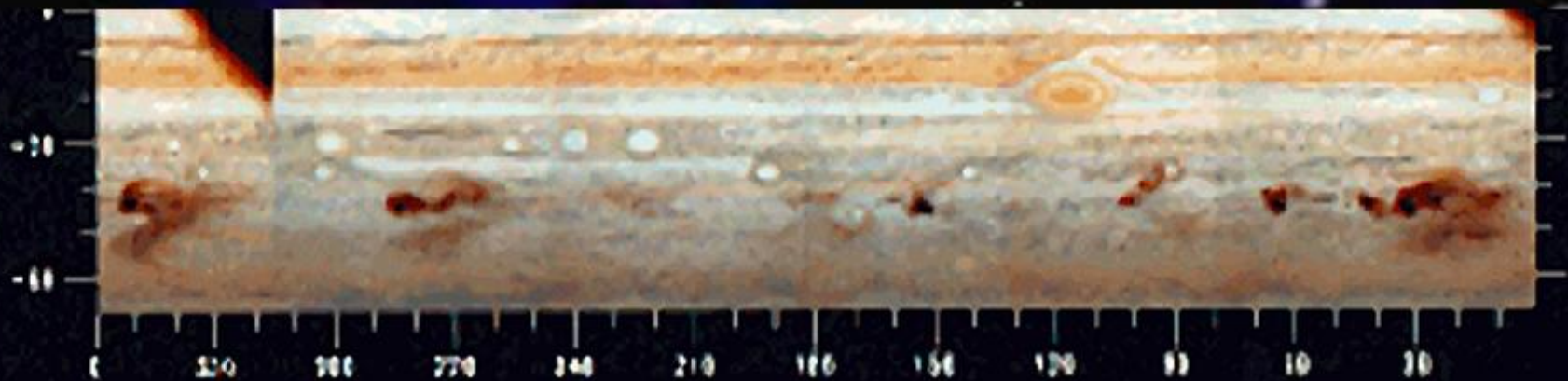


Hale-Bopp, 1996





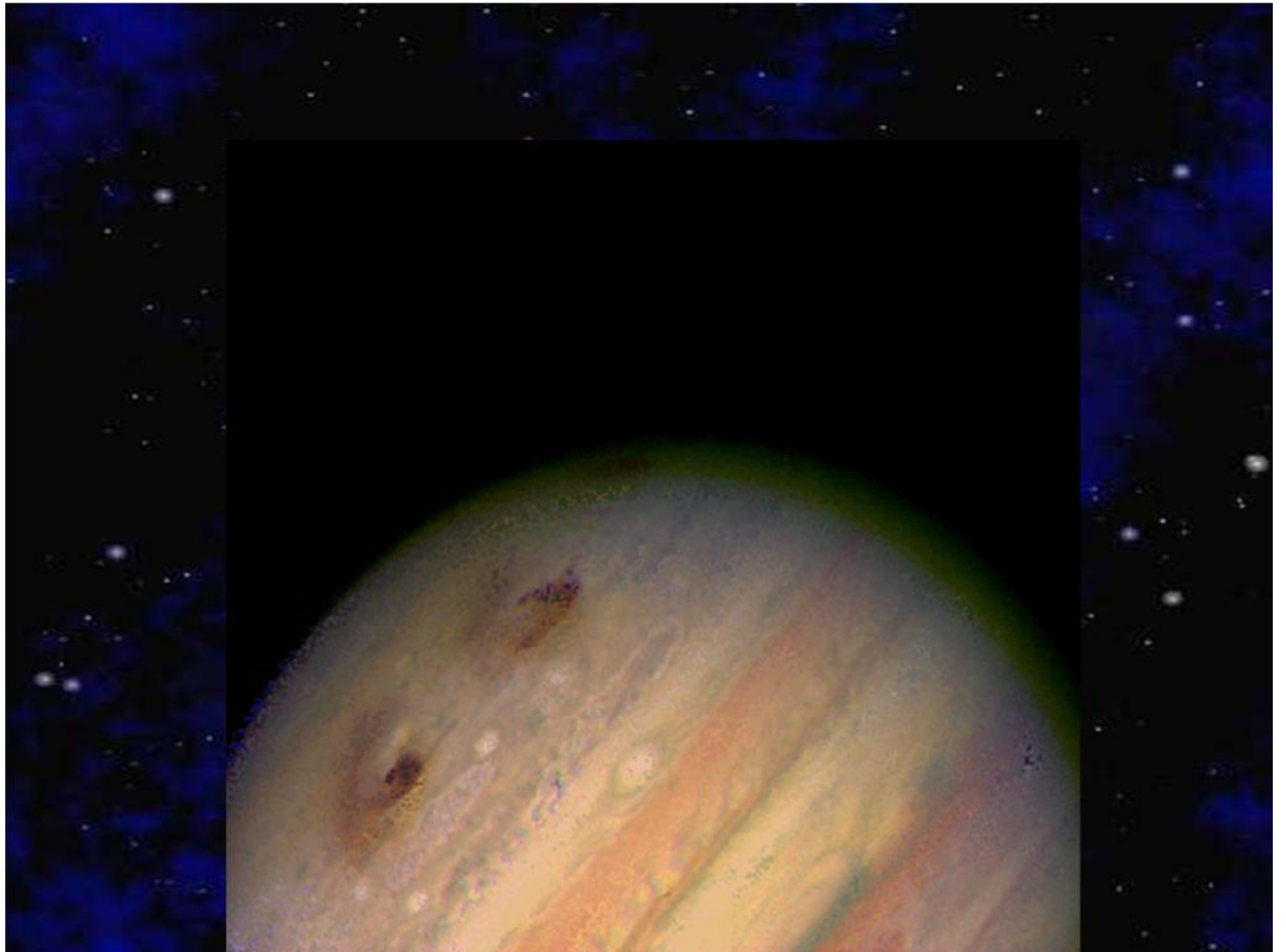
Comet Shoemaker-Levy

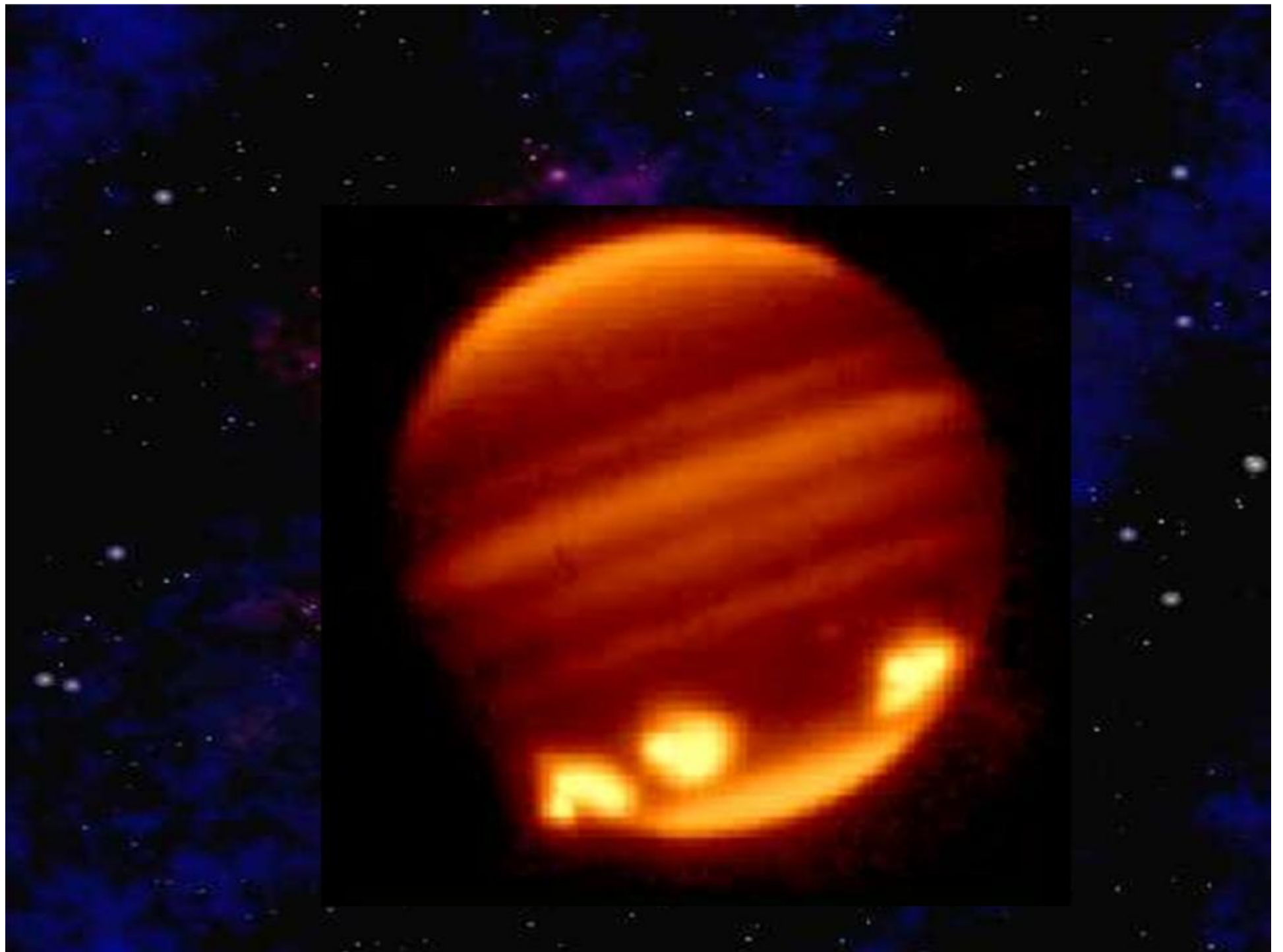


Comet Shoemaker-Levy

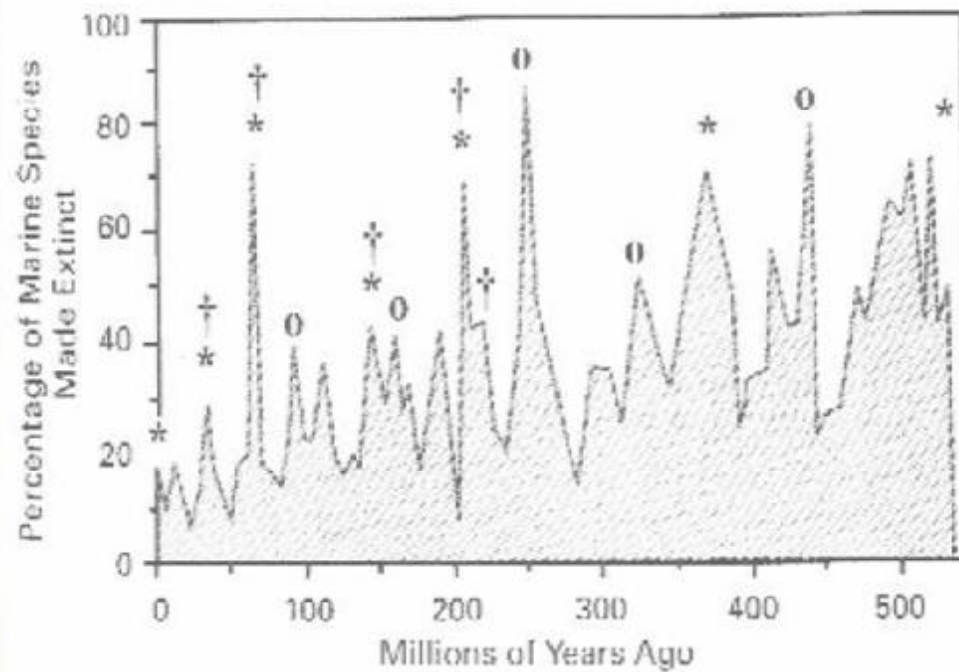


www.spacetelescope.org





Mass Extinctions in 540 Million Years



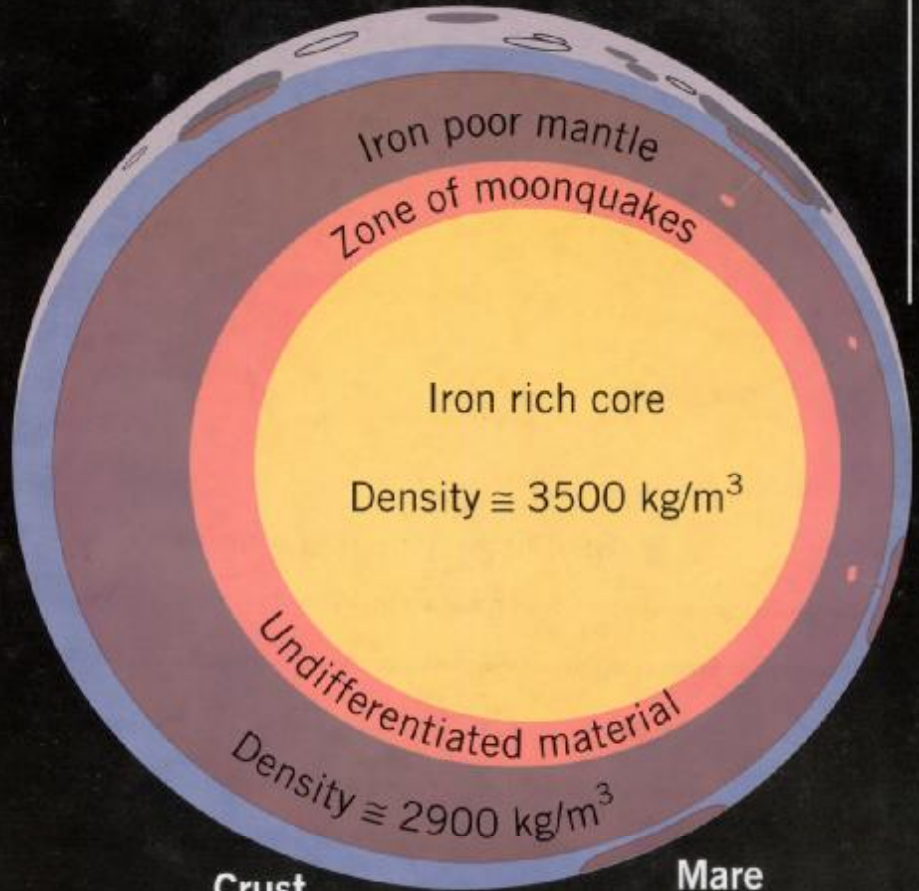


Lunar Maria



Moon

1740 km



Iron rich core

Density $\approx 3500 \text{ kg/m}^3$

Undifferentiated material

Density $\approx 2900 \text{ kg/m}^3$

Iron poor mantle

Zone of moonquakes

Crust

Mare

To earth

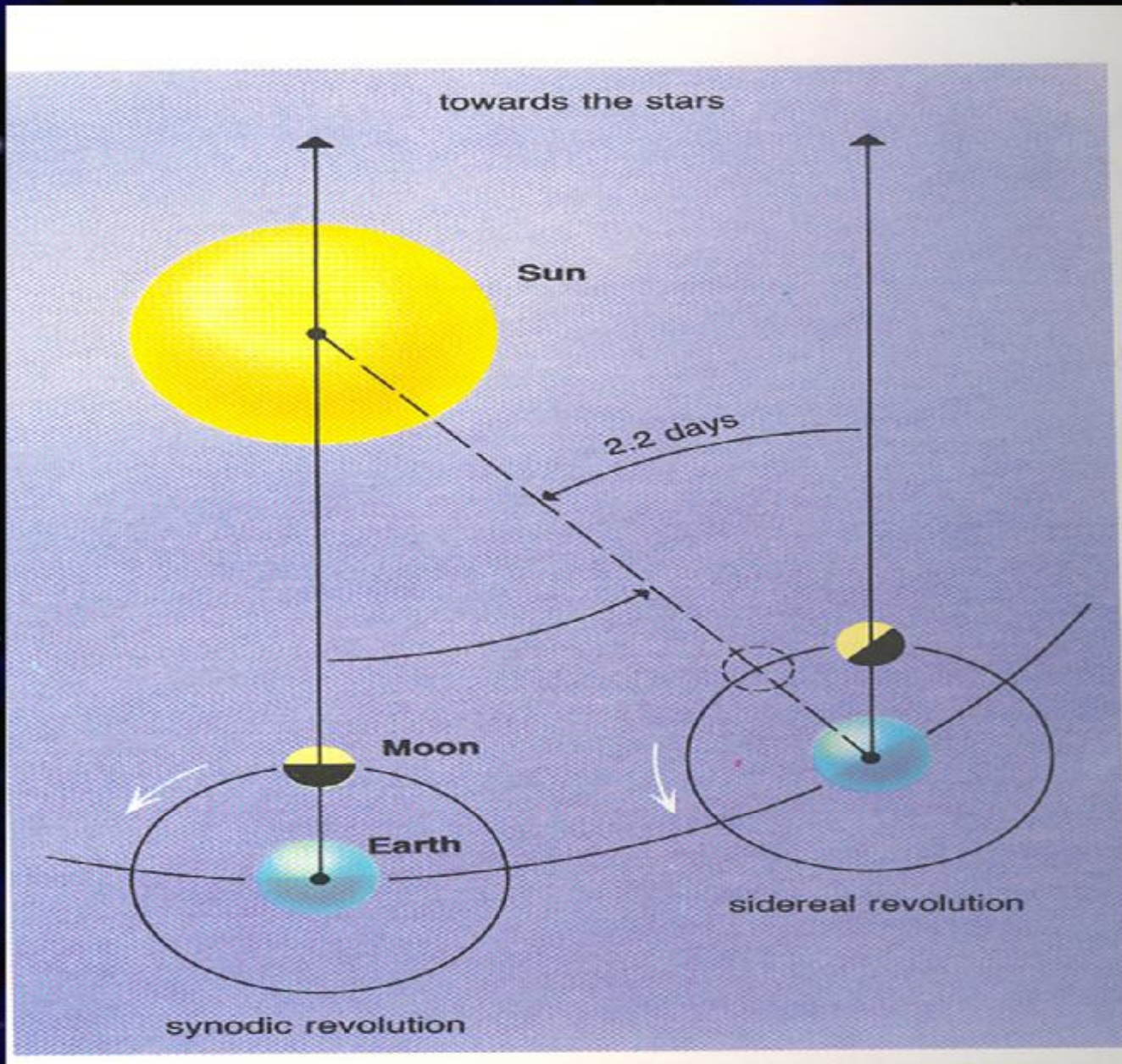
Motion of the Moon with respect to the Celestial Sphere:

§ Diurnal Motion (east –to- west)

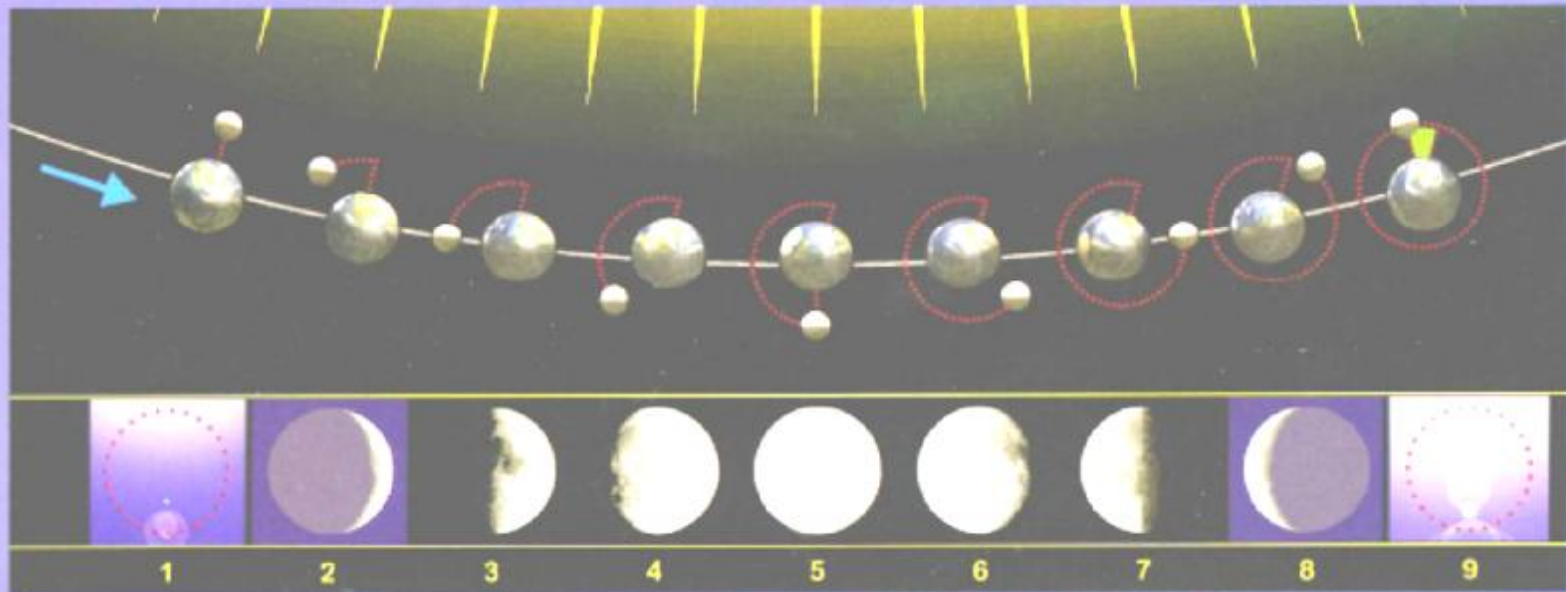
§ Monthly (west –to- east)

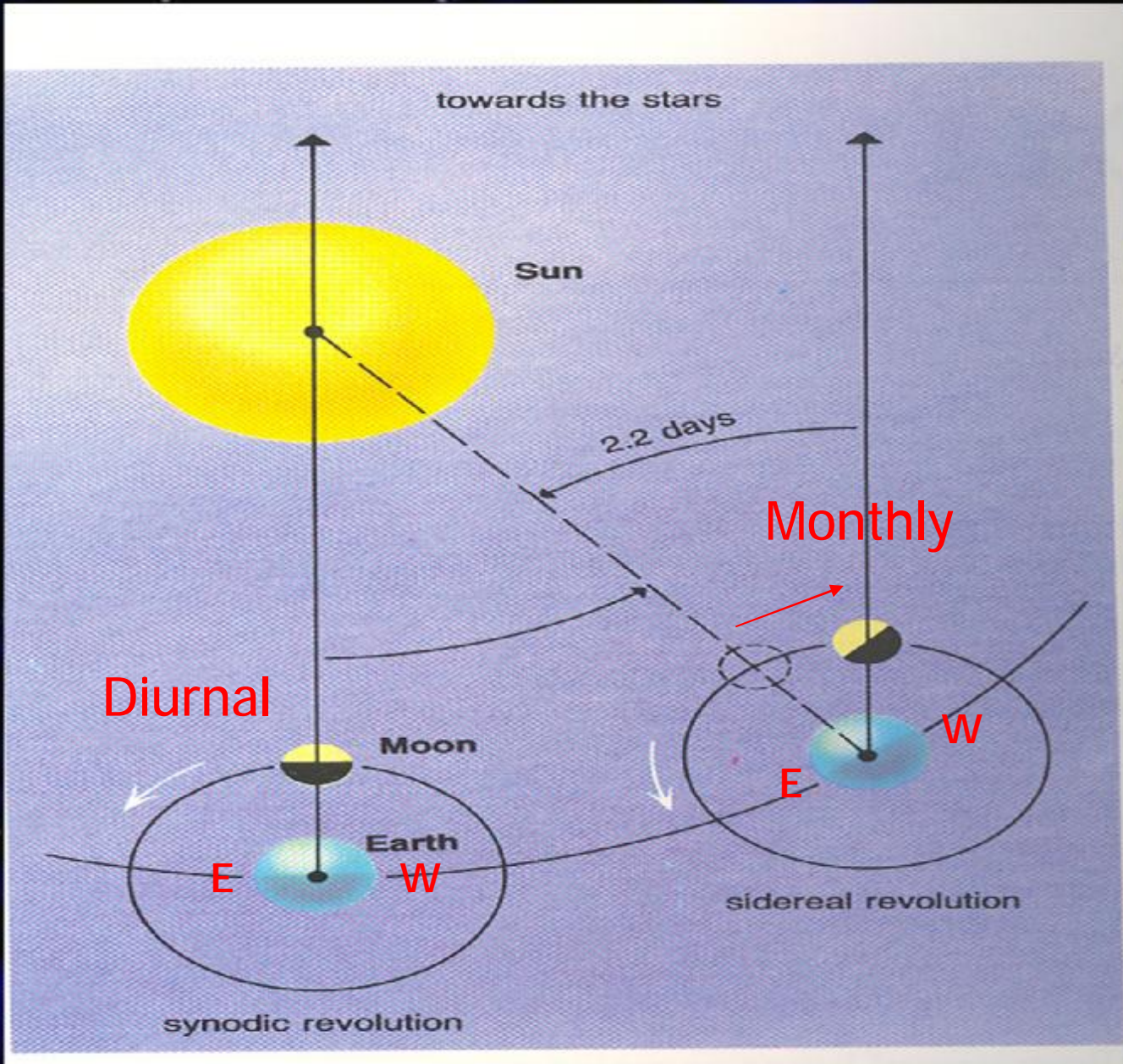
§ Moves $\sim 13^\circ$ each day

$360 \text{ degrees in a circle} / 28 \text{ days} = 12.9 \text{ degrees/day}$



Moon Phases or Moon Shine?







LUNAR PHASES

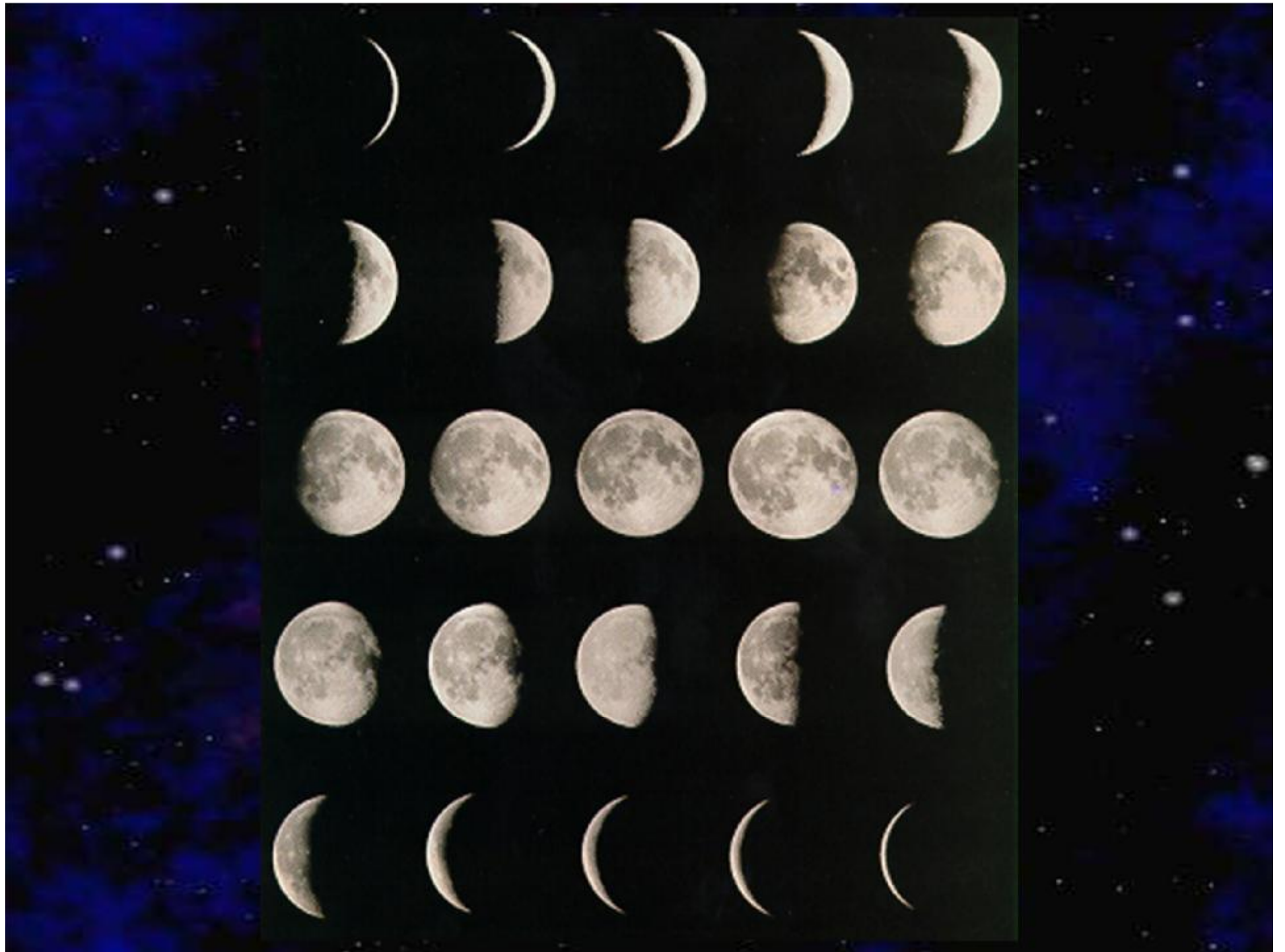
Mother Goose & Grimm

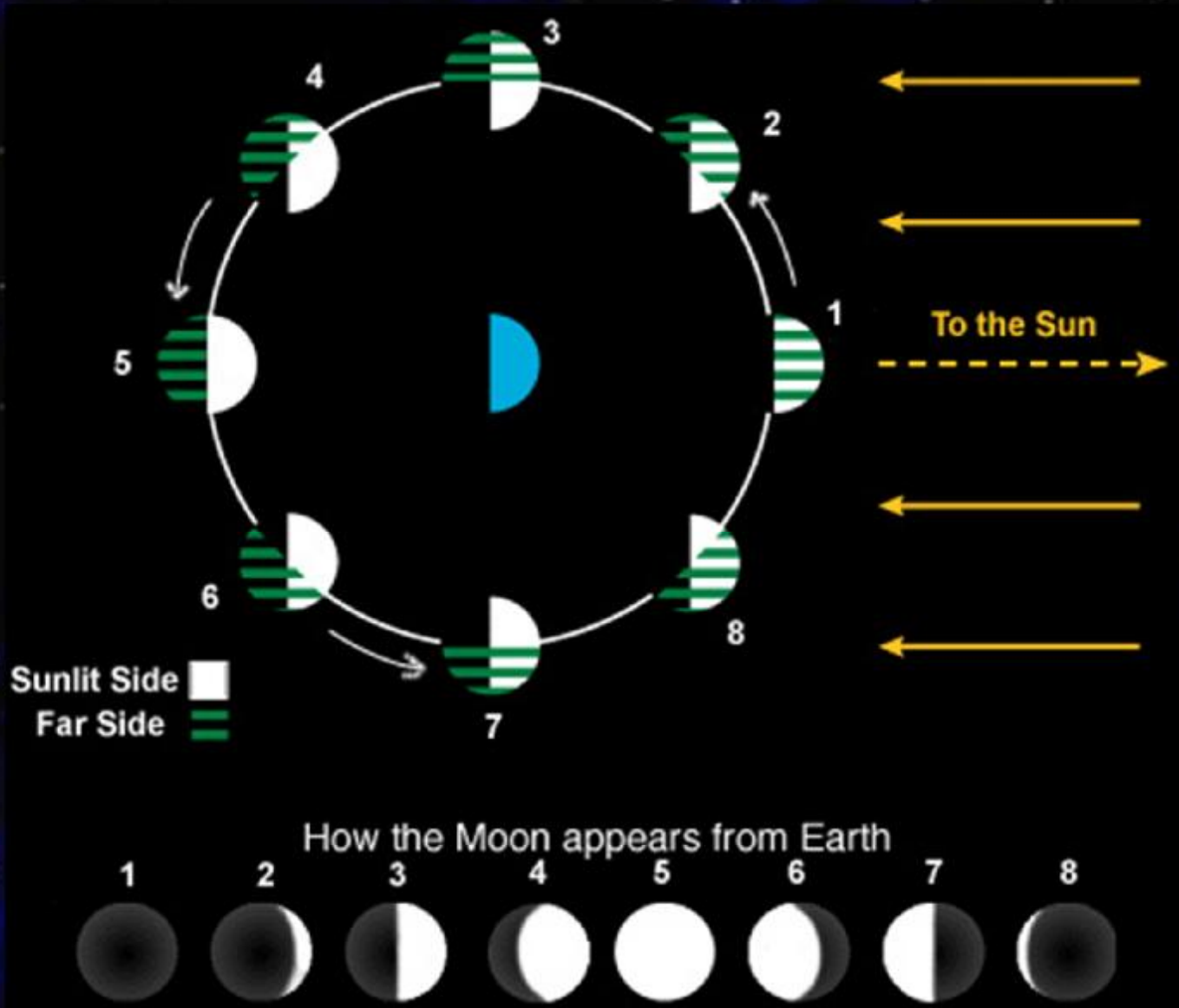
By Mike Peters



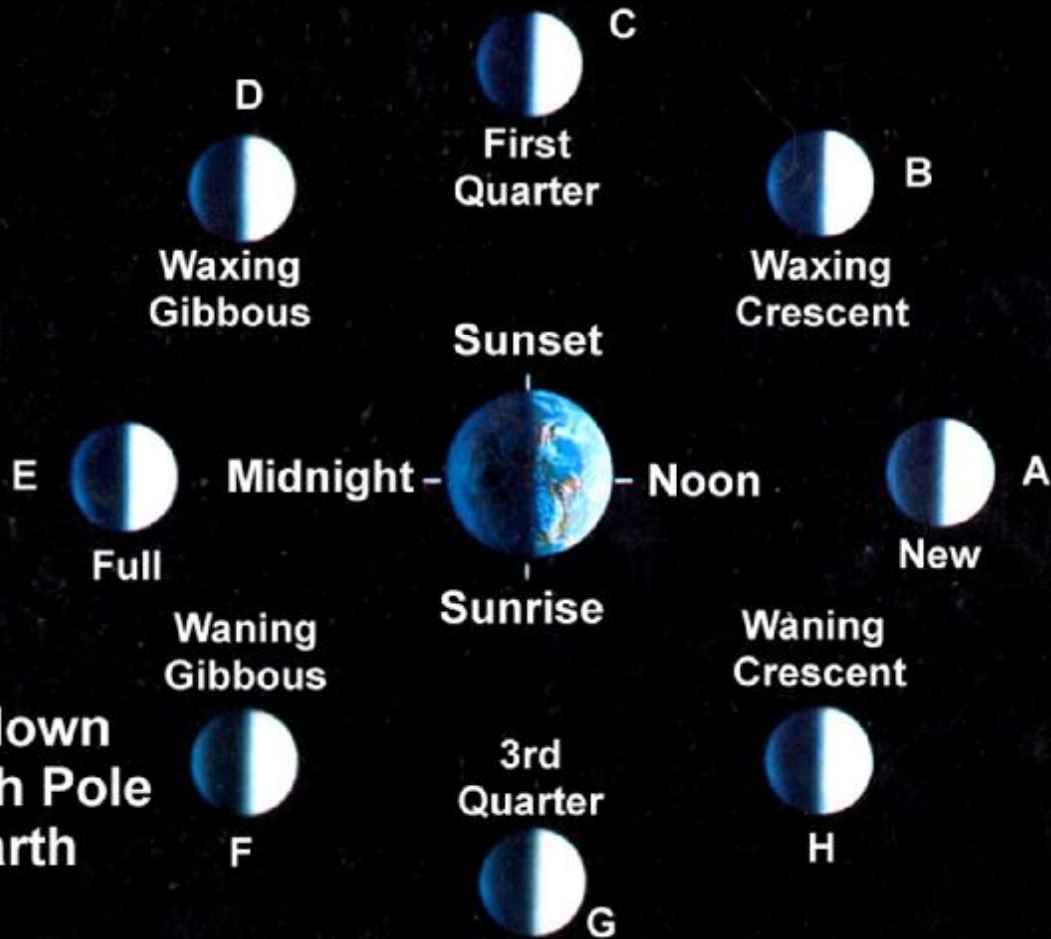
Lunar Phases:

Caused by the change in the orientation between the Earth, Sun, & Moon

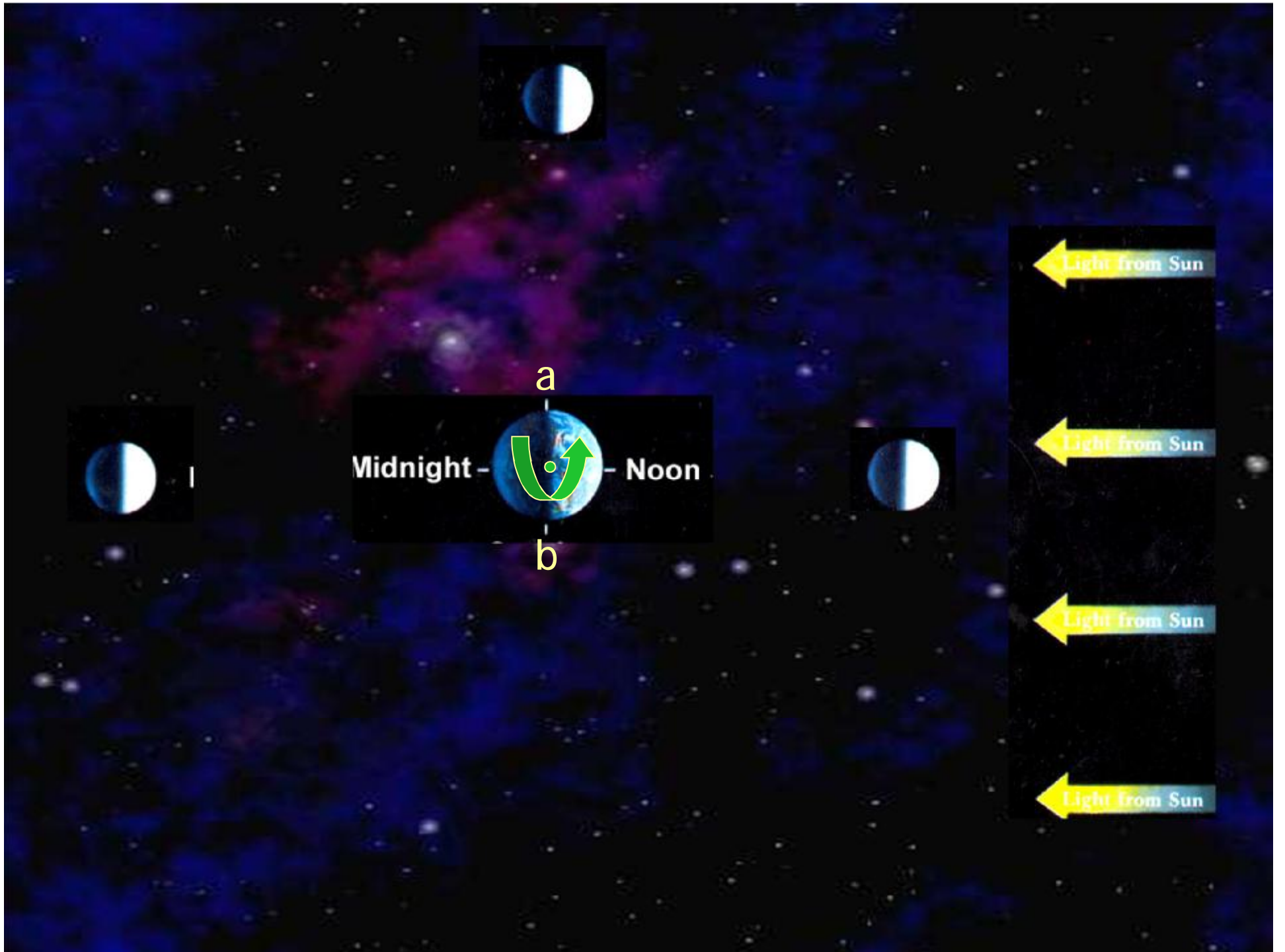




The Moon as seen from Earth



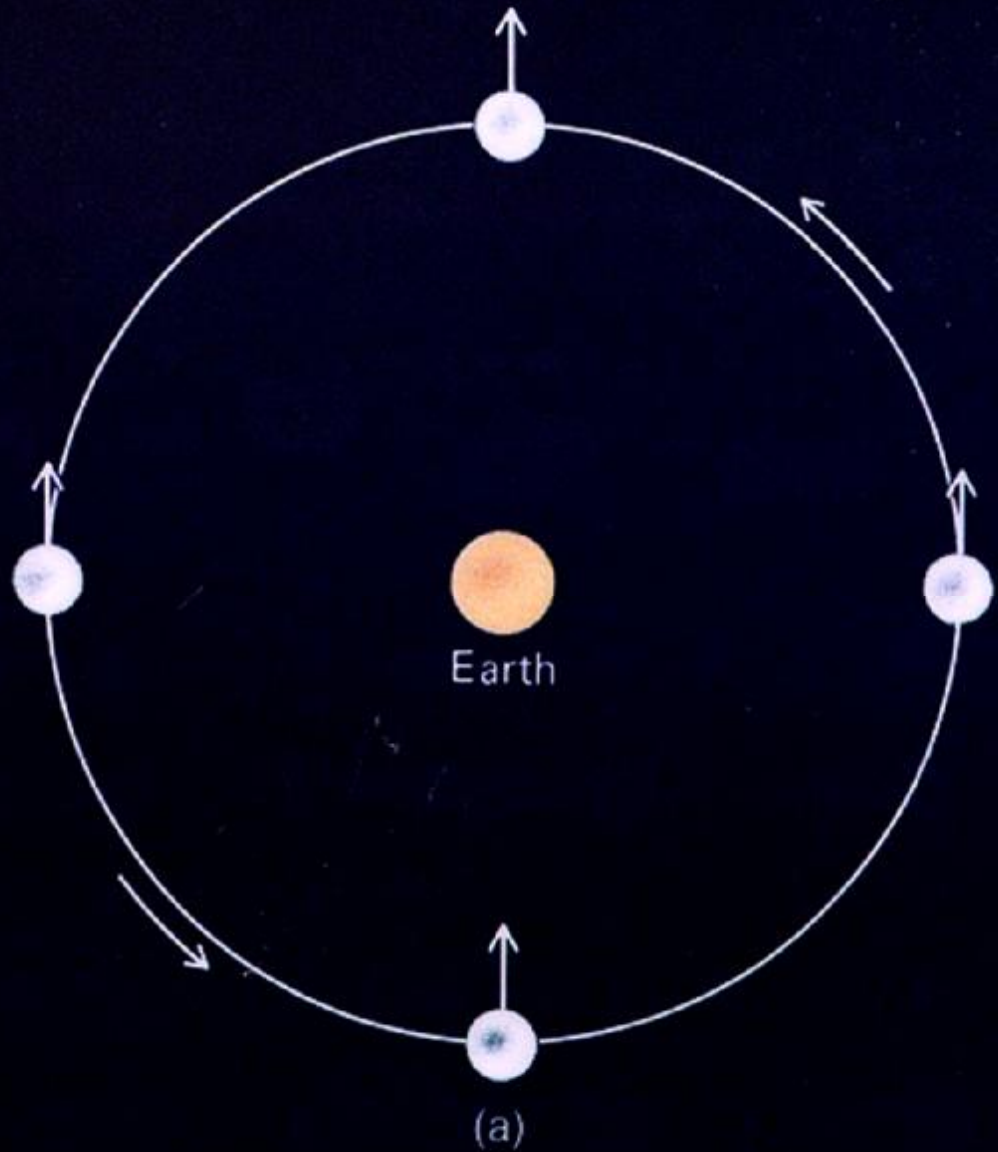
Looking down on the North Pole of the Earth



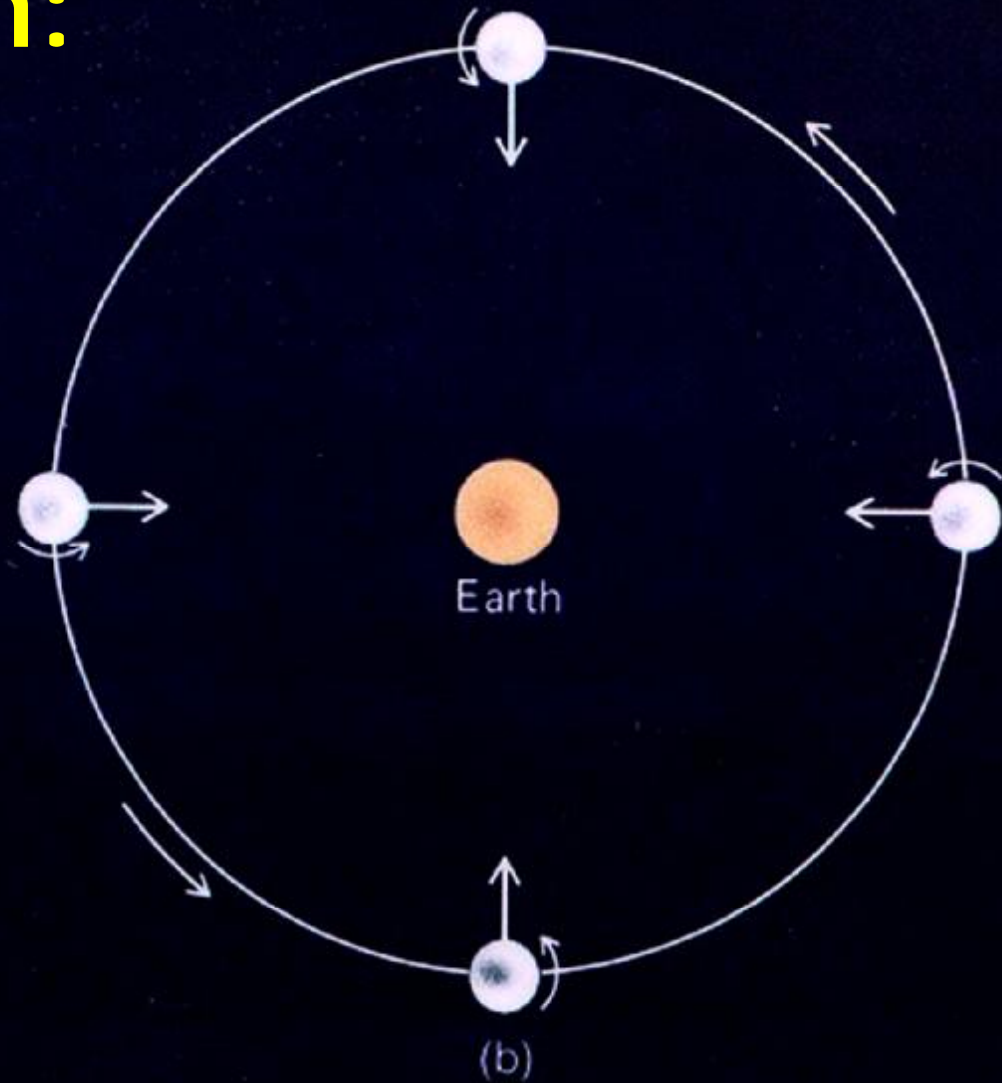


**Does the Moon Rotate on its
Axis?**

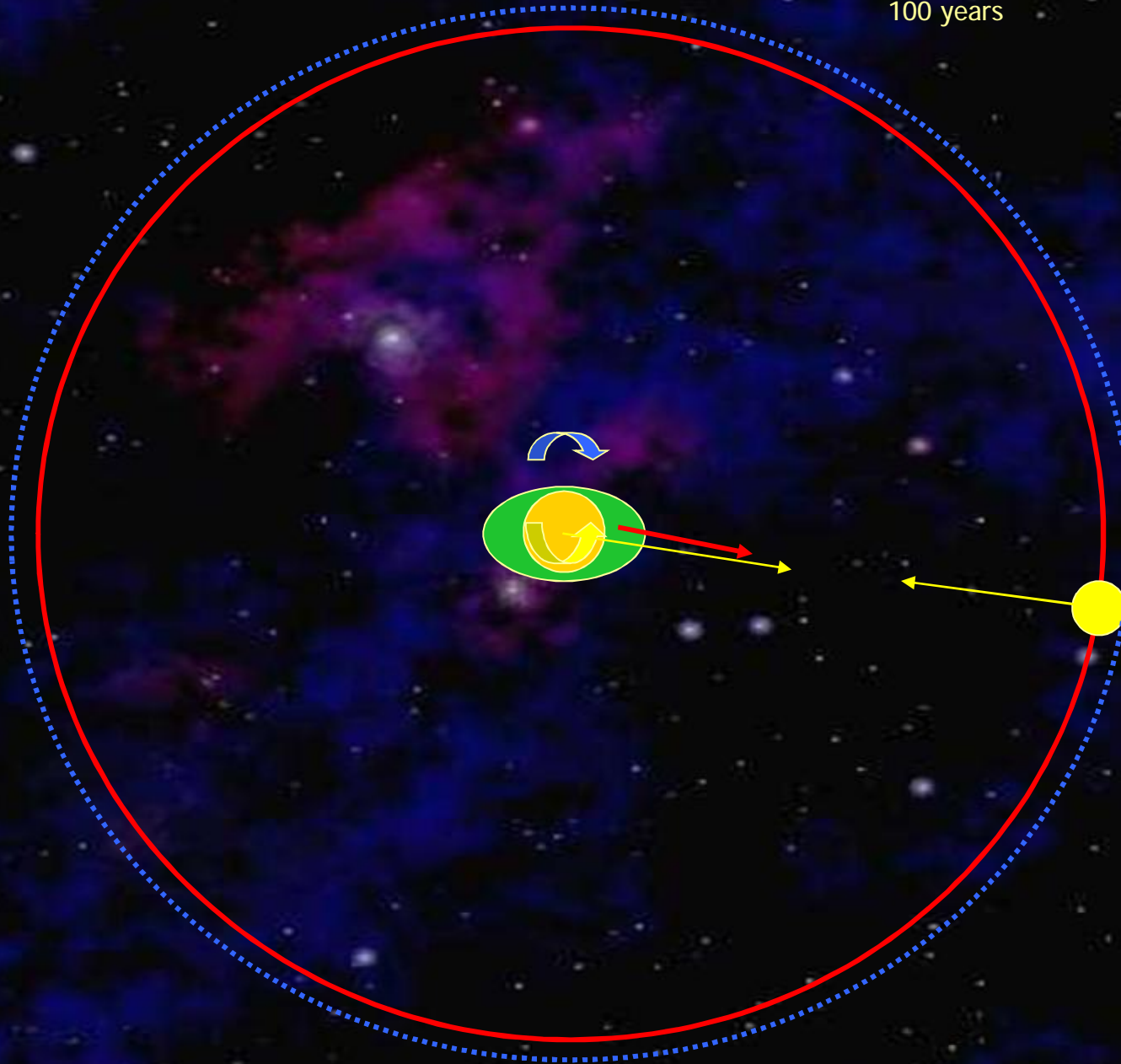
No rotation:



With rotation:



Day lengthens 0.0017 seconds per 100 years
Moon moves away from earth 3.84 m per 100 years



SHORT TERM CHANGES



THE SEASONS

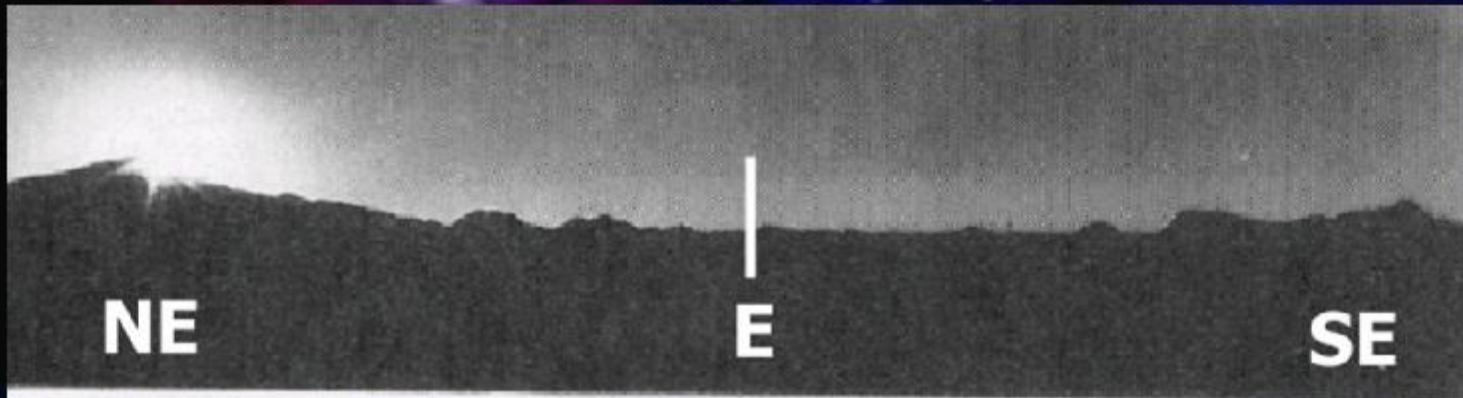
The revolution of the Earth around the Sun:

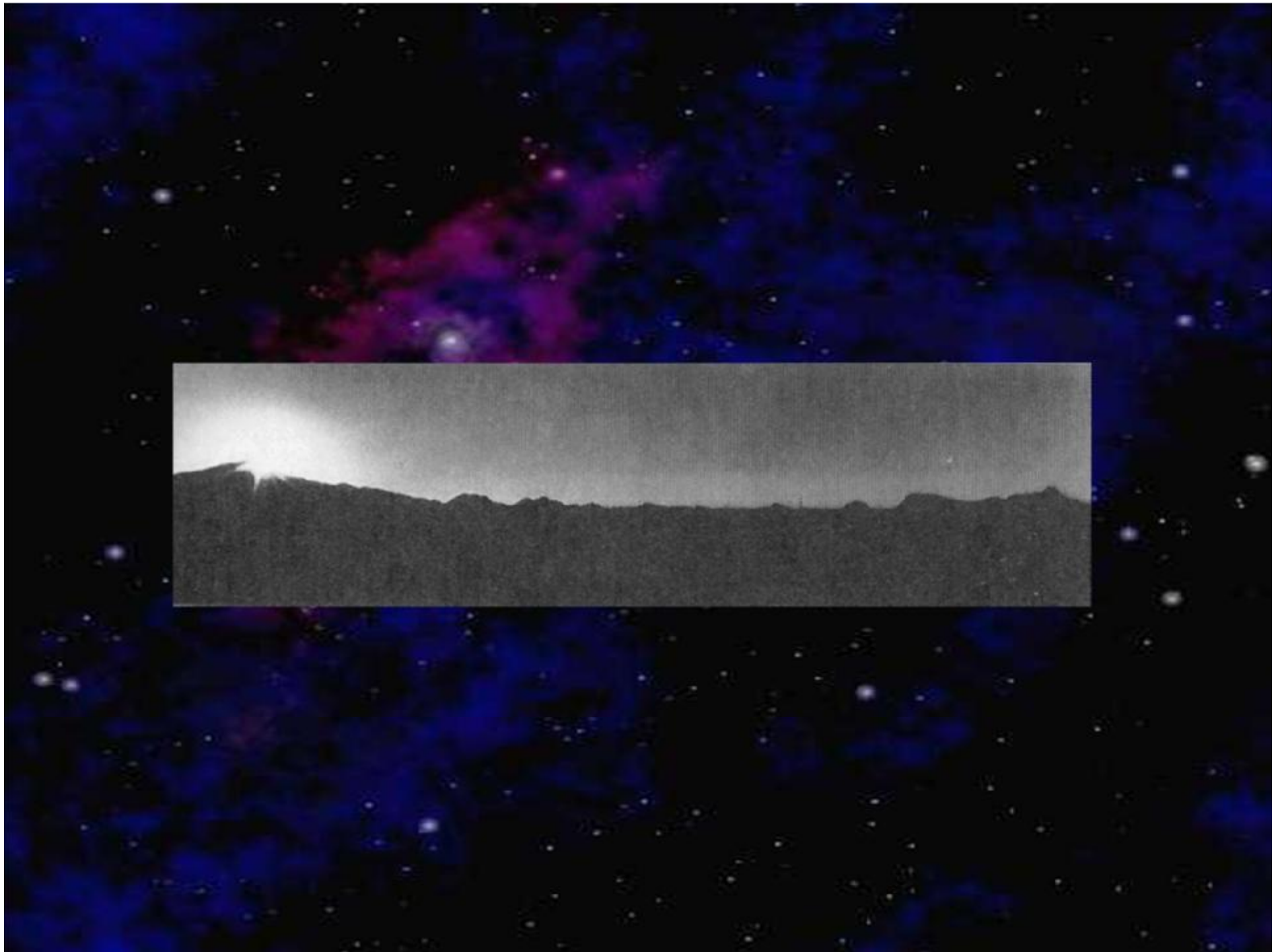
Tropical Year:

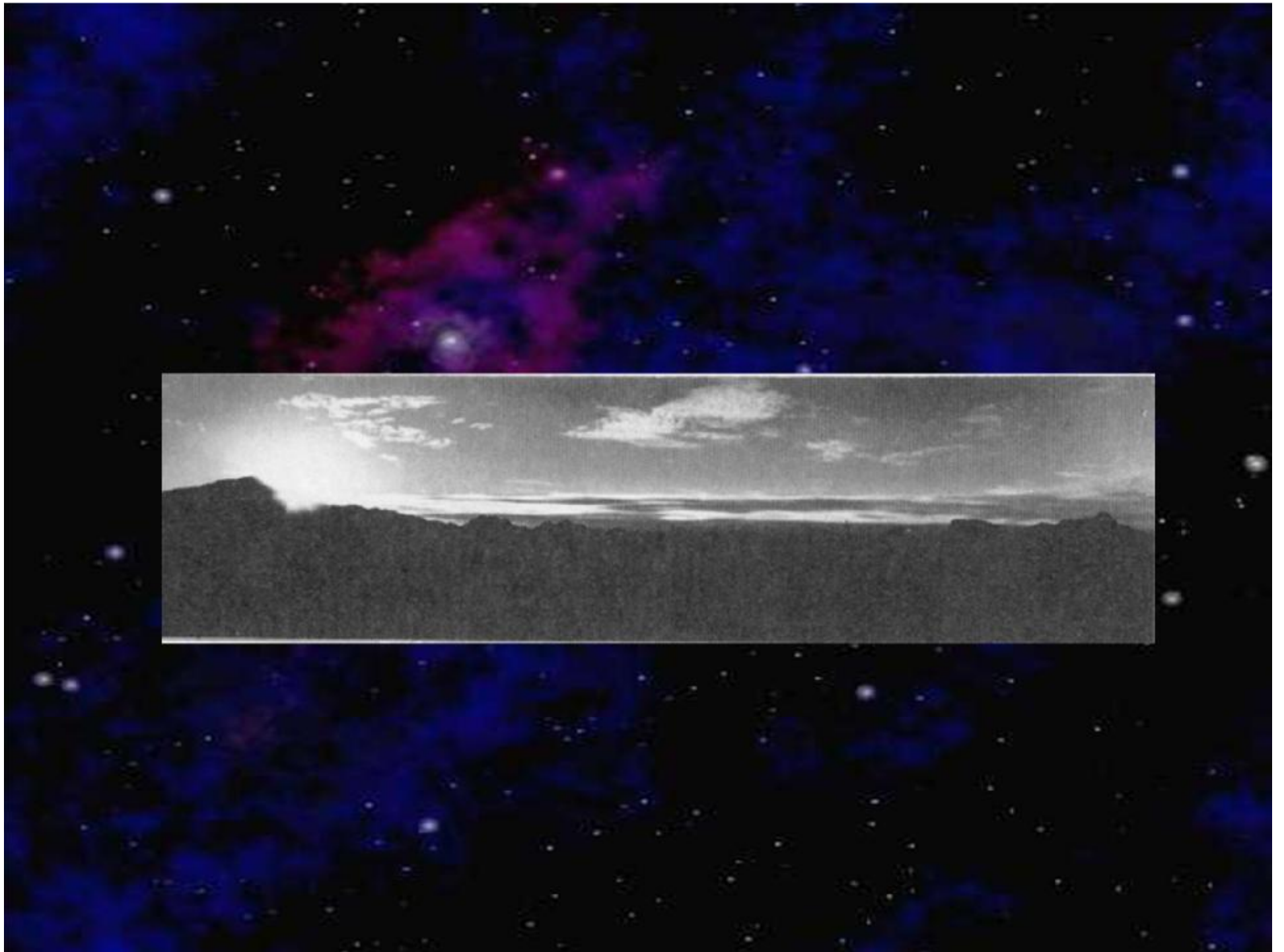
365.2422 days

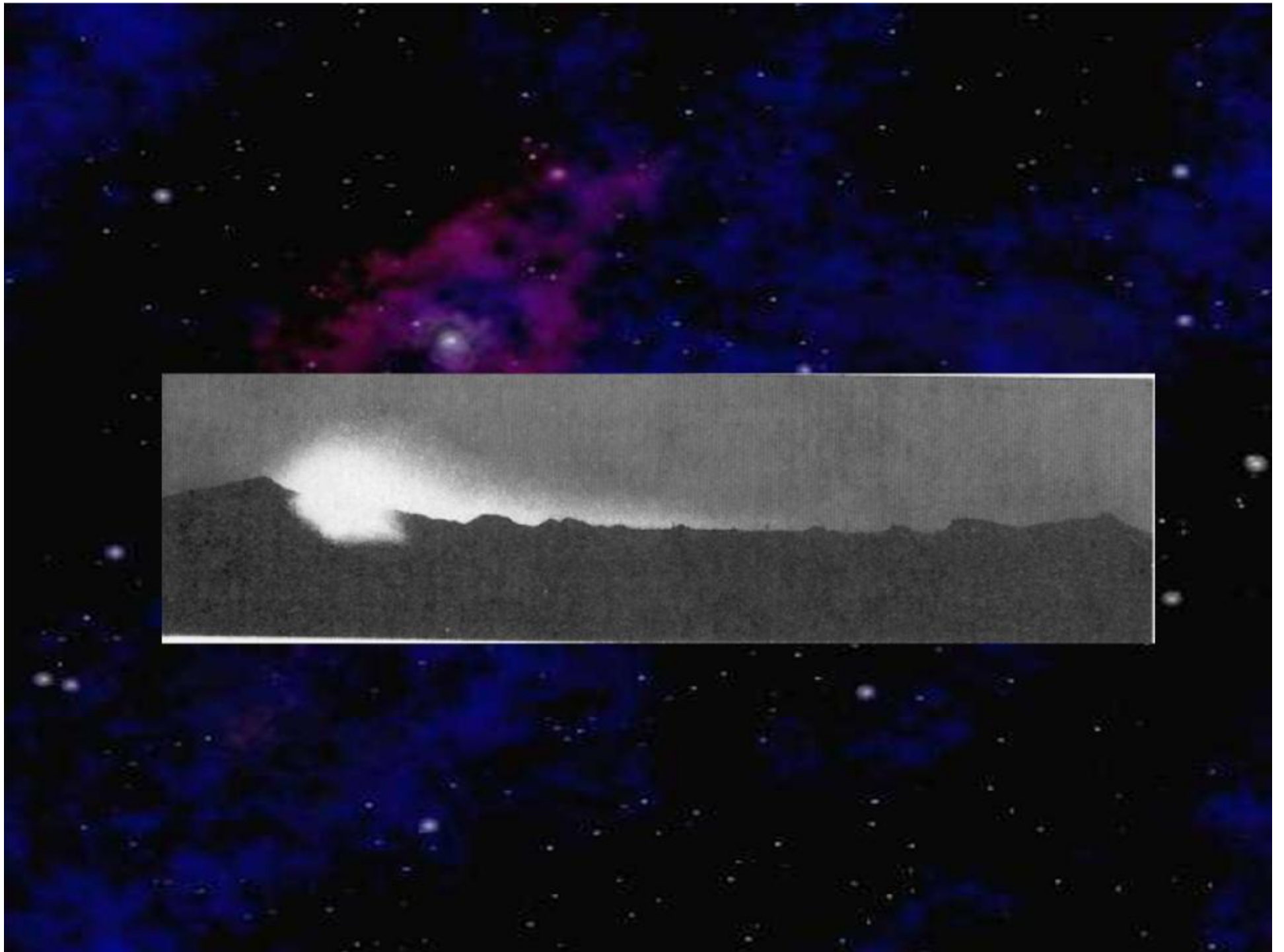
Why does the Earth experience
Seasons?

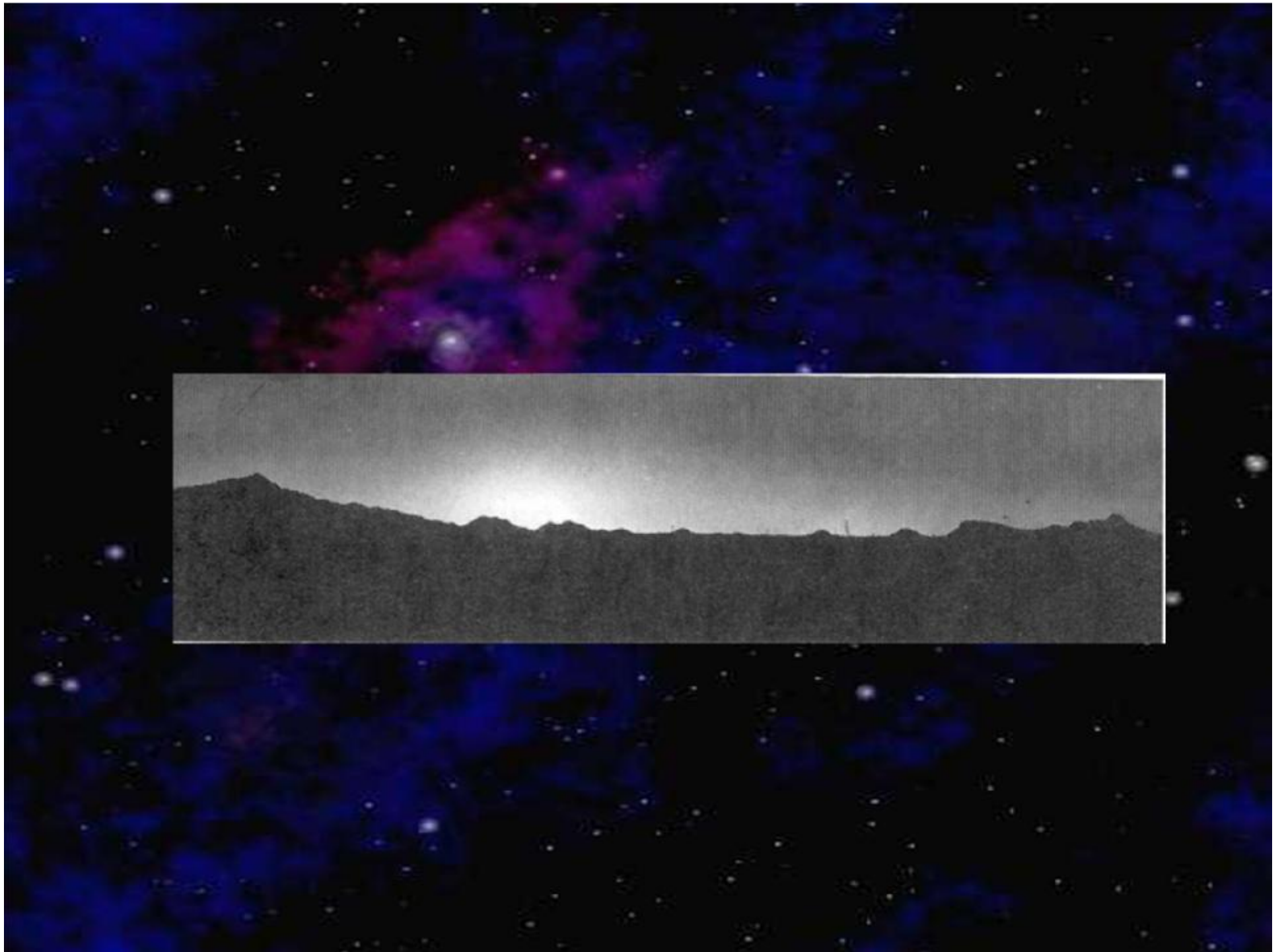
1. The Sun rises and sets at different places along the horizon.

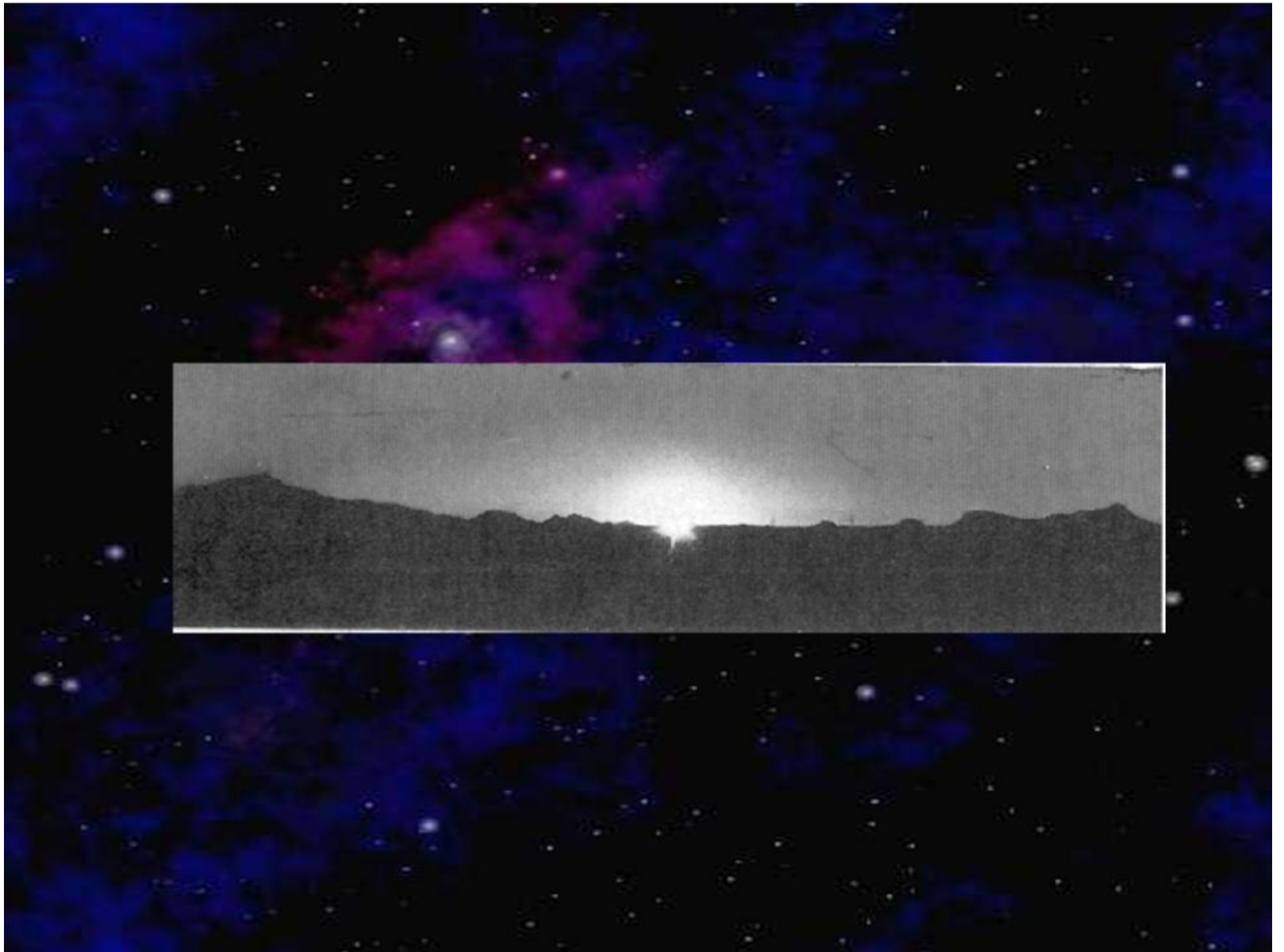


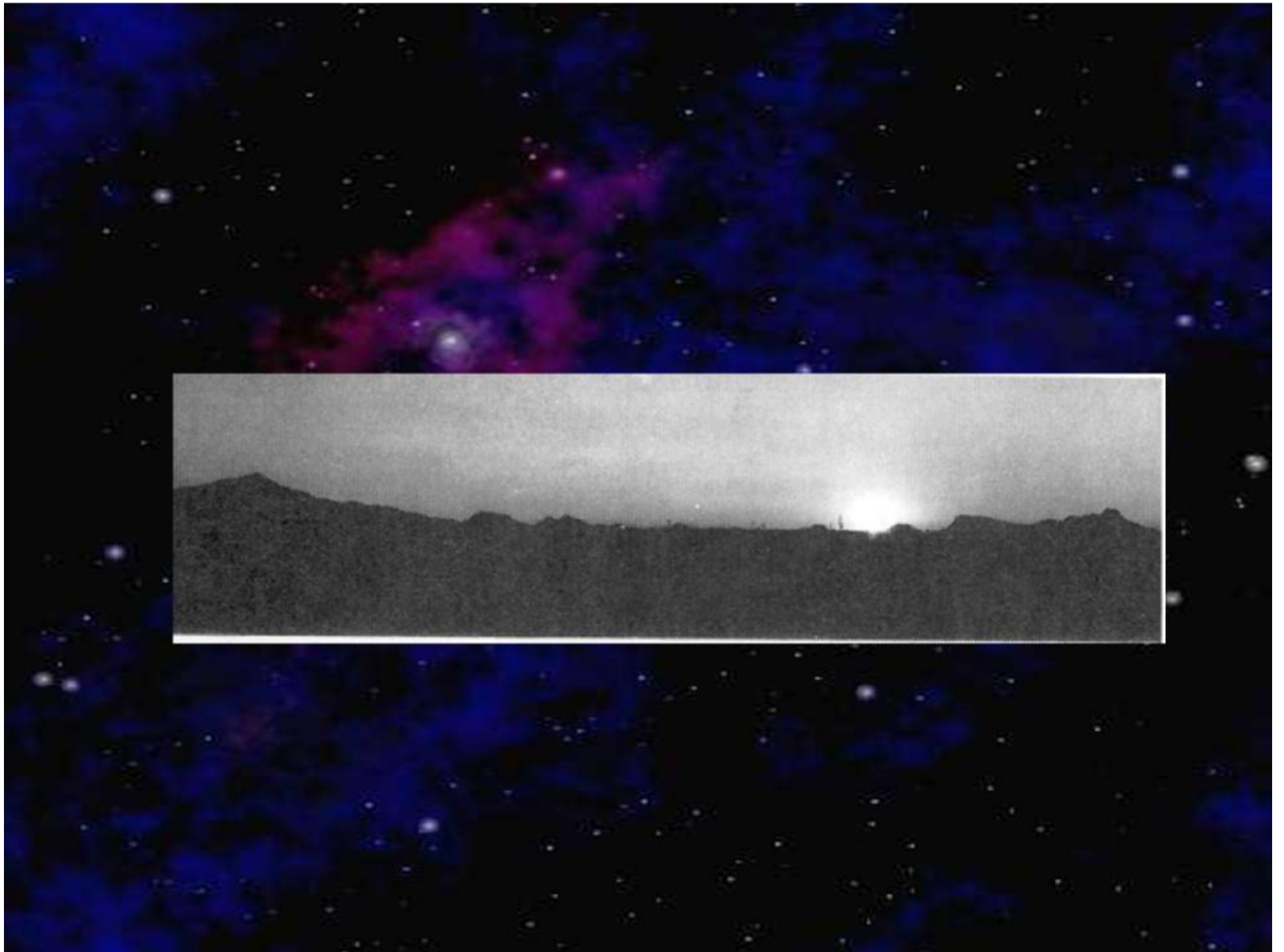


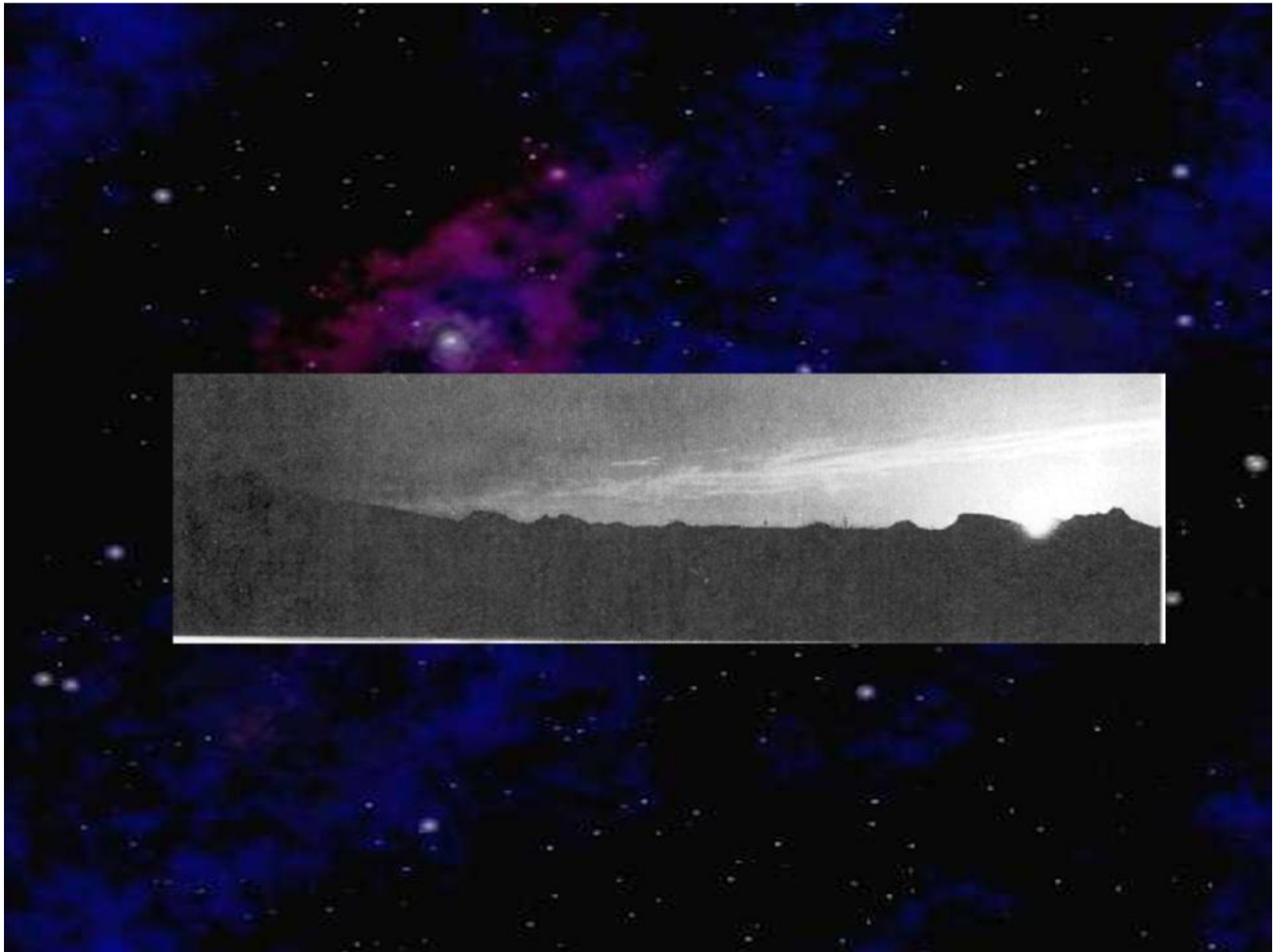


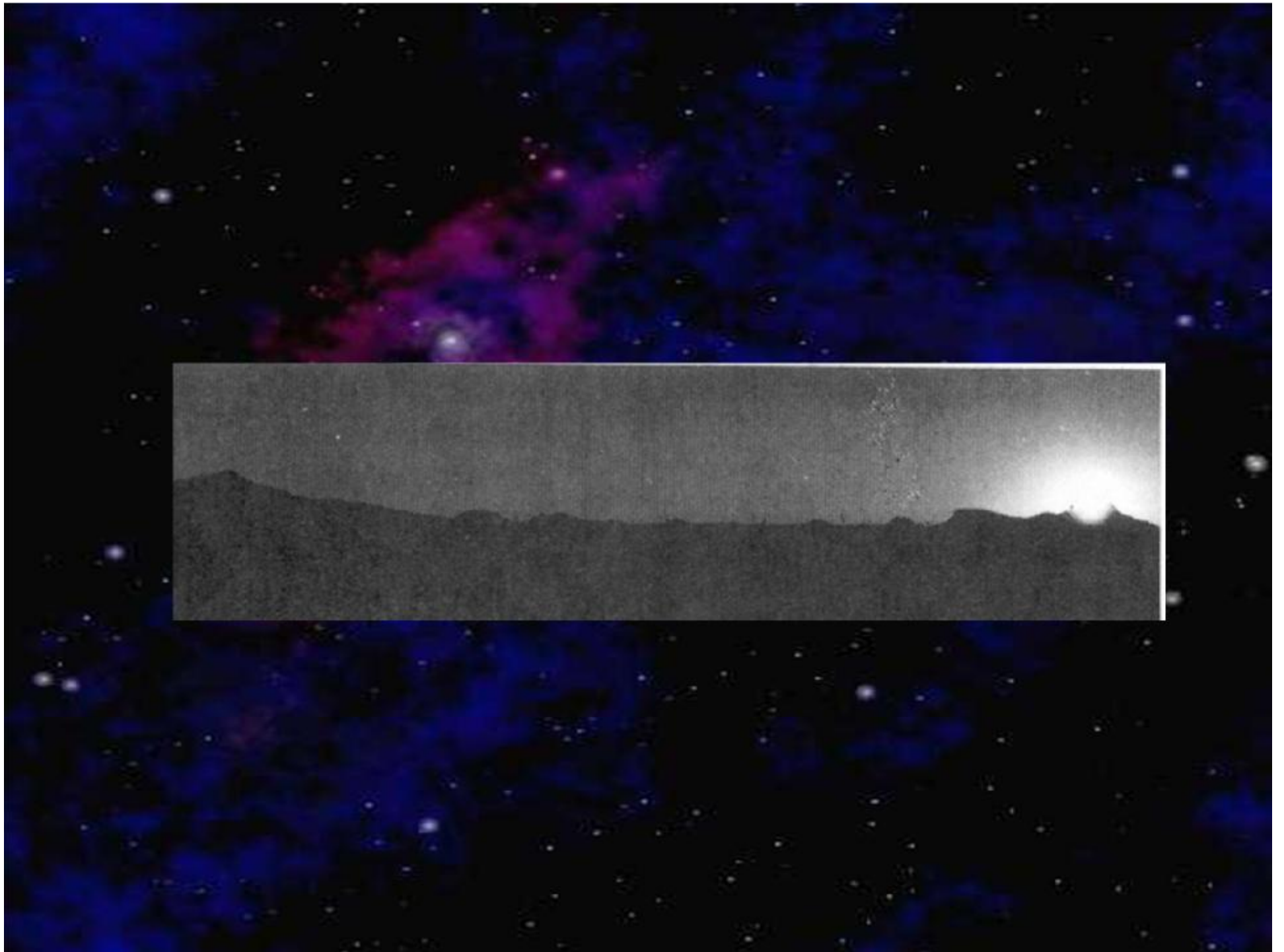




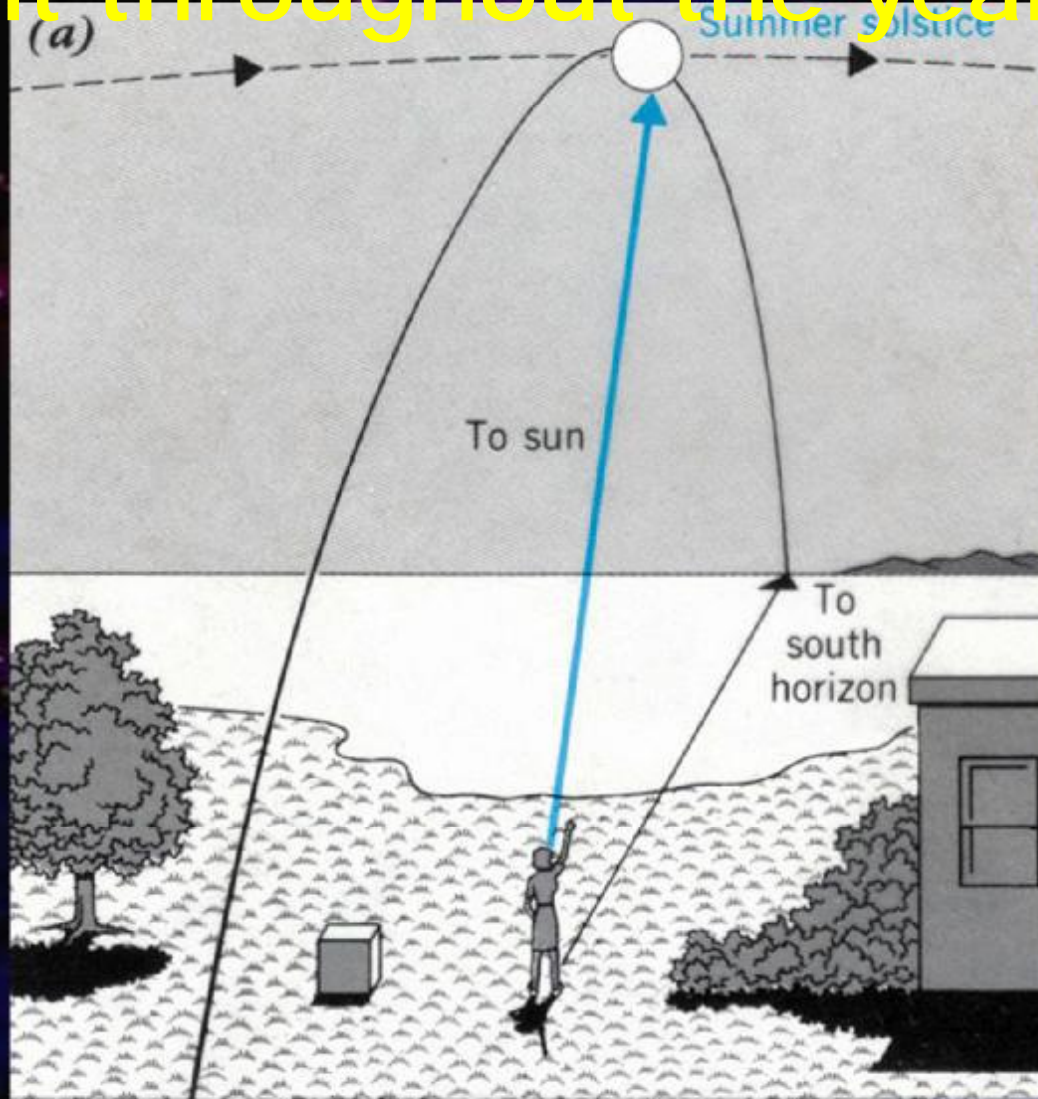


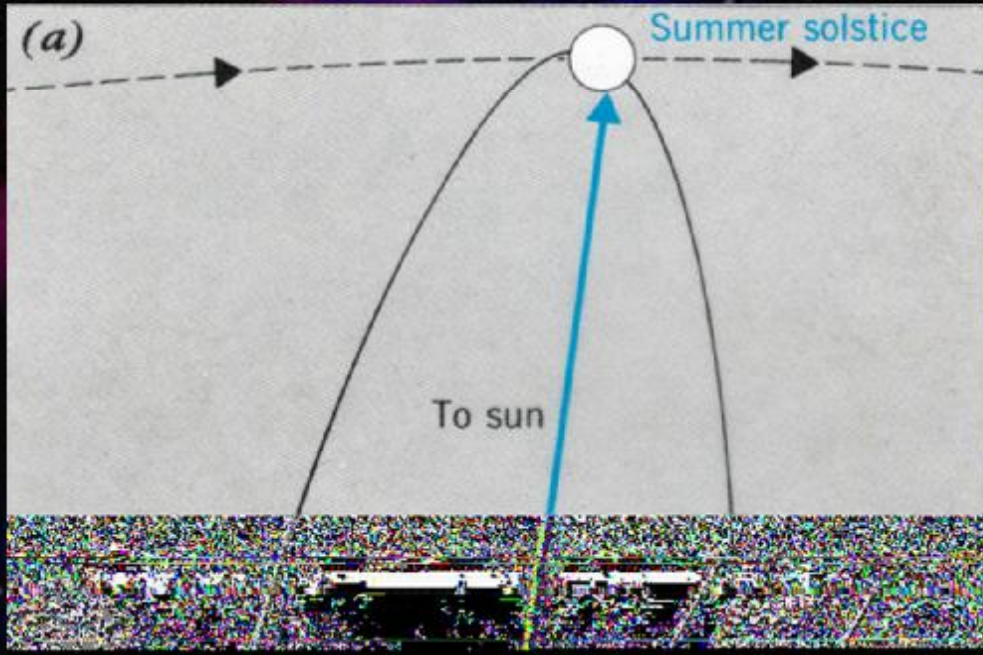


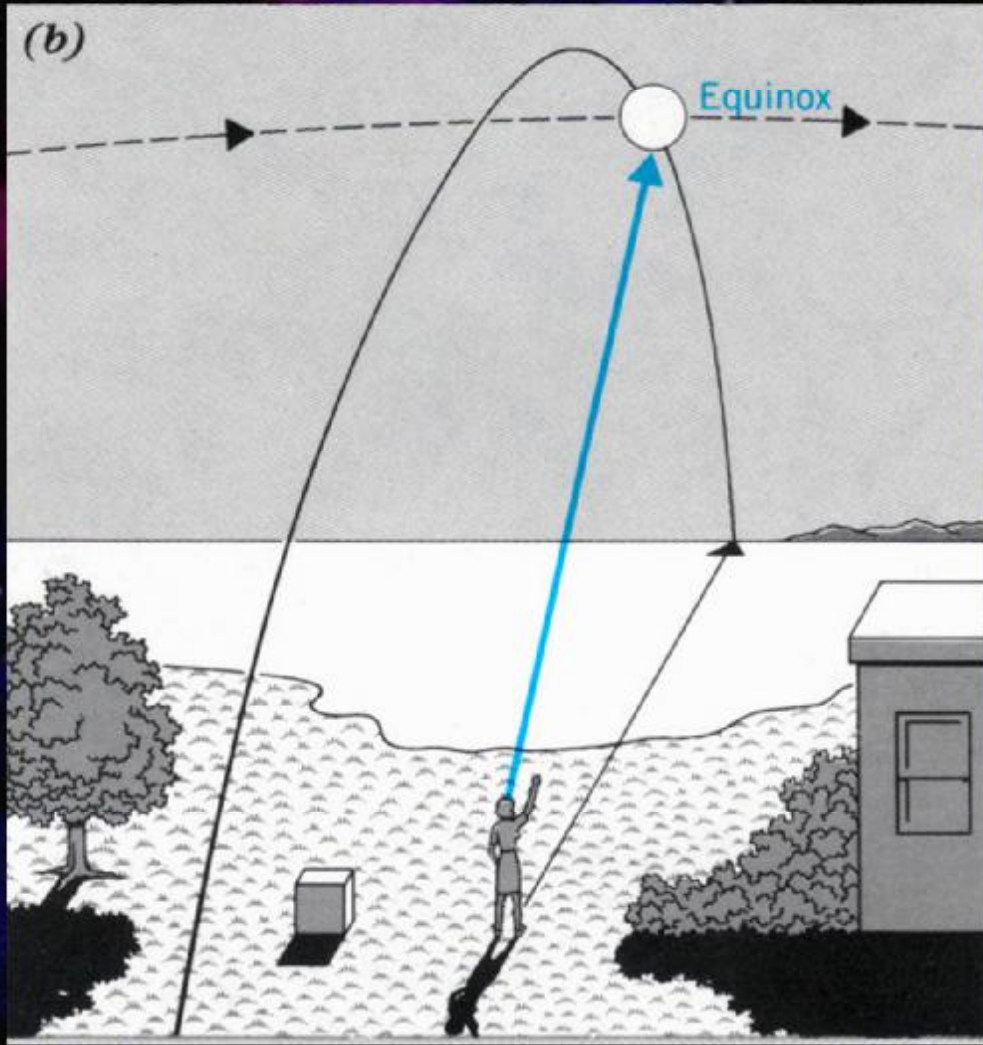


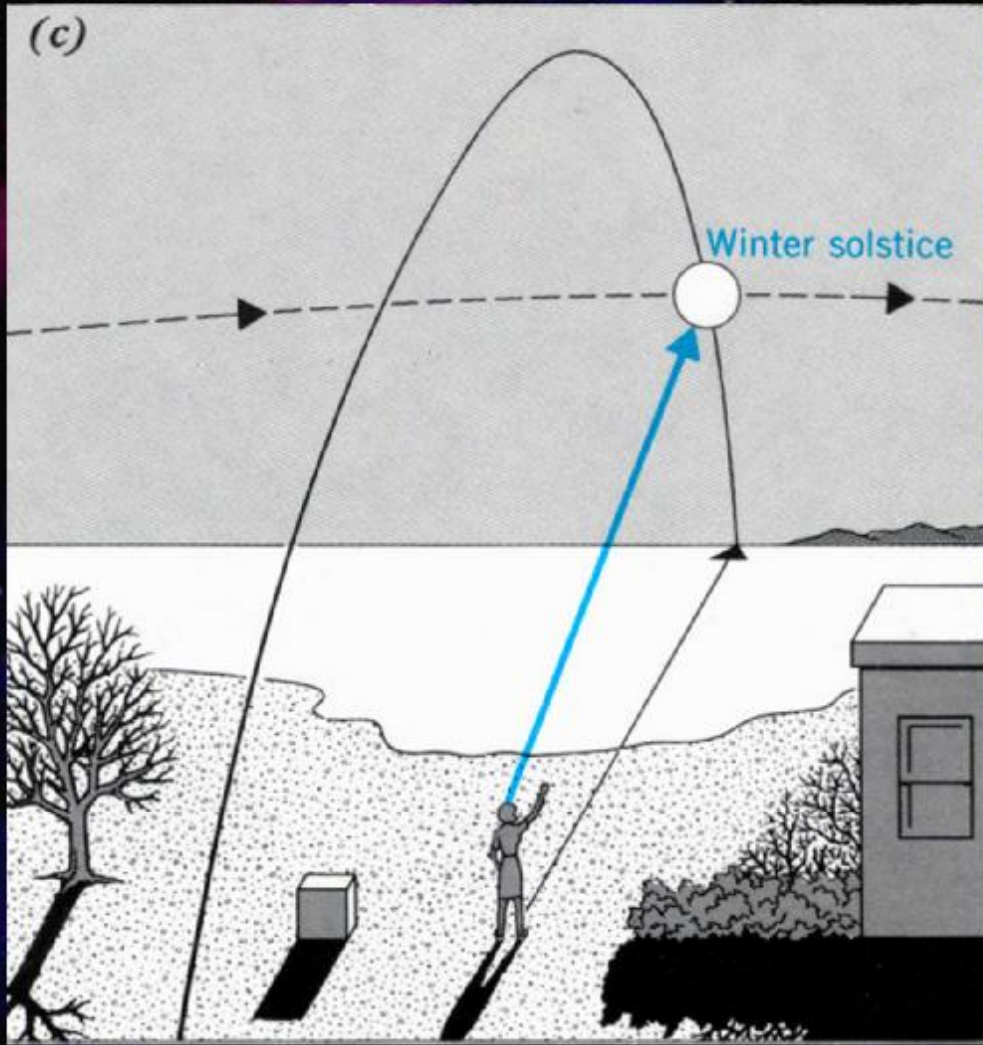


2. The height of the Sun at noon is different throughout the year.

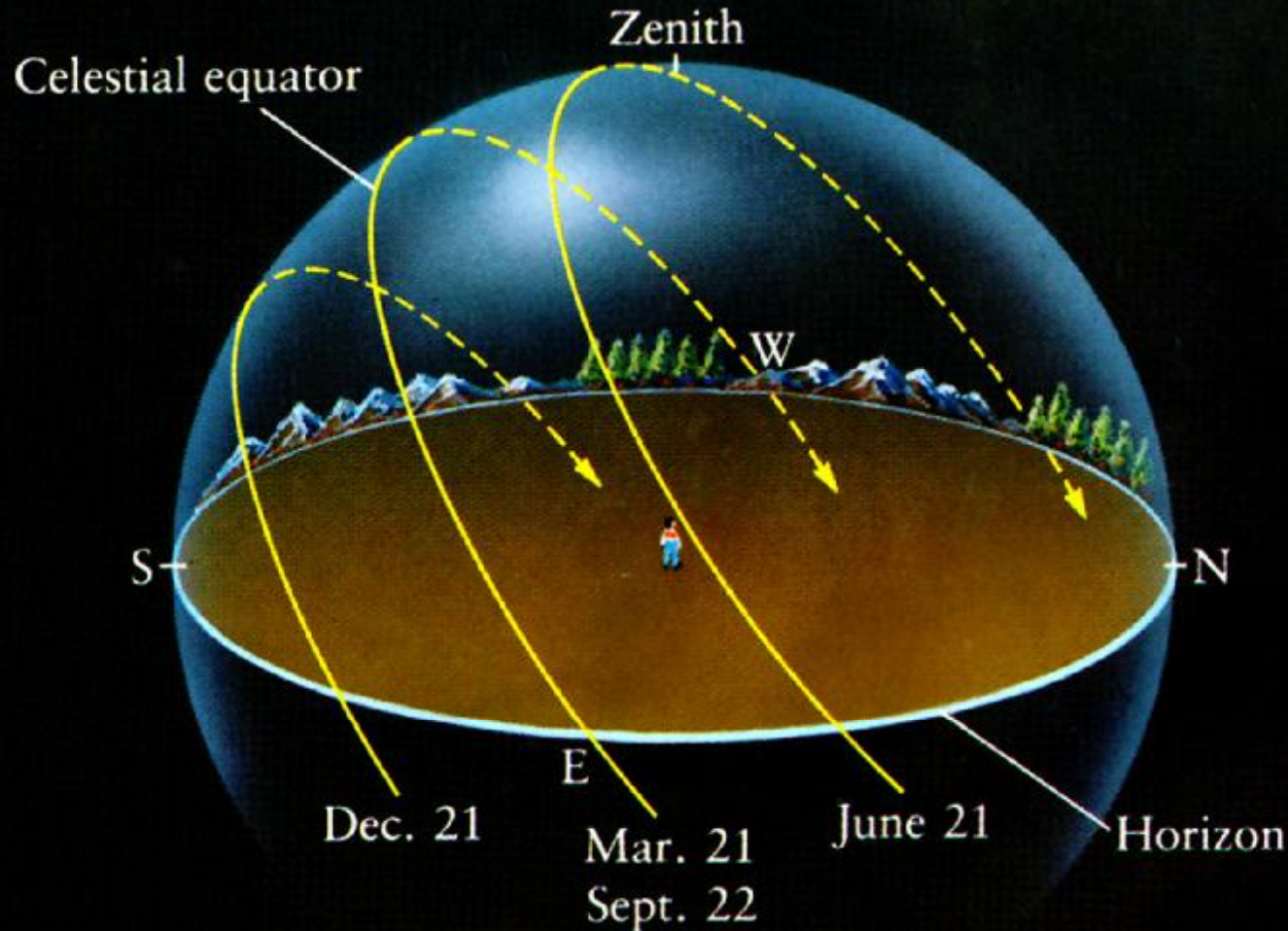




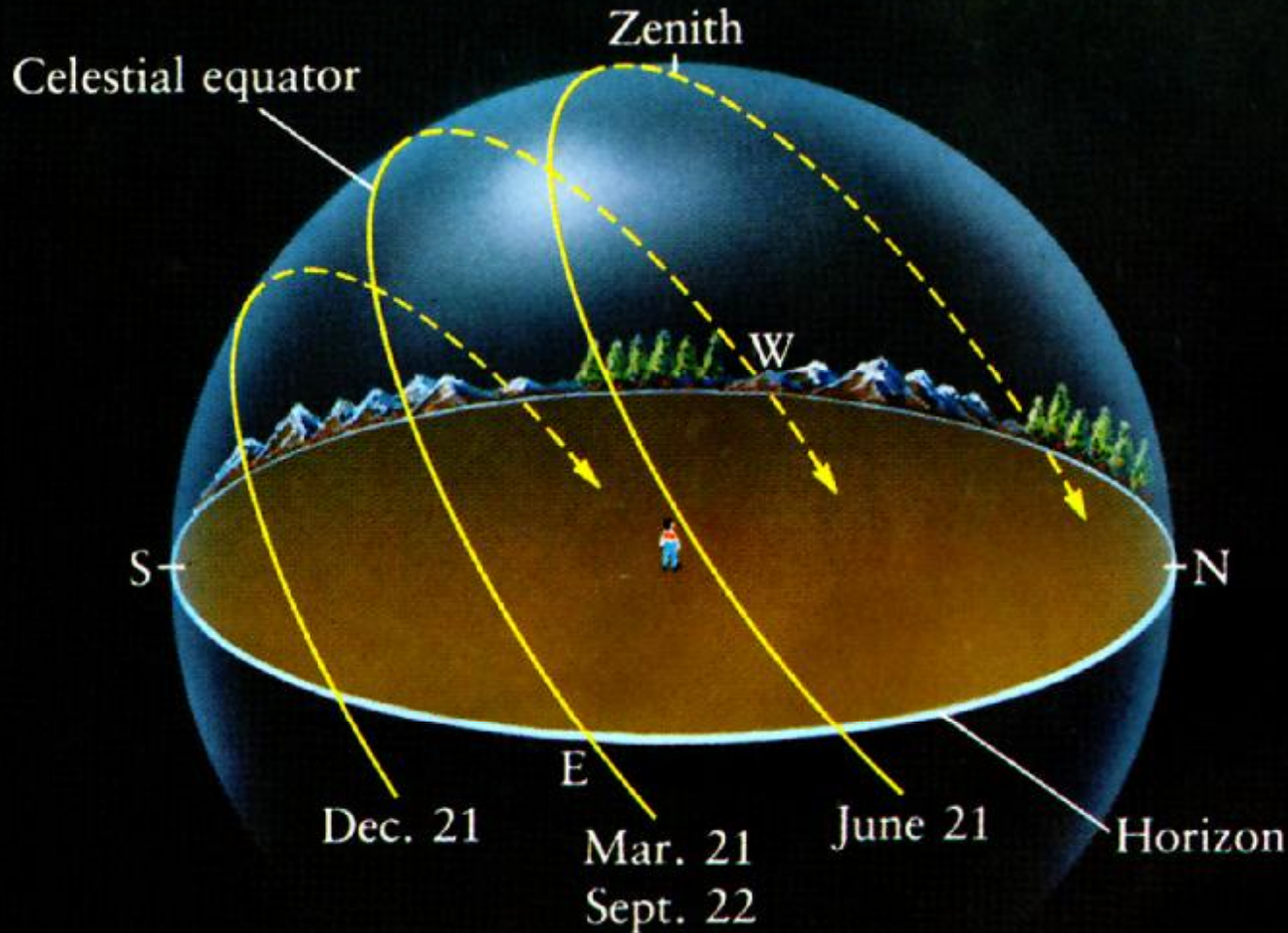


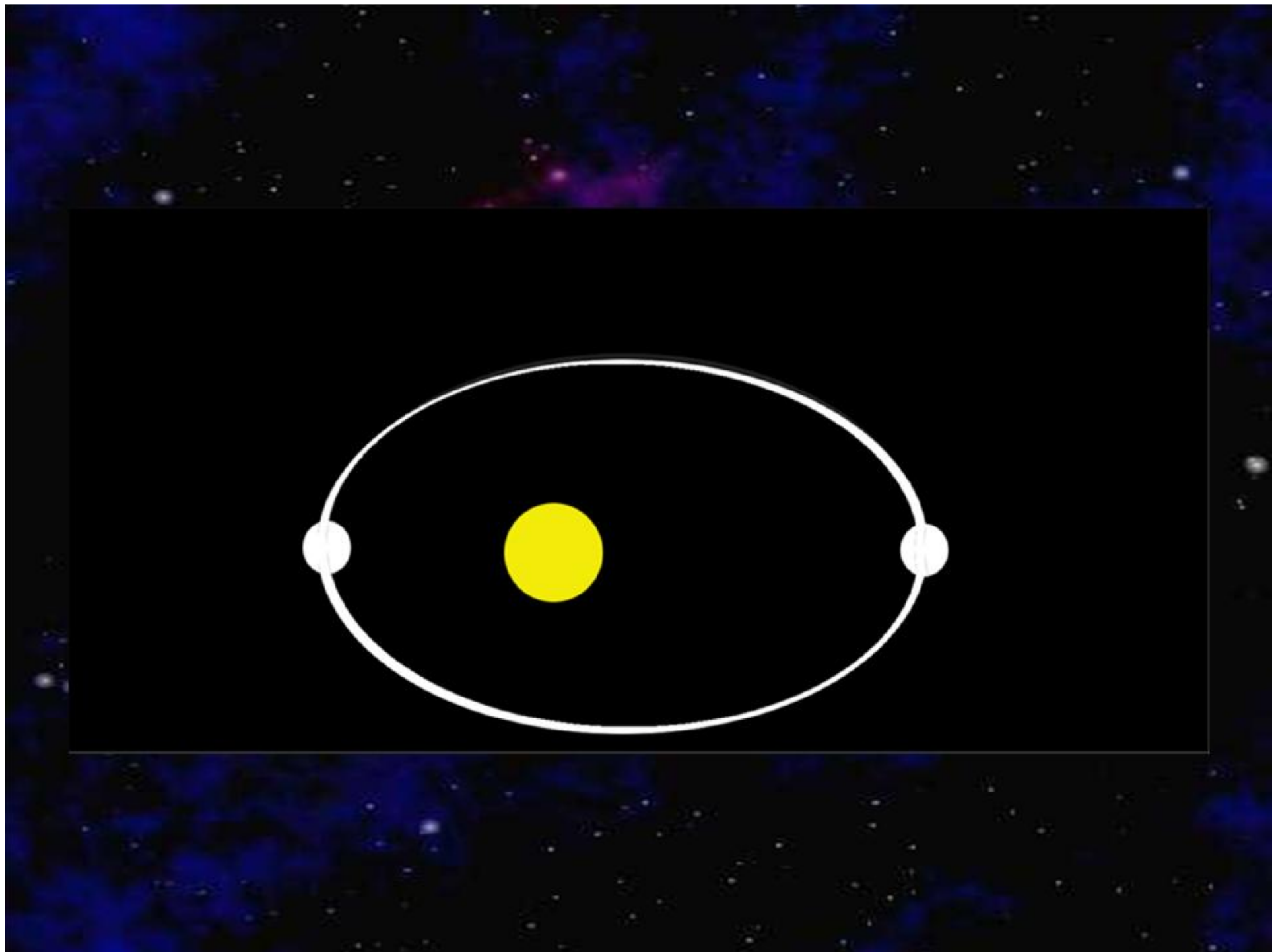


3. The size of the Sun's path across the sky is different throughout the year.

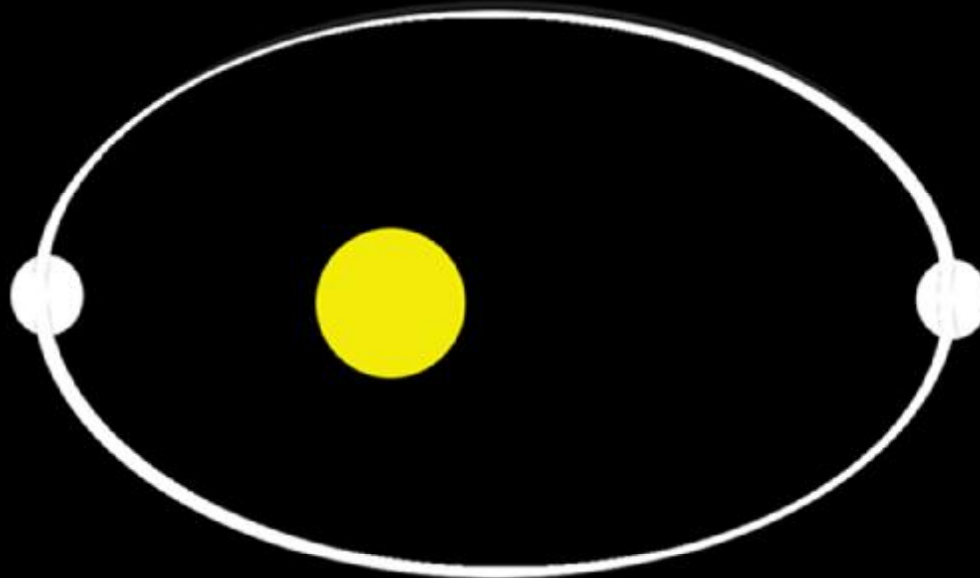


4. The length of daylight compared to the length of night changes each day.



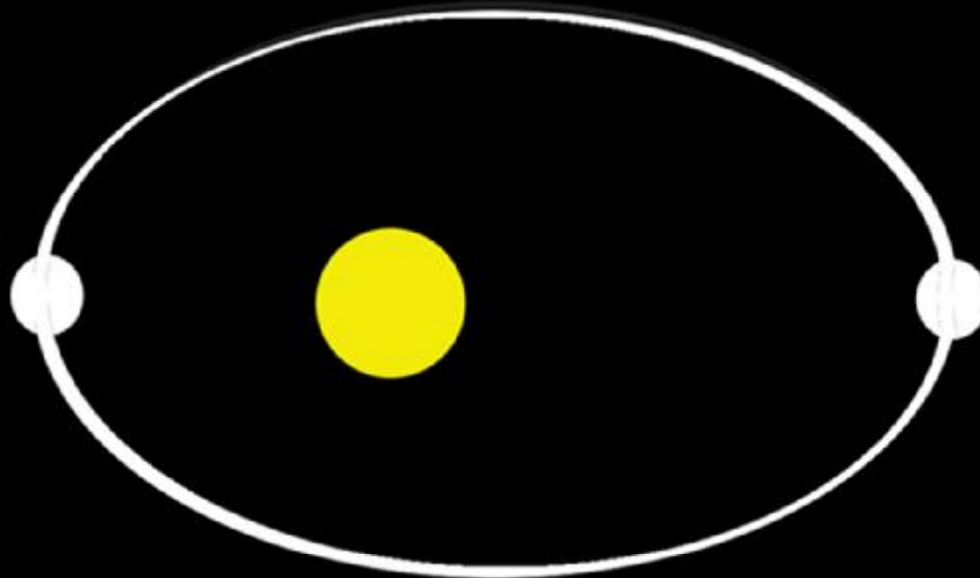


**Earth-Sun Distance
(Astronomical Unit) = 93 million miles**



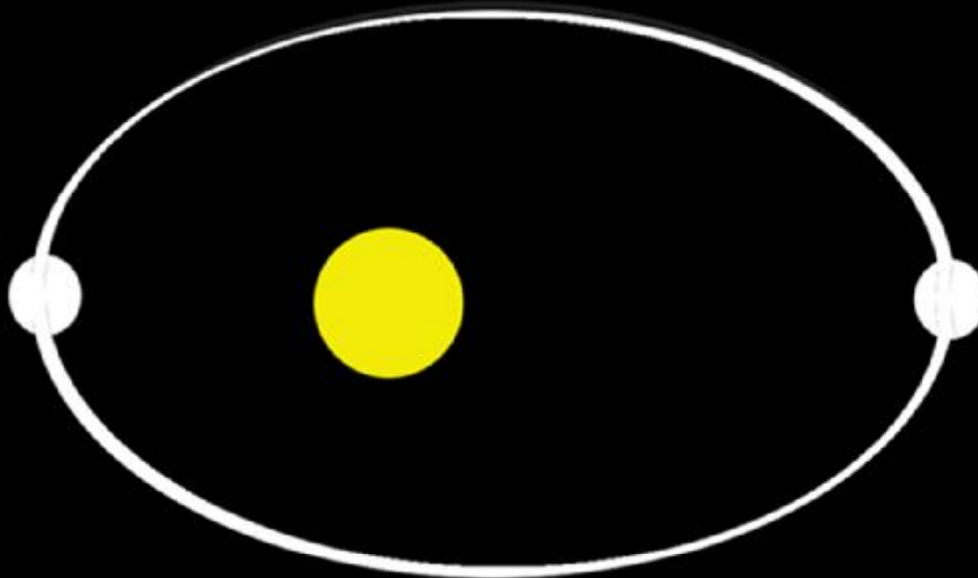
**Earth-Sun Distance
(Astronomical Unit) = 93 million miles**

Perihelion



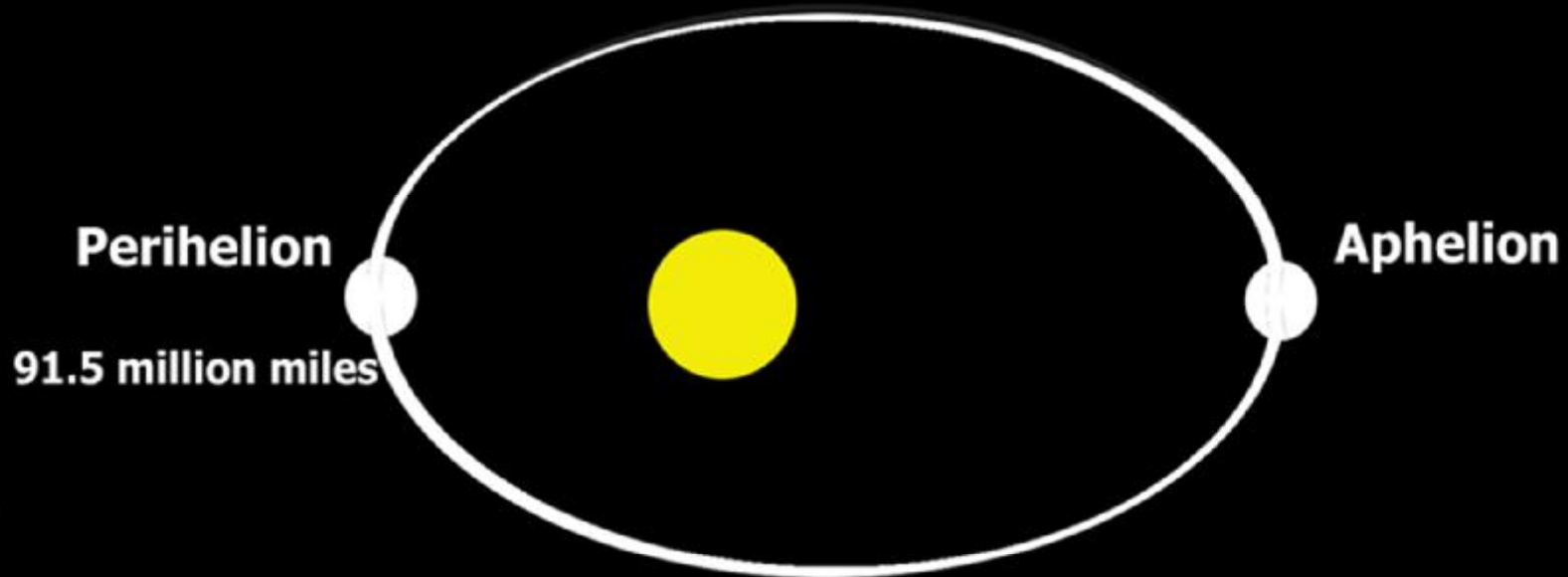
**Earth-Sun Distance
(Astronomical Unit) = 93 million miles**

Perihelion

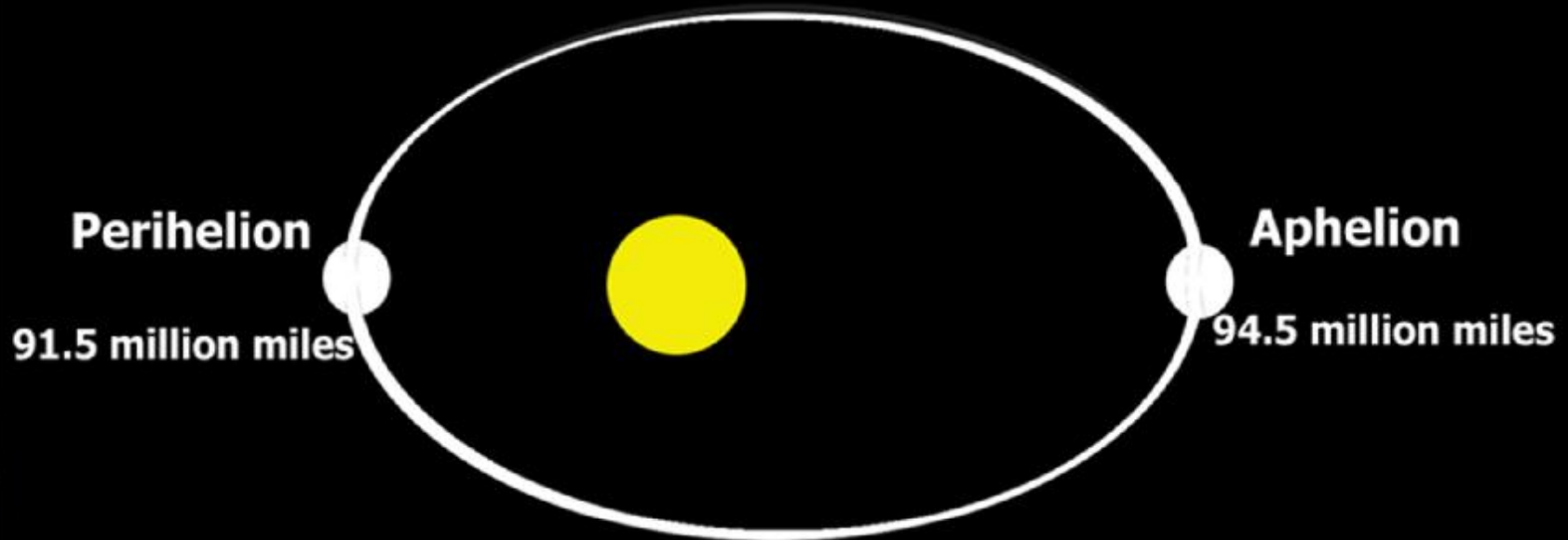


Aphelion

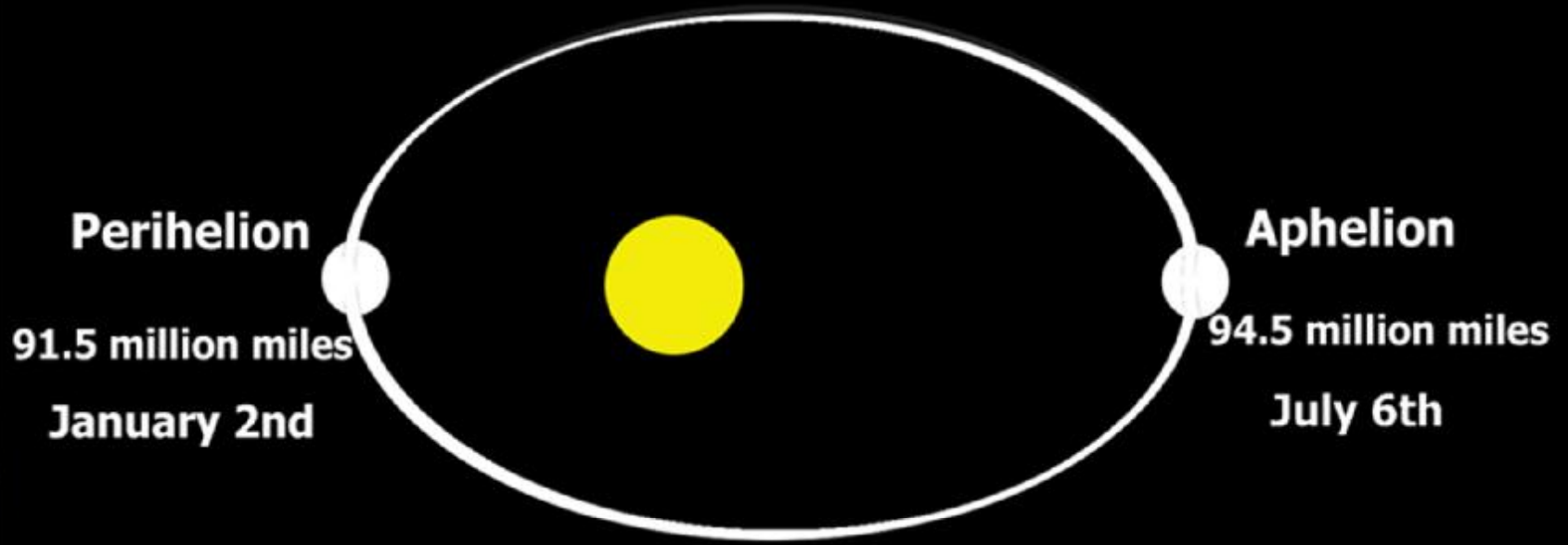
**Earth-Sun Distance
(Astronomical Unit) = 93 million miles**



**Earth-Sun Distance
(Astronomical Unit) = 93 million miles**



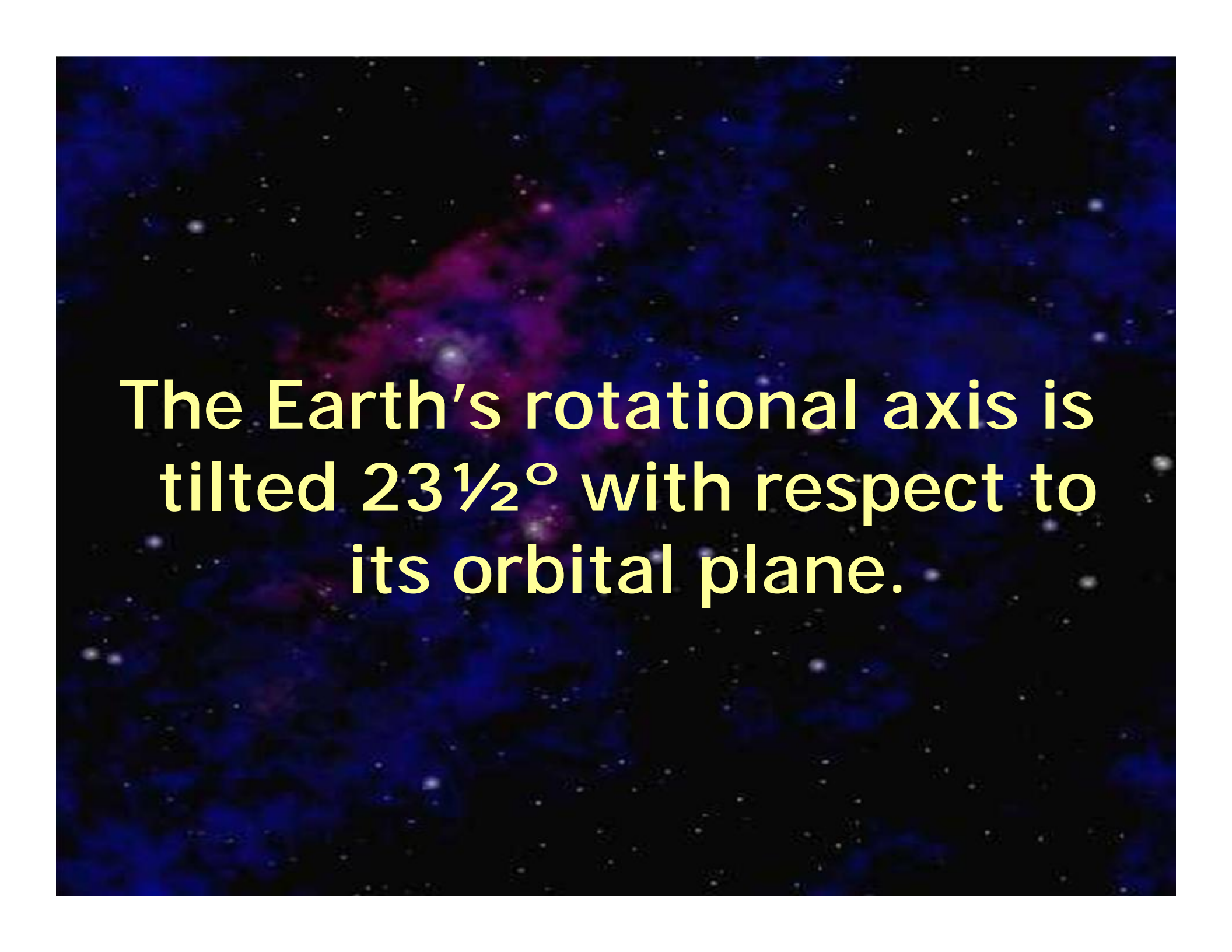
**Earth-Sun Distance
(Astronomical Unit) = 93 million miles**





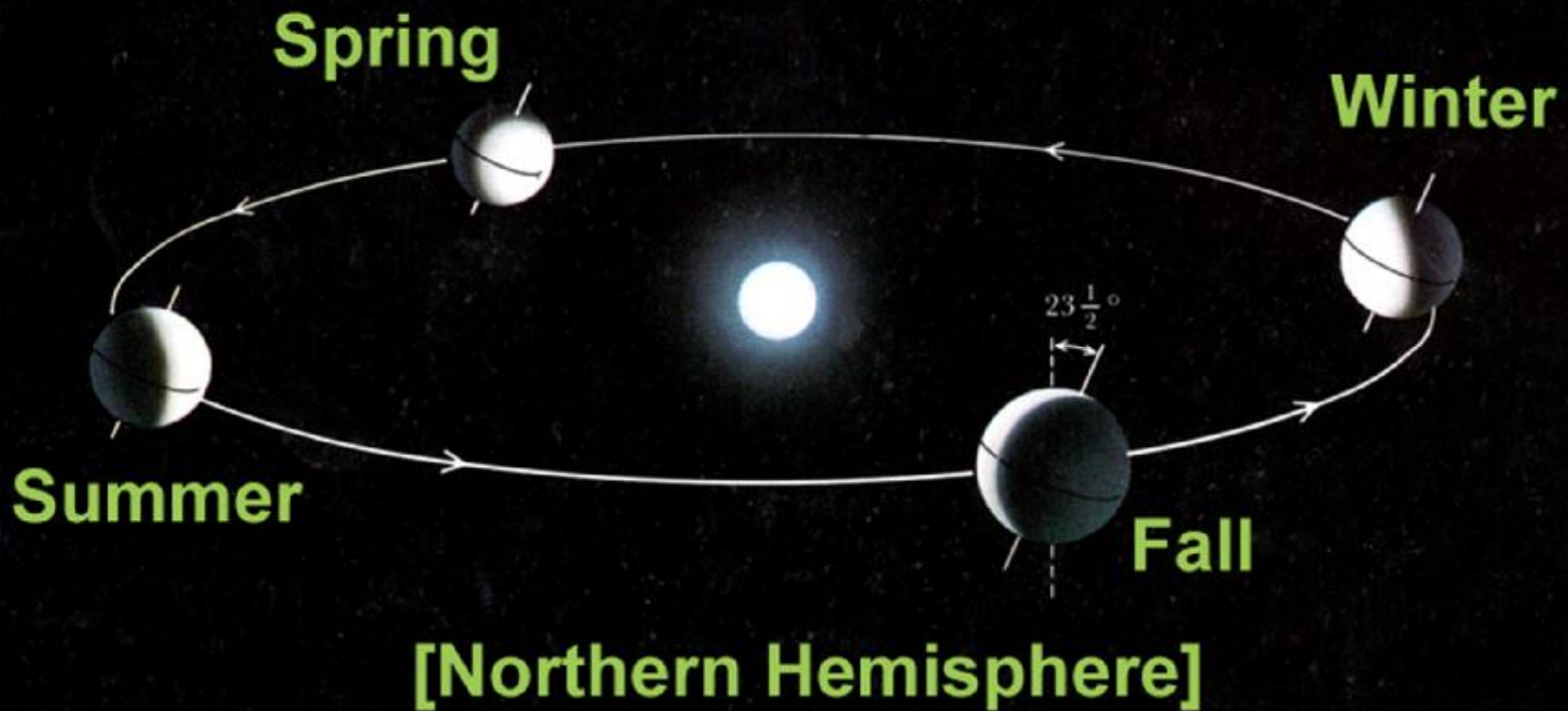
A diagram illustrating the orbit of a celestial body around the Sun. The Sun is represented by a yellow dot at the center of a circular orbit. The center of the orbit is labeled "Center". The Sun is labeled "Sun". The point on the orbit furthest from the Sun is labeled "Aphelion". The point on the orbit closest to the Sun is labeled "Perihelion". A horizontal line passes through the Sun and the center, with white dots at the Aphelion and Perihelion points.

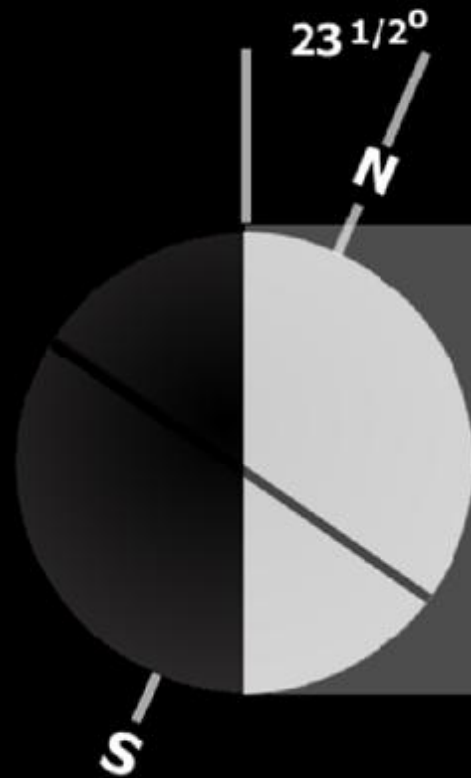
Aphelion **Center** **Sun** **Perihelion**



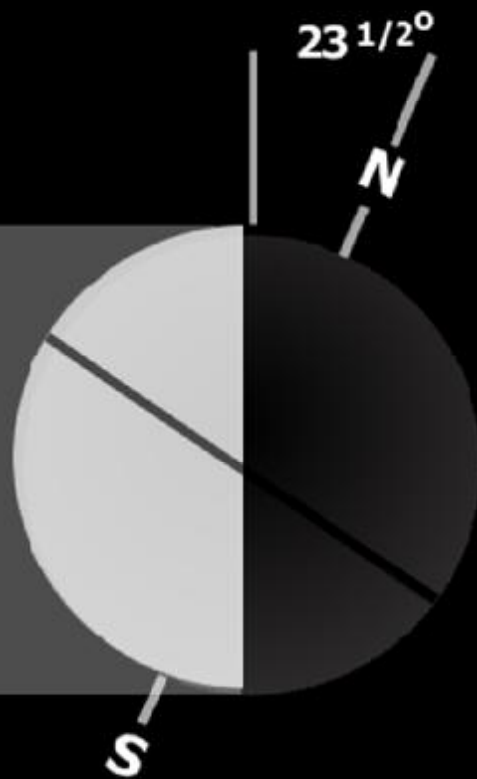
The Earth's rotational axis is tilted $23\frac{1}{2}^{\circ}$ with respect to its orbital plane.

Earth's Orbit

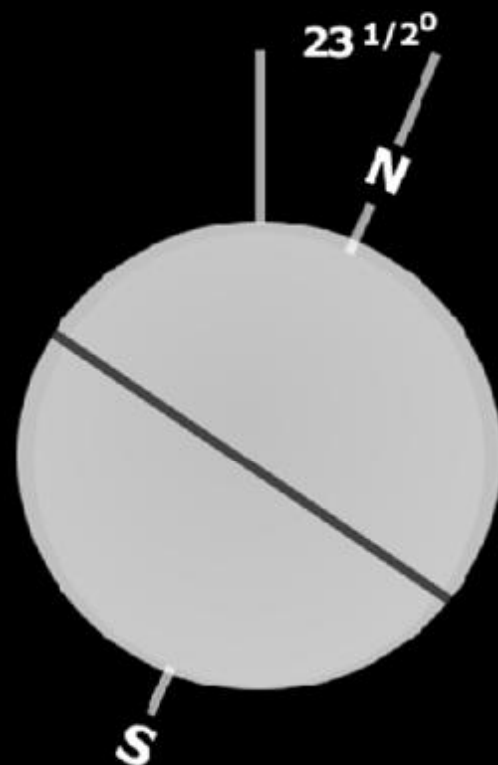




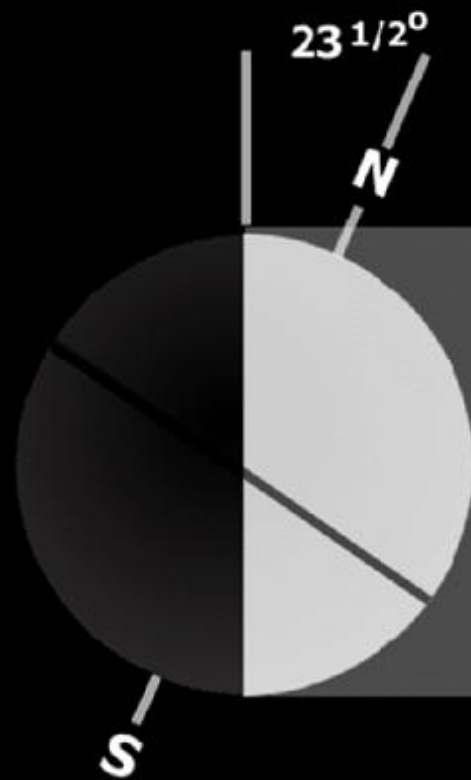
Summer



Winter

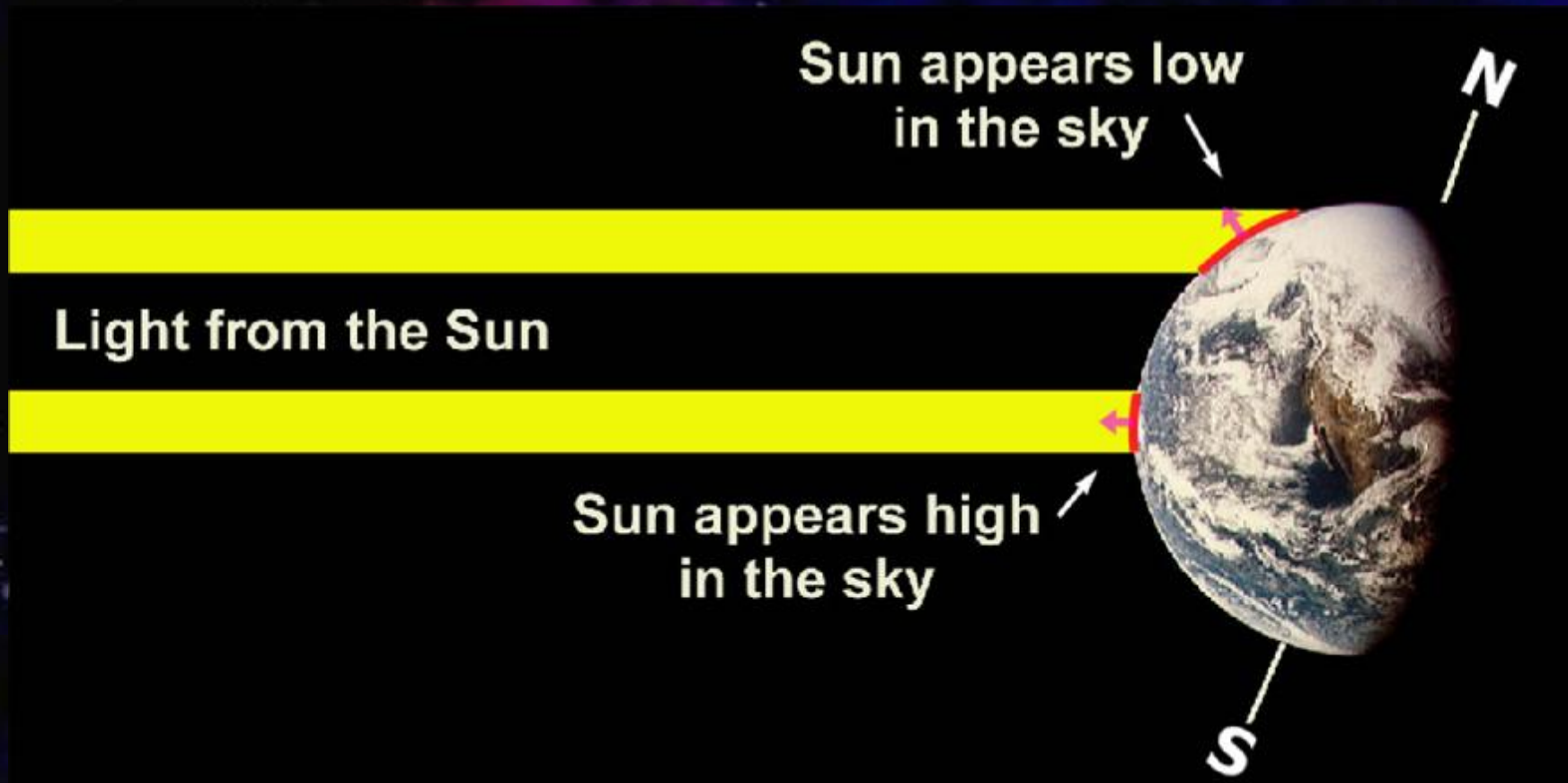


Spring or Fall



Summer

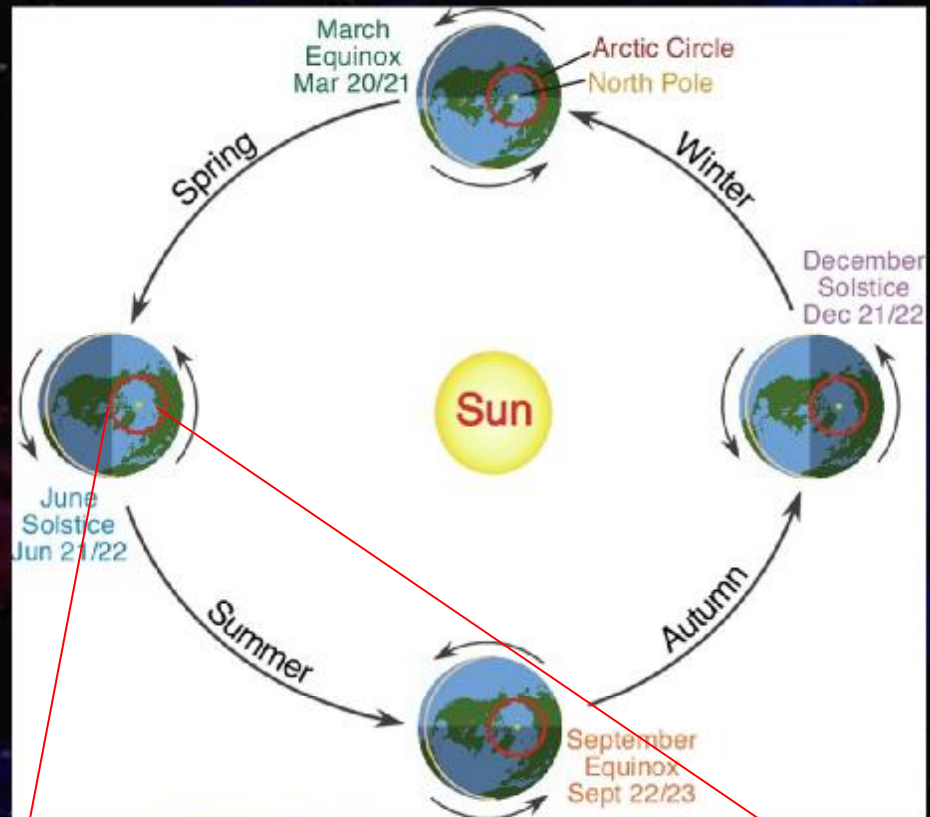
The intensity of sunlight striking the Earth varies with location:



Where on Earth is this photo taken?

What was the calendar date?

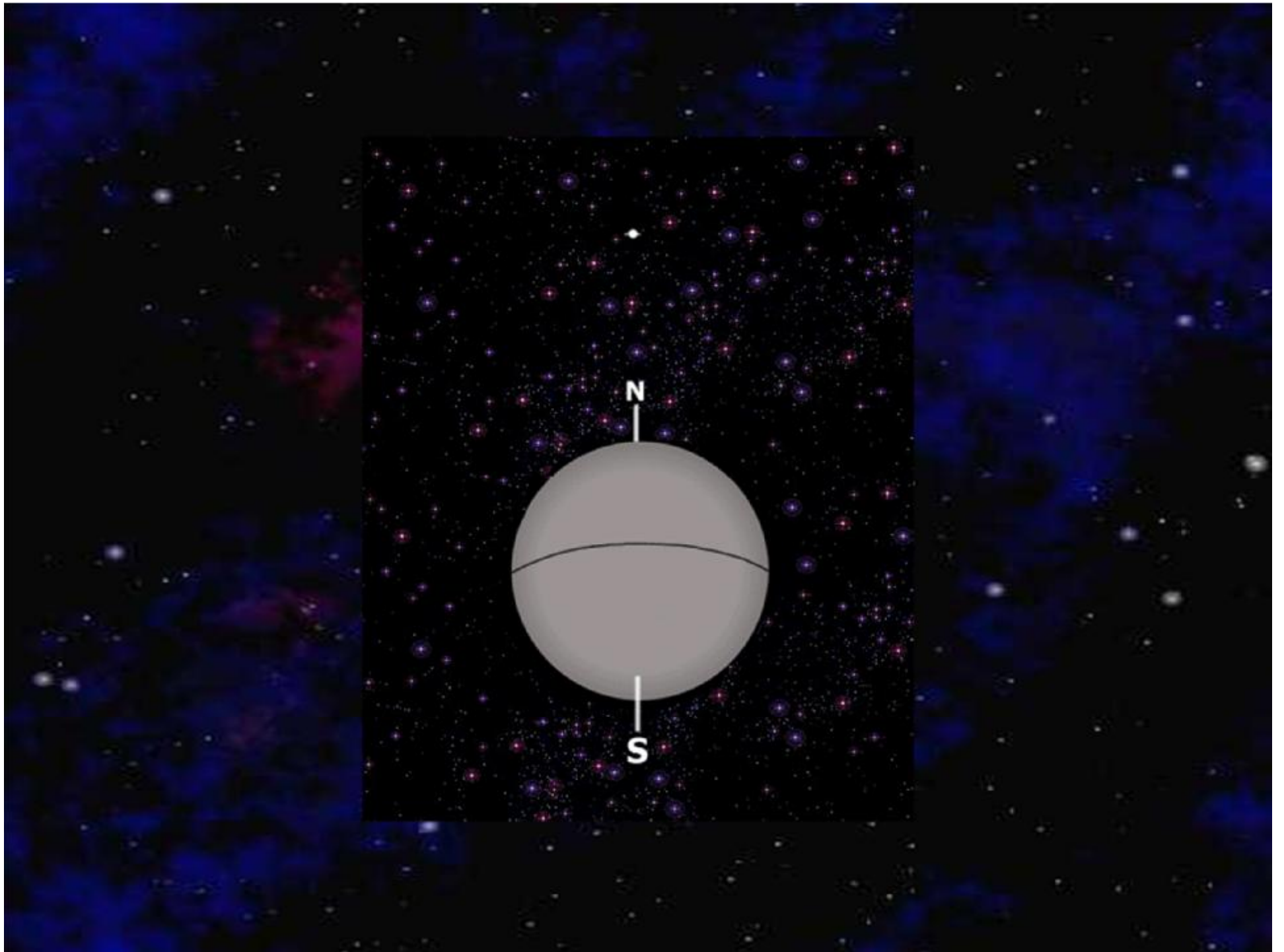


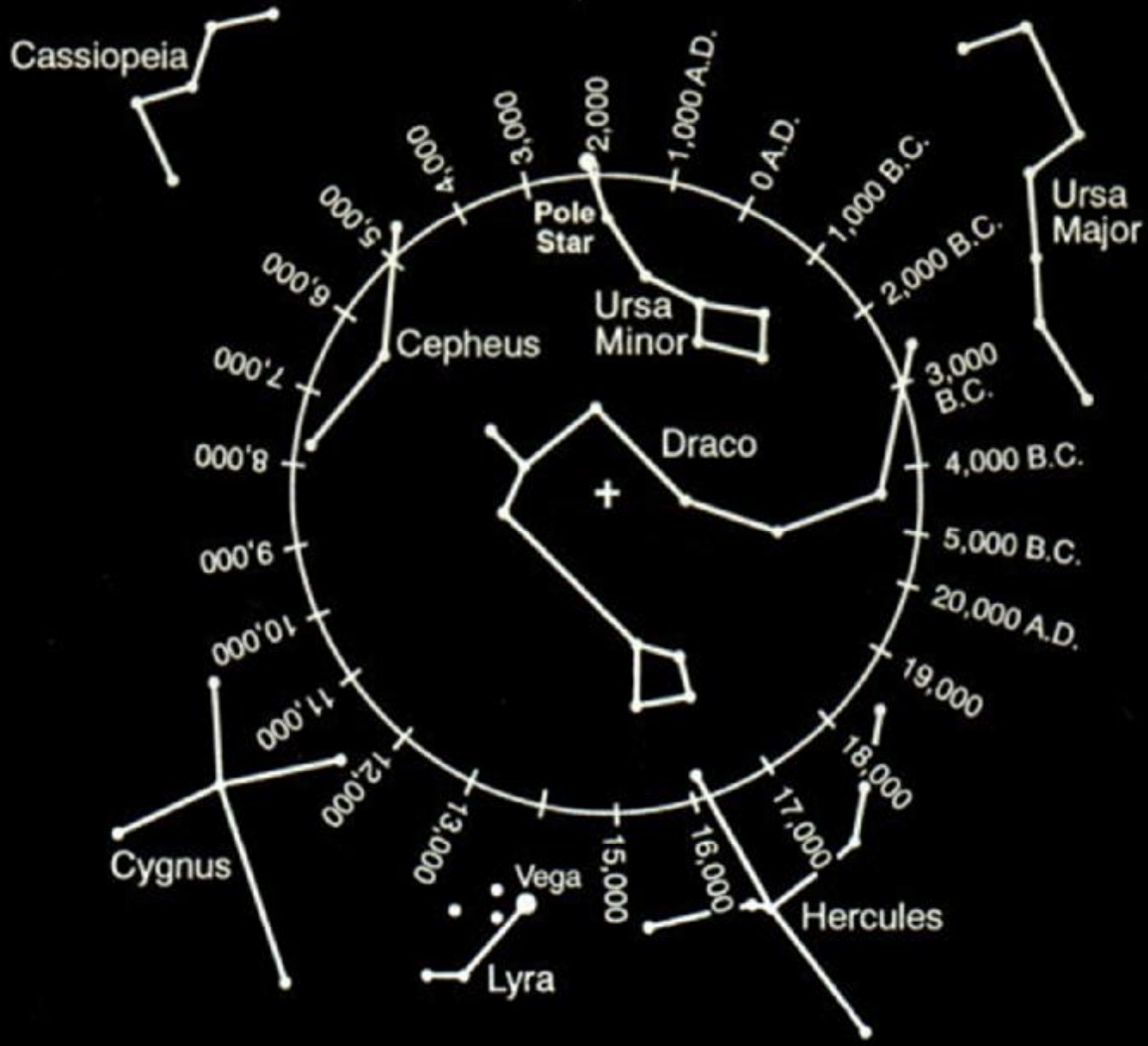


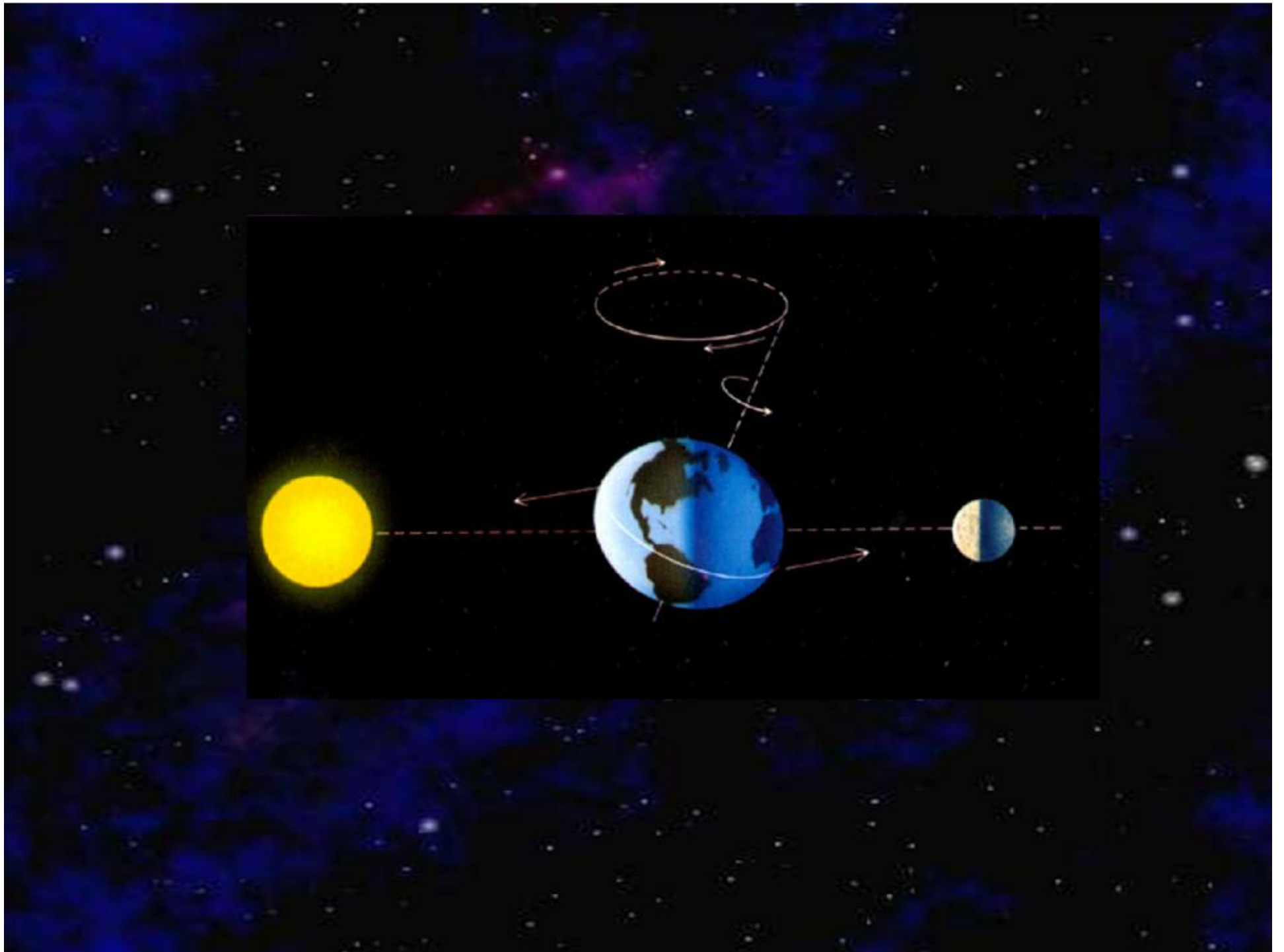
A deep space photograph showing a vast field of stars against a dark blue background. In the center, there is a prominent nebula with a mix of purple and pink hues. The text "LONG TERM CHANGES" is overlaid in yellow, bold, sans-serif font.

LONG TERM CHANGES

PRECESSION







PRECESSION

Slow change in the orientation of the Earth's axis of rotation.

§ Caused by the gravitational interaction between the Sun, Earth, and Moon.

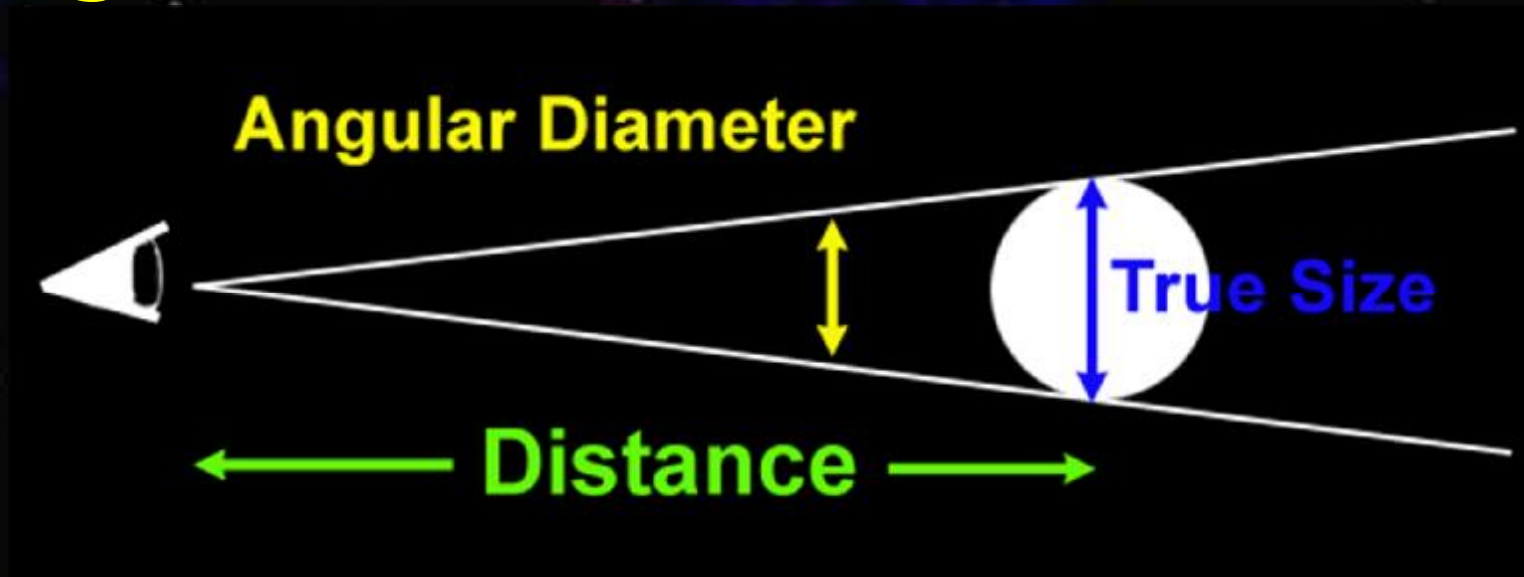
§ Tilt of the Earth remains $23\frac{1}{2}^{\circ}$ -BUT- changes orientation.

§ Period of precession ~ 26,000 years



ECLIPSES

Angular Diameter



Angular diameter of the Sun and Moon are approximately the same.

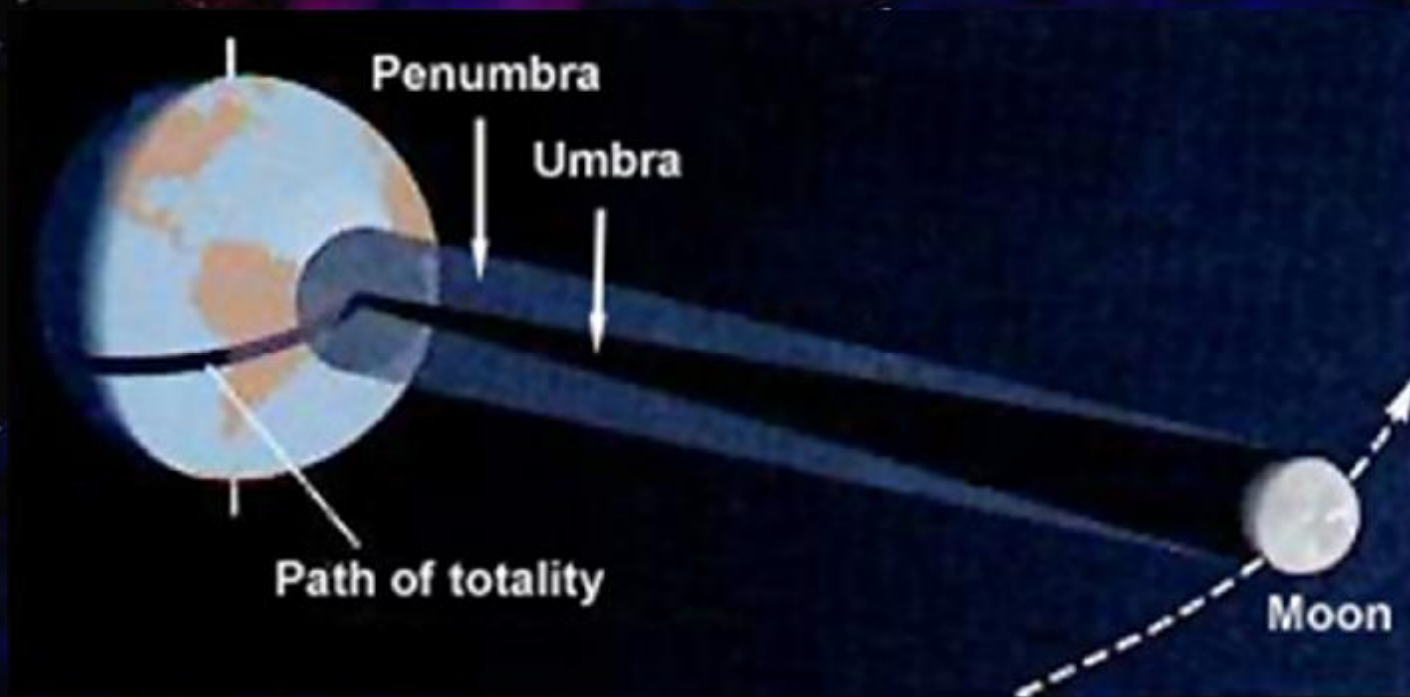
Both Appear to be $\sim 1/2^\circ$

Solar Eclipses

Occurs when the shadow of the Moon is cast upon the Earth.

Can only occur at NEW MOON

Anatomy of a Solar Eclipse

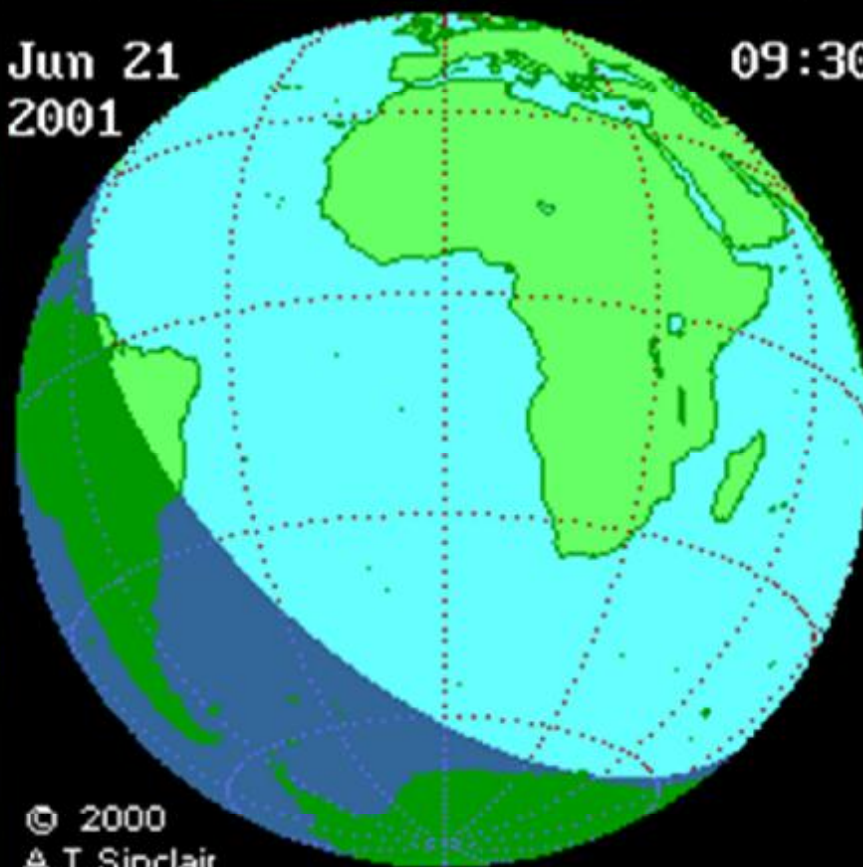


Solar Eclipse seen from space:



Jun 21
2001

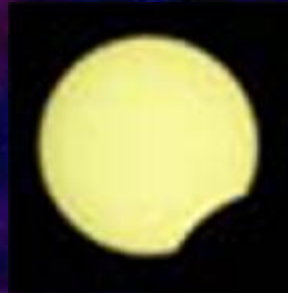
09:30

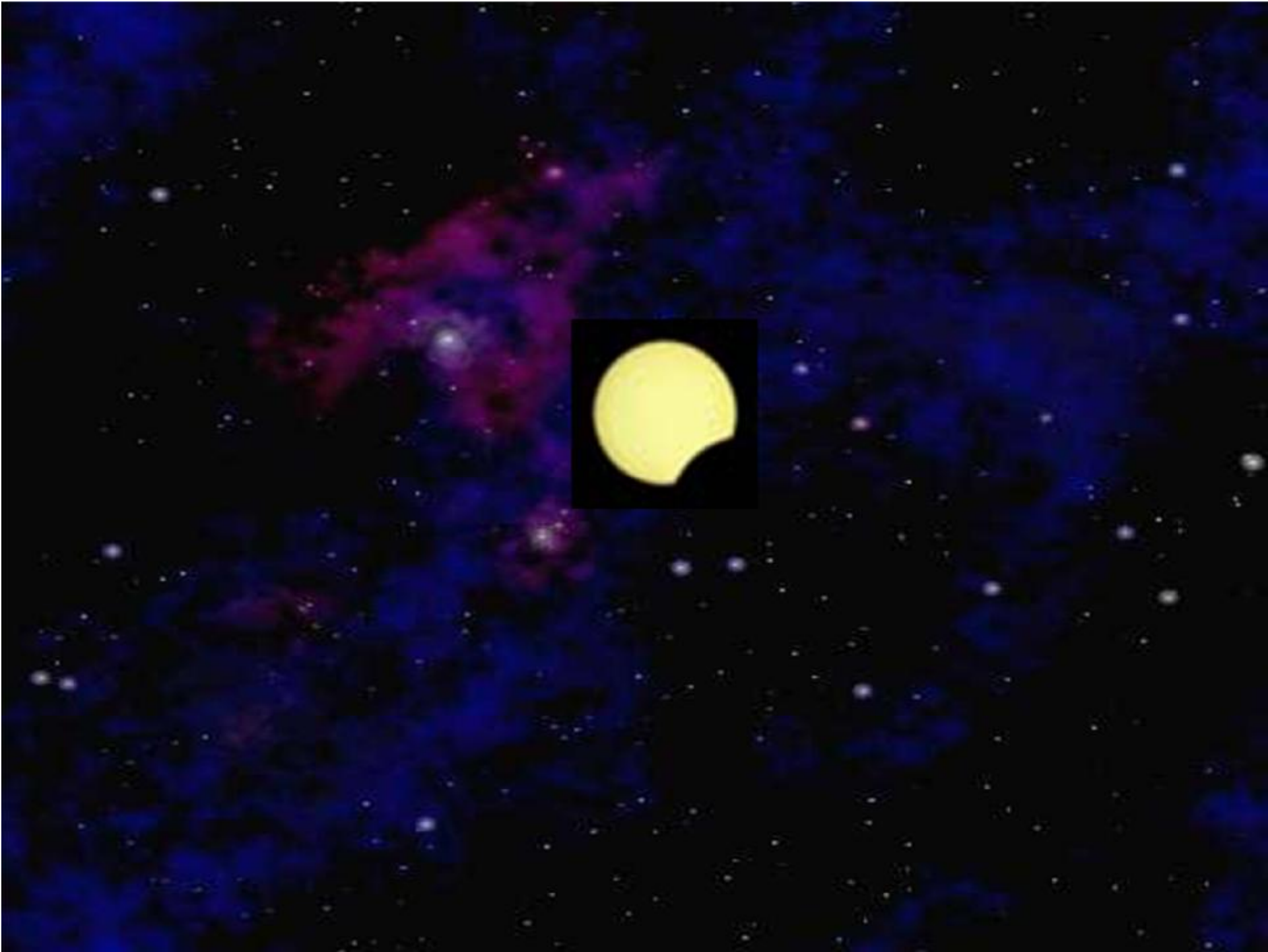


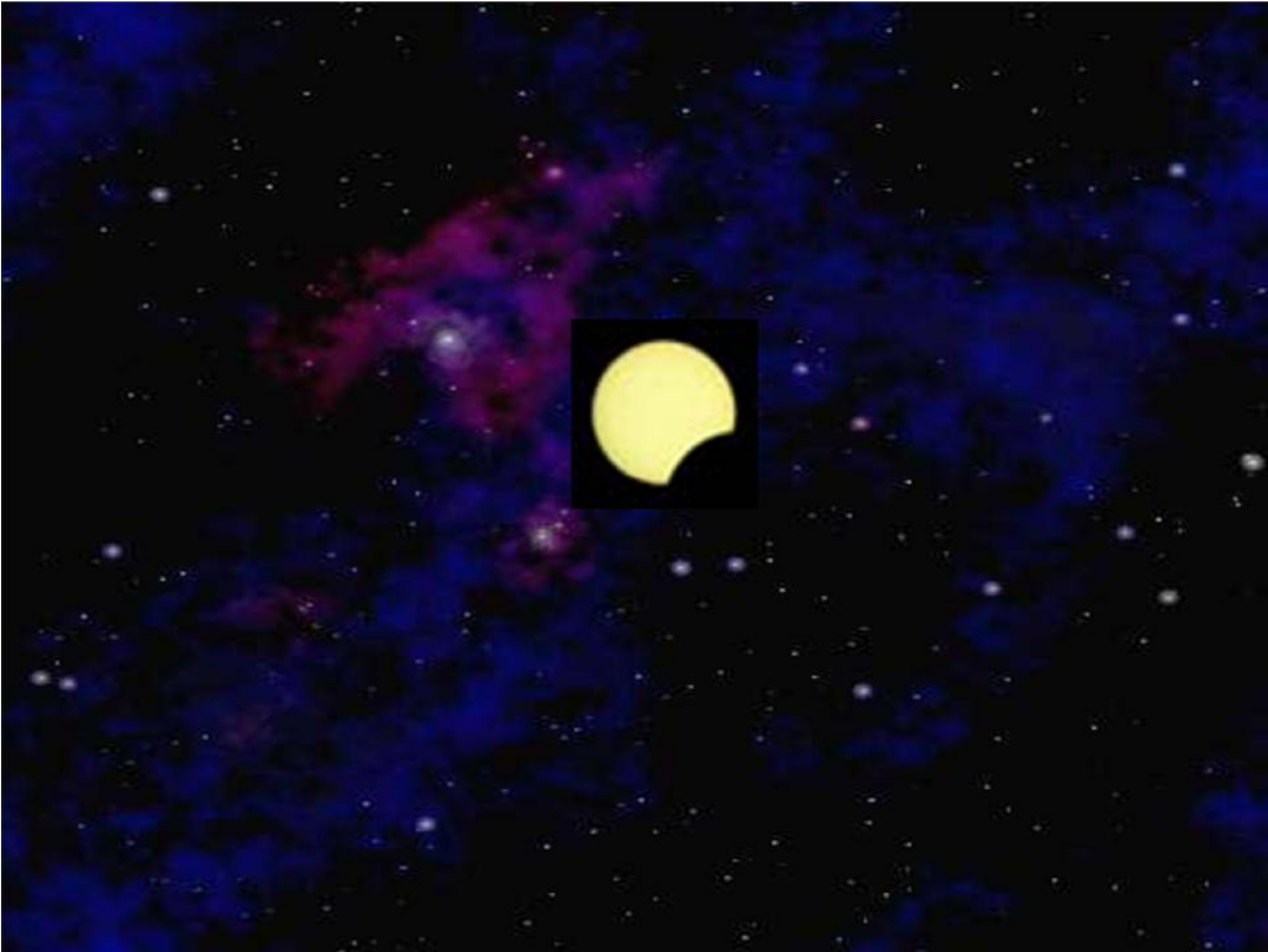
© 2000
A.T. Sinclair

Espenak's Eclipse Home Page

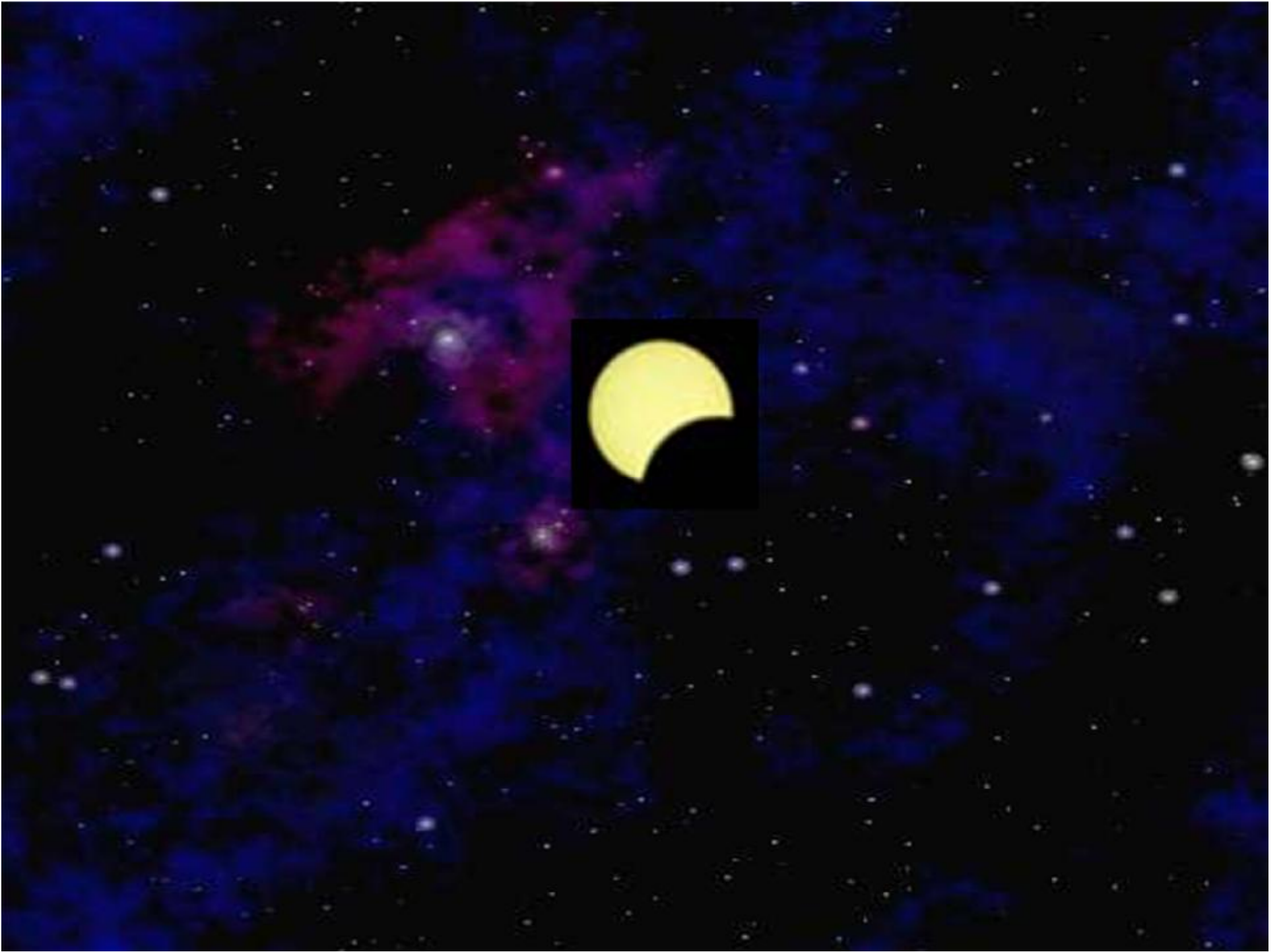
Solar Eclipse seen from Earth:

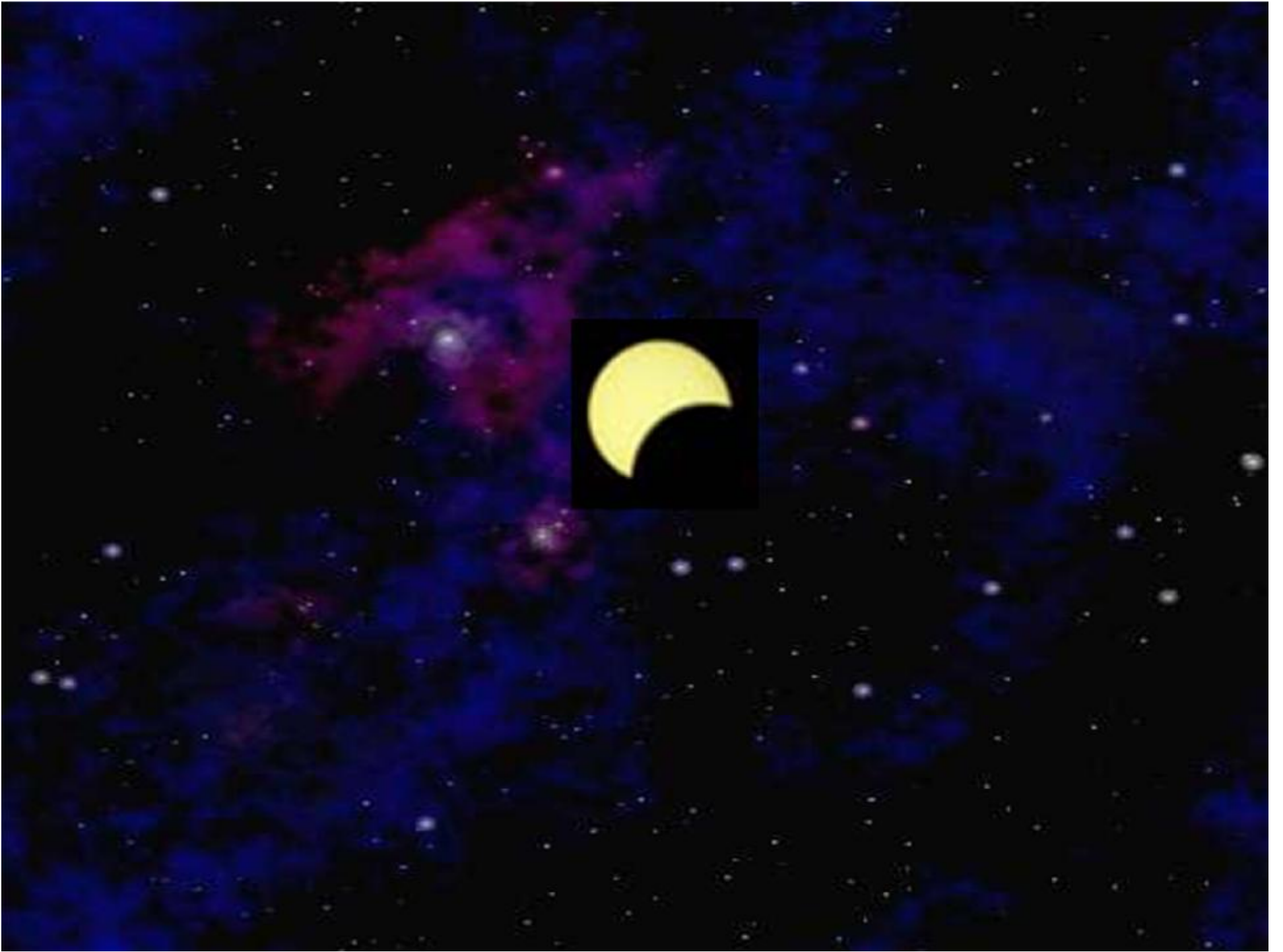


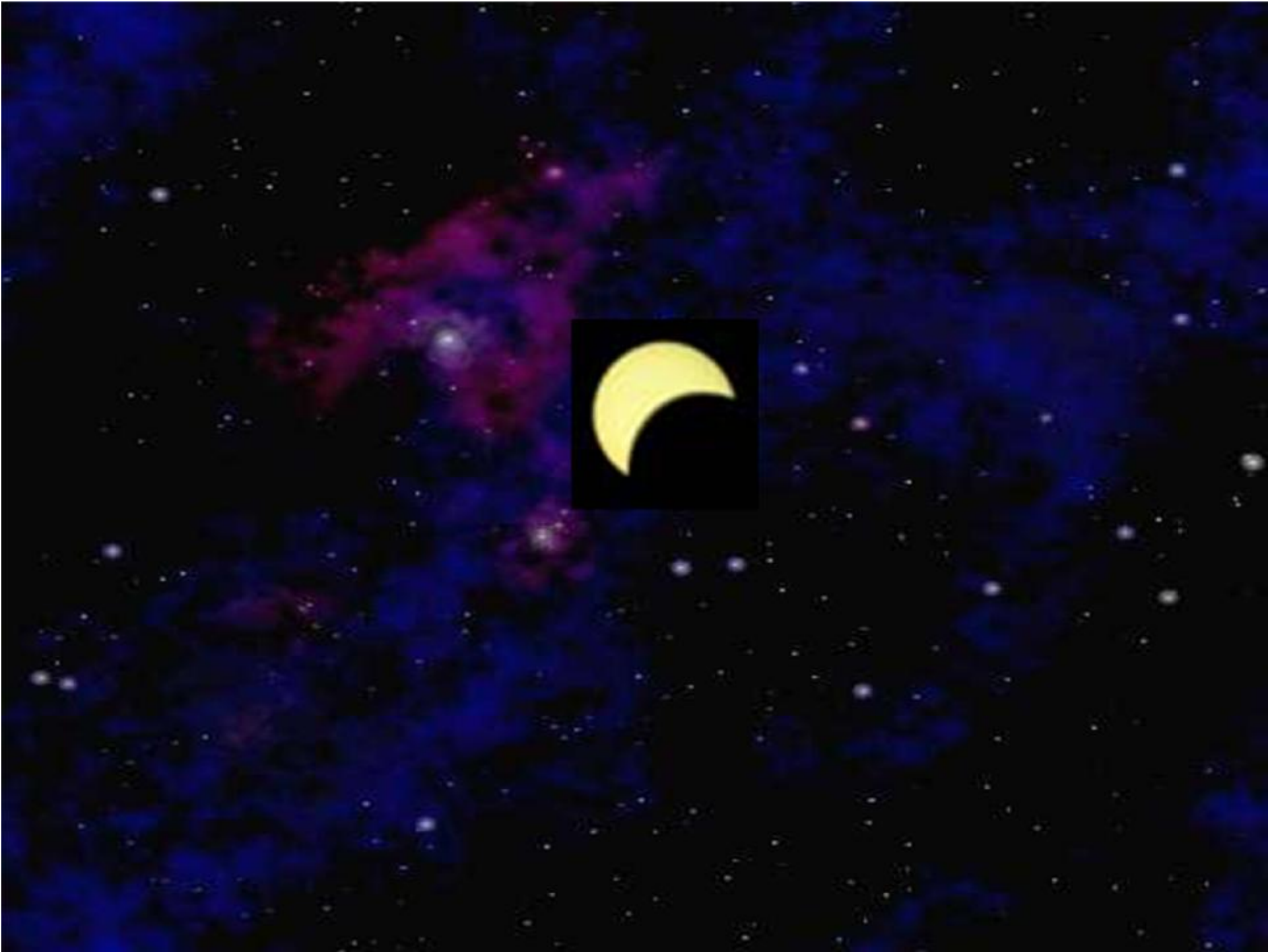


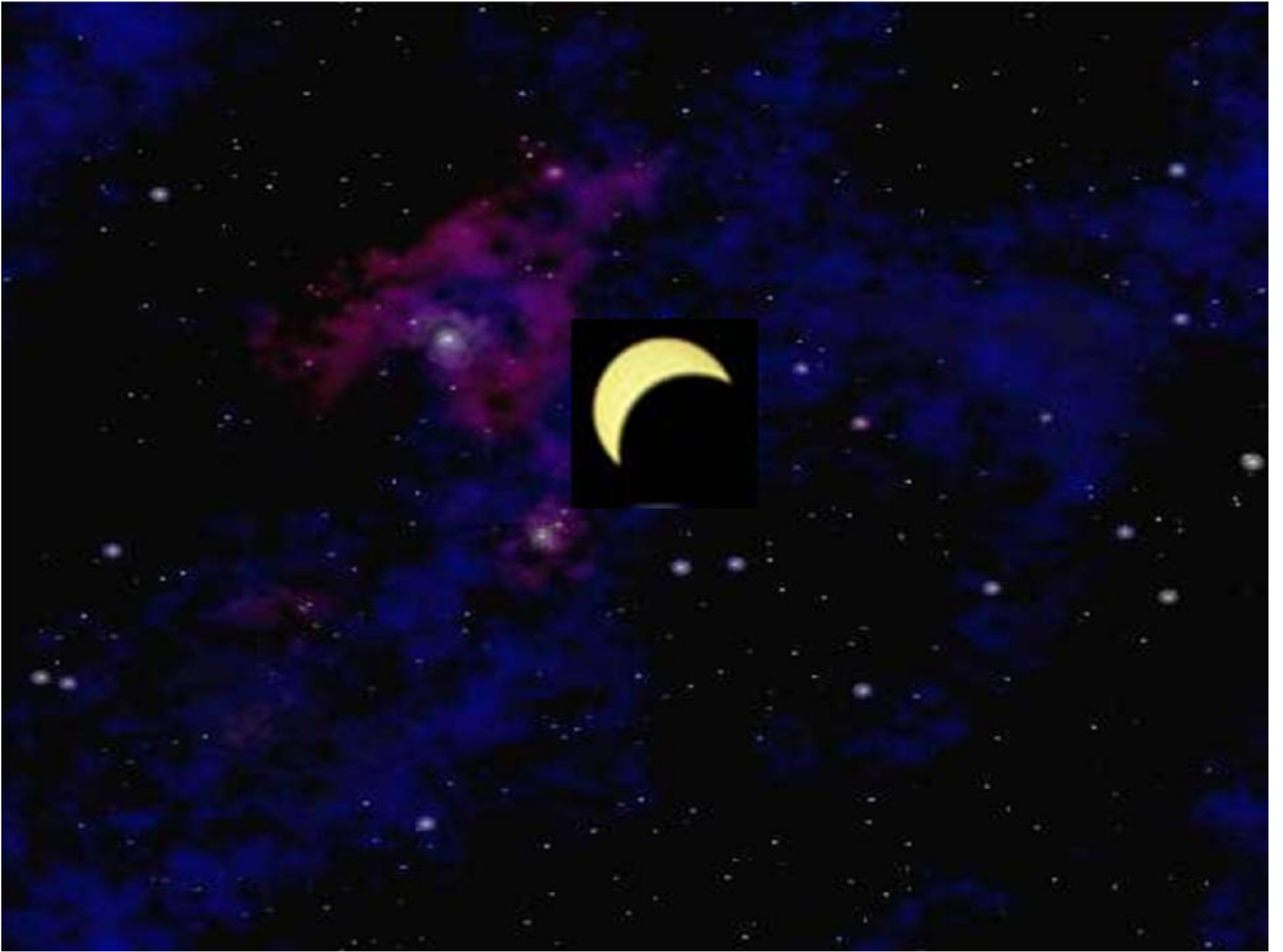


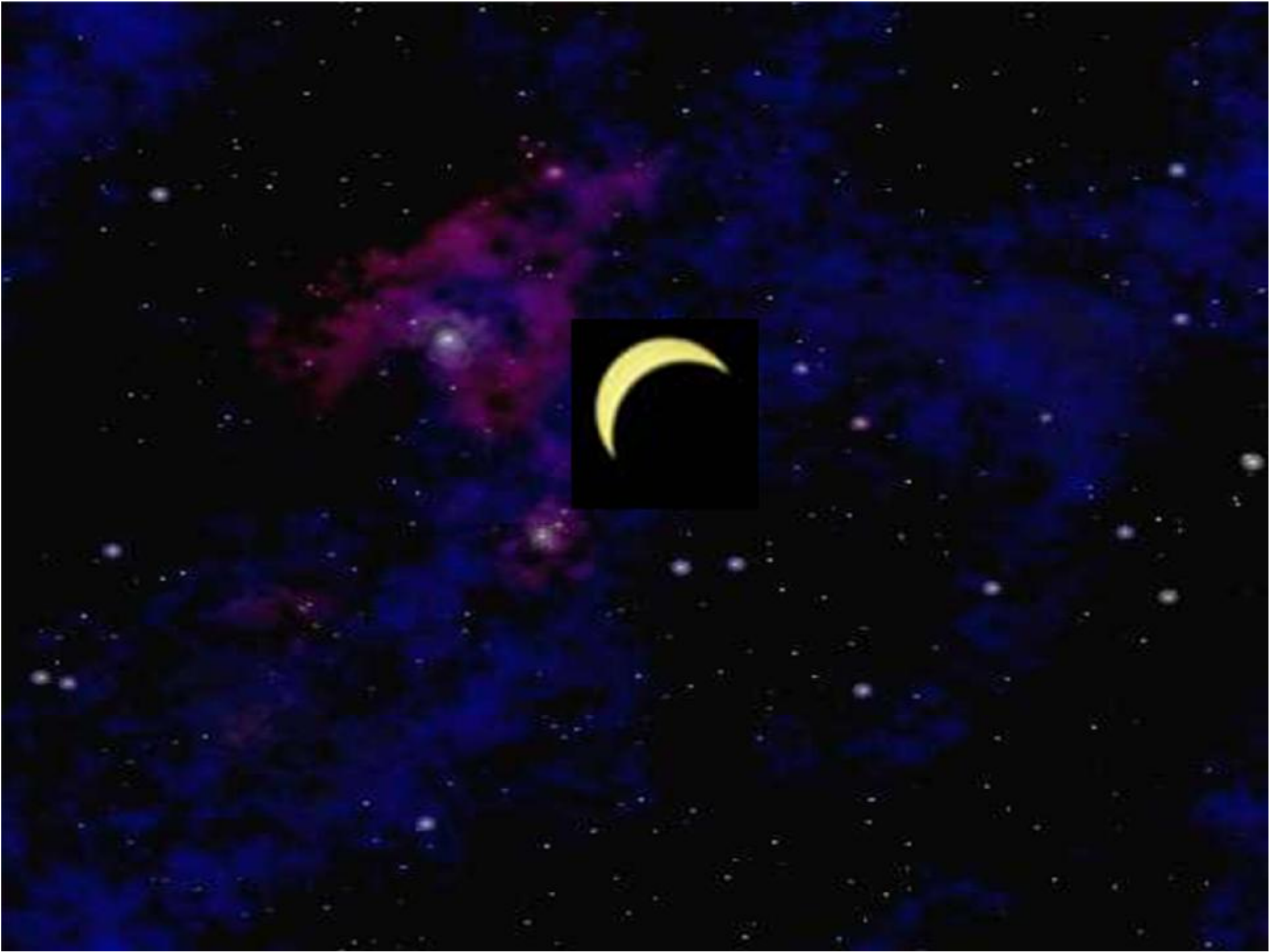


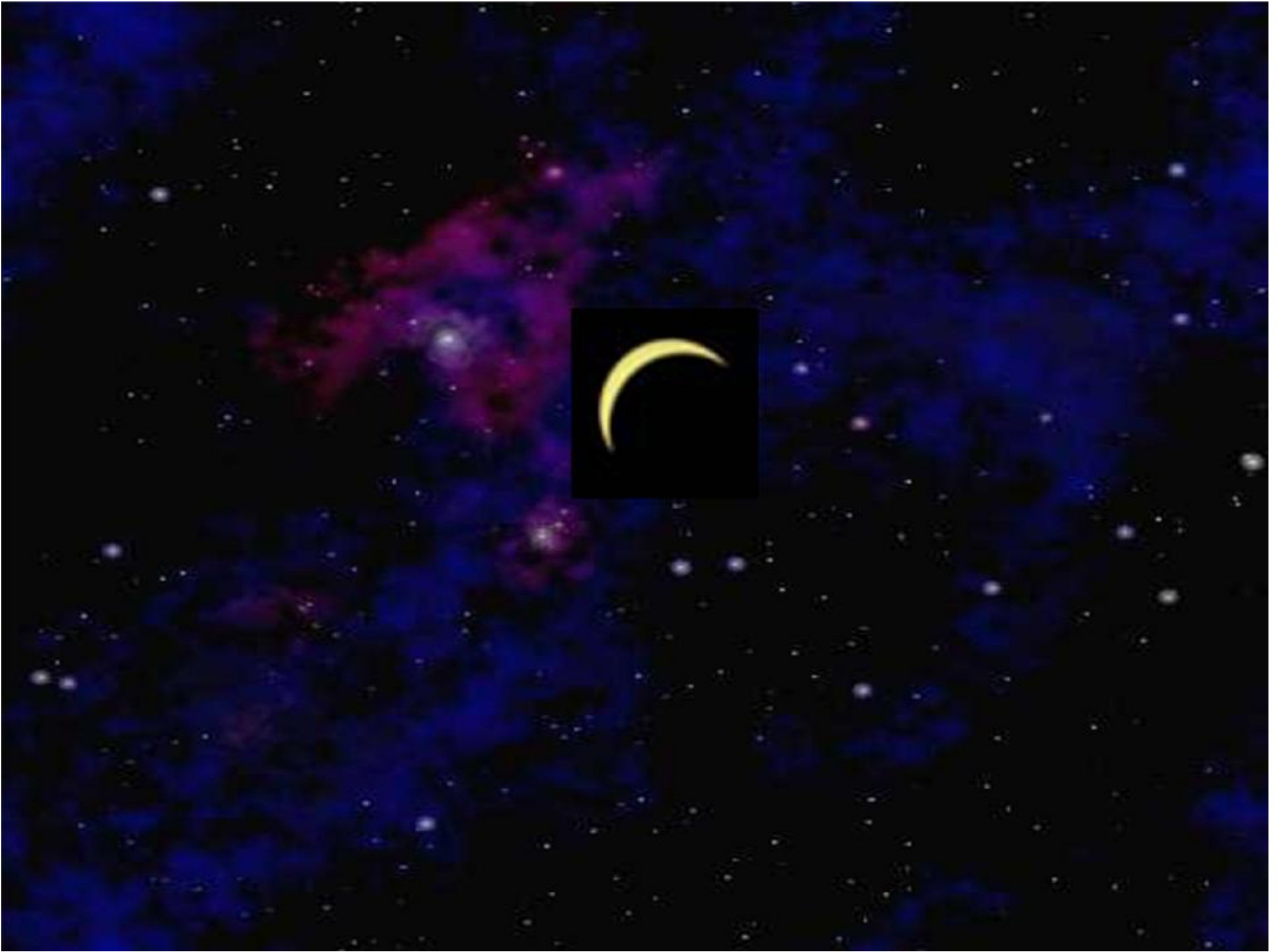


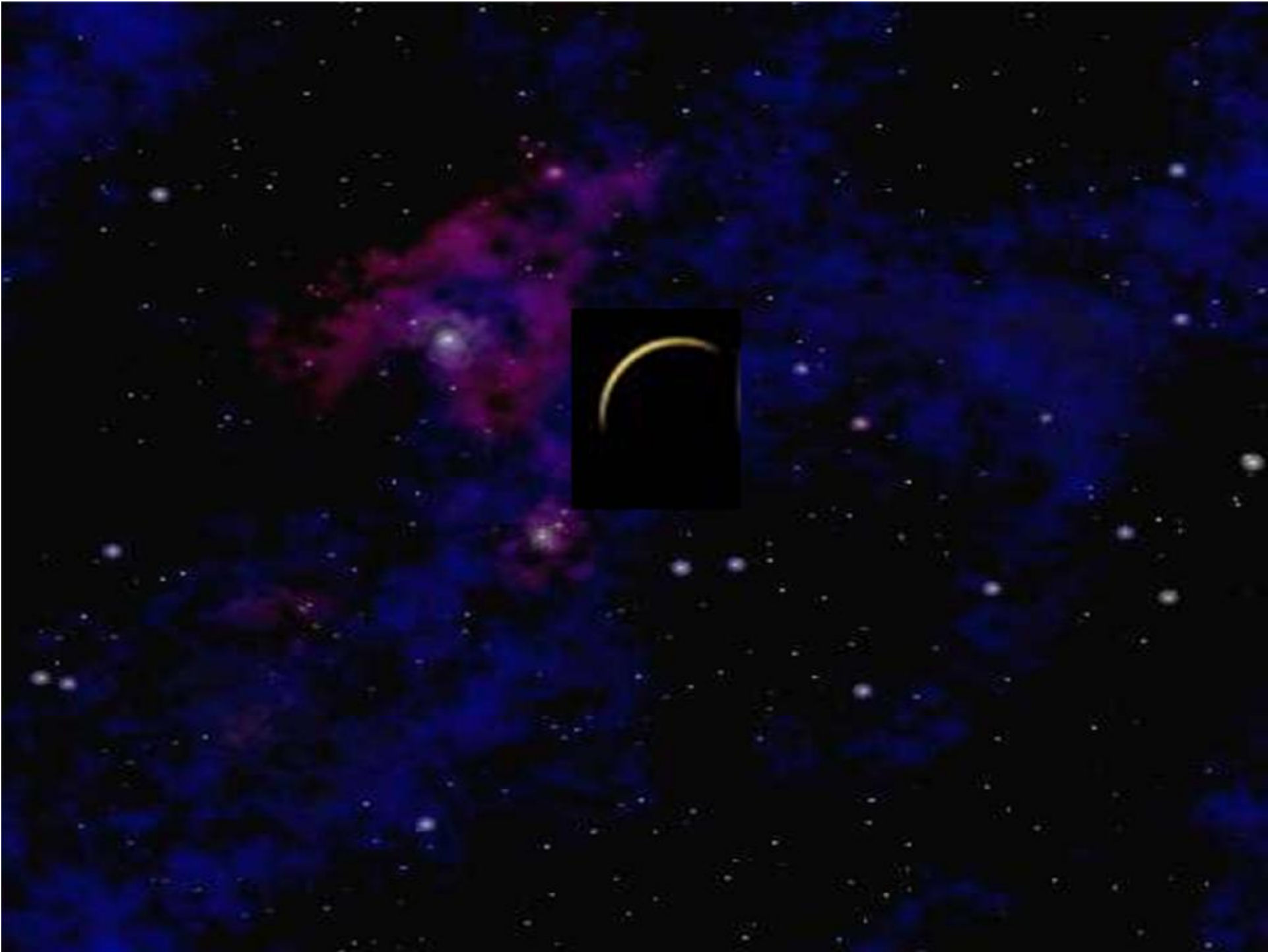


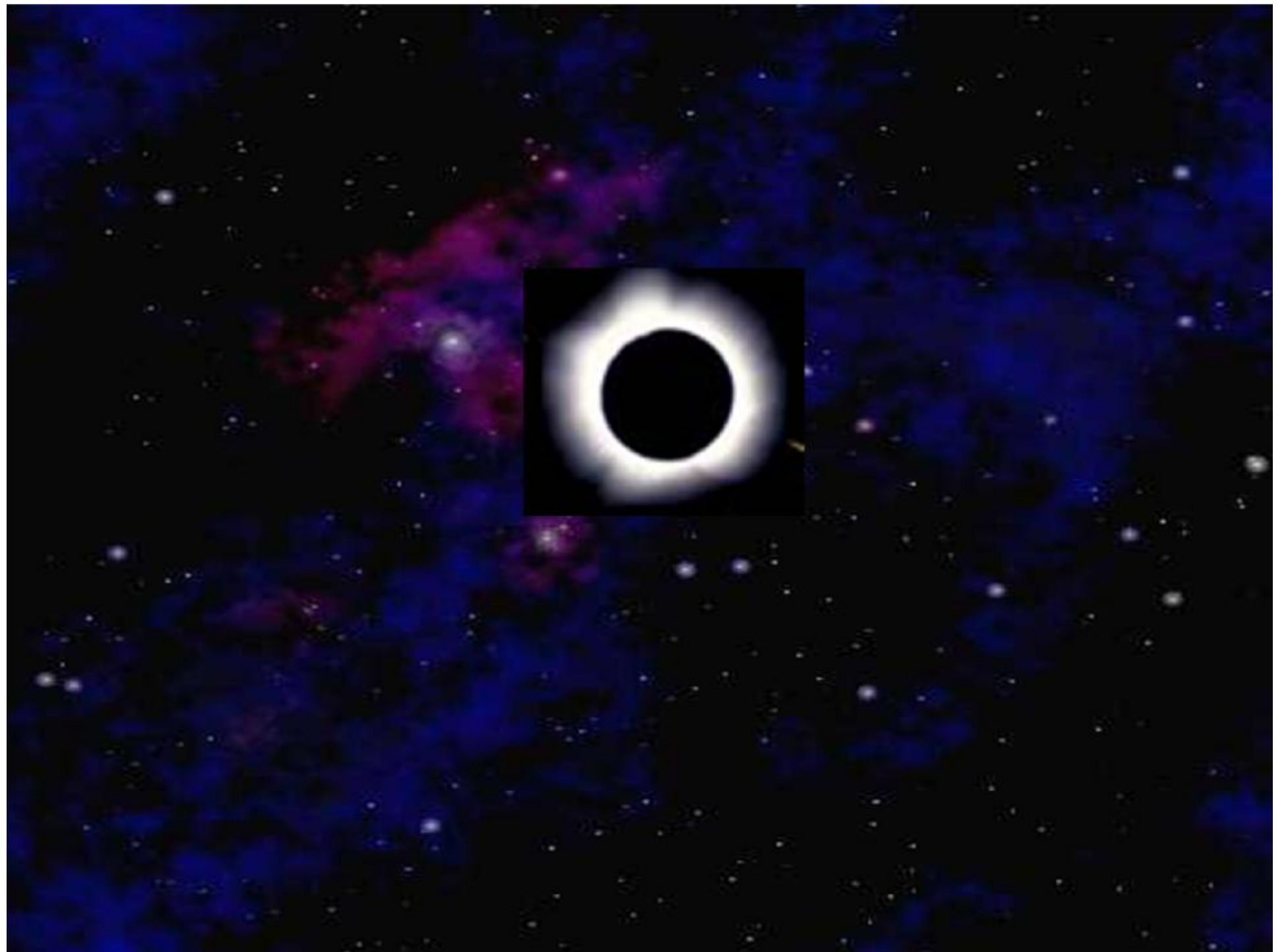


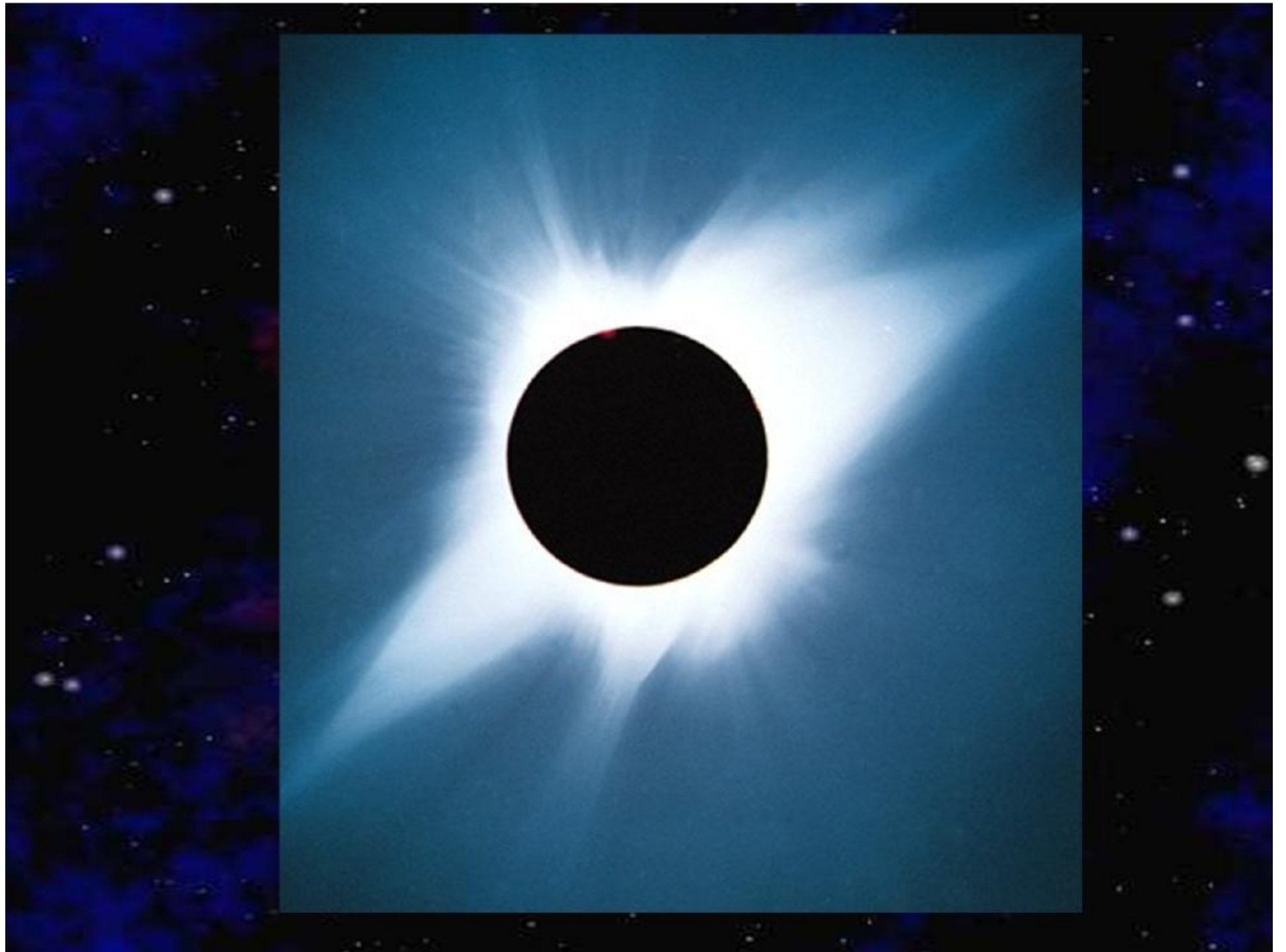


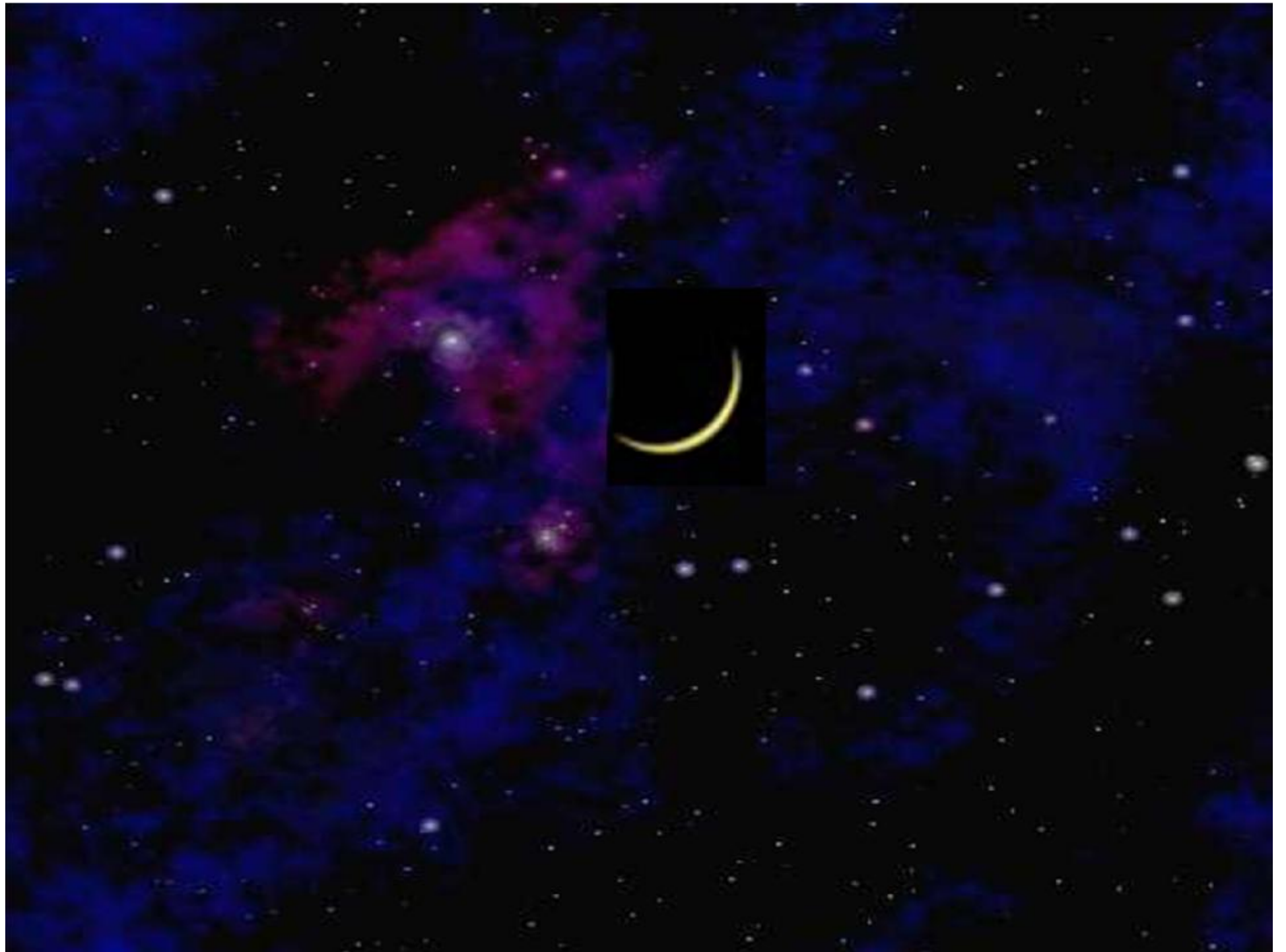


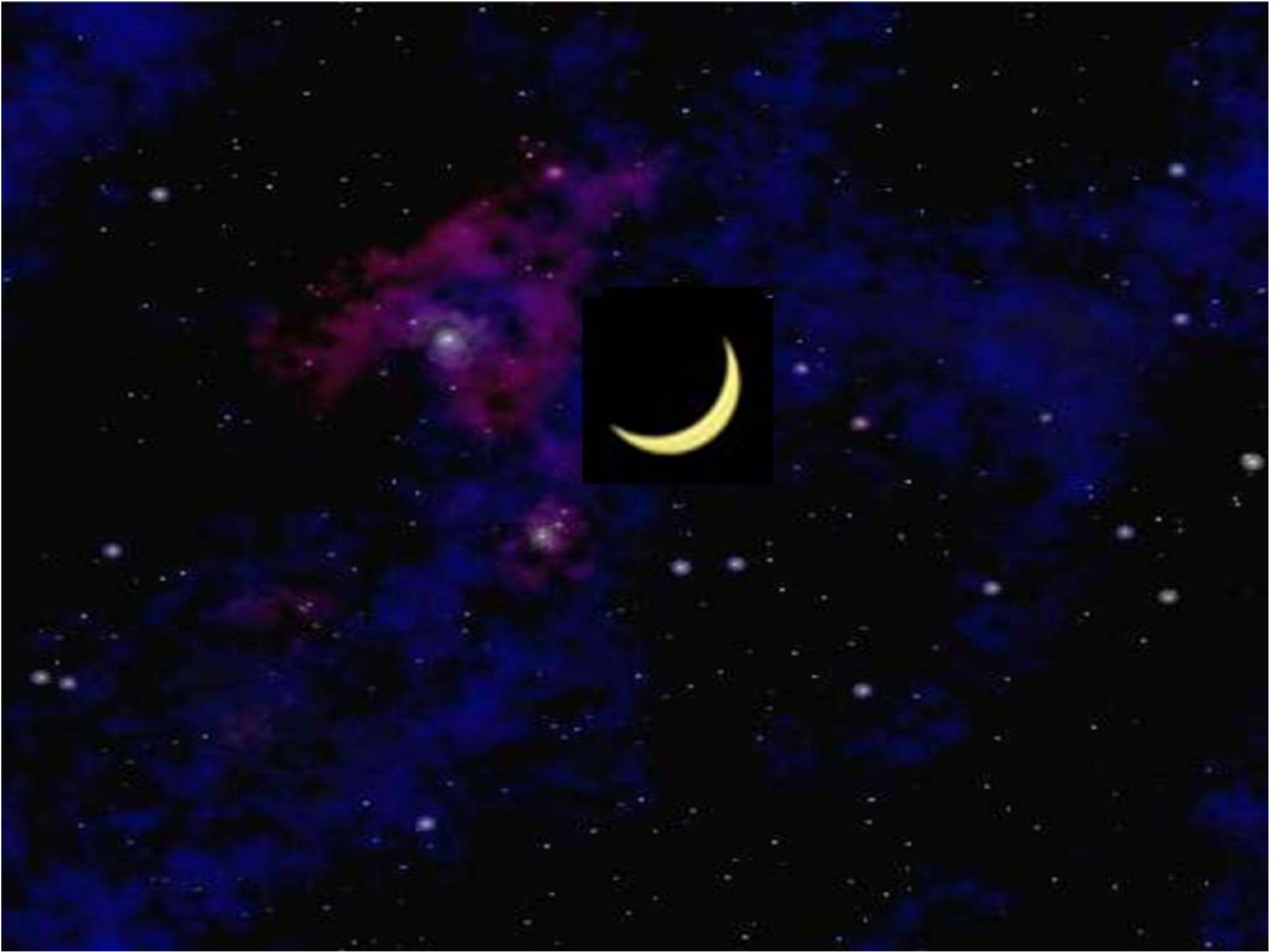


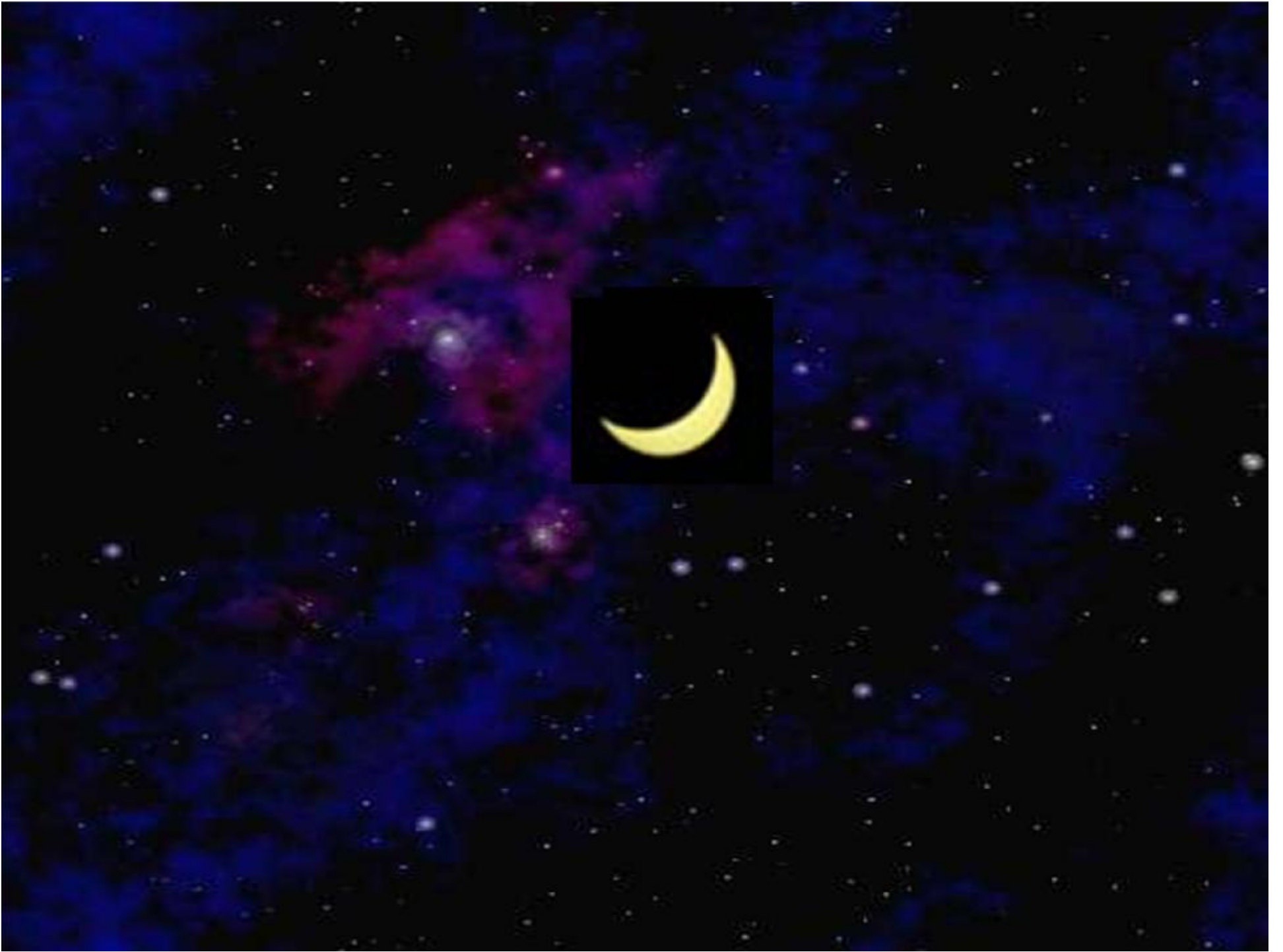


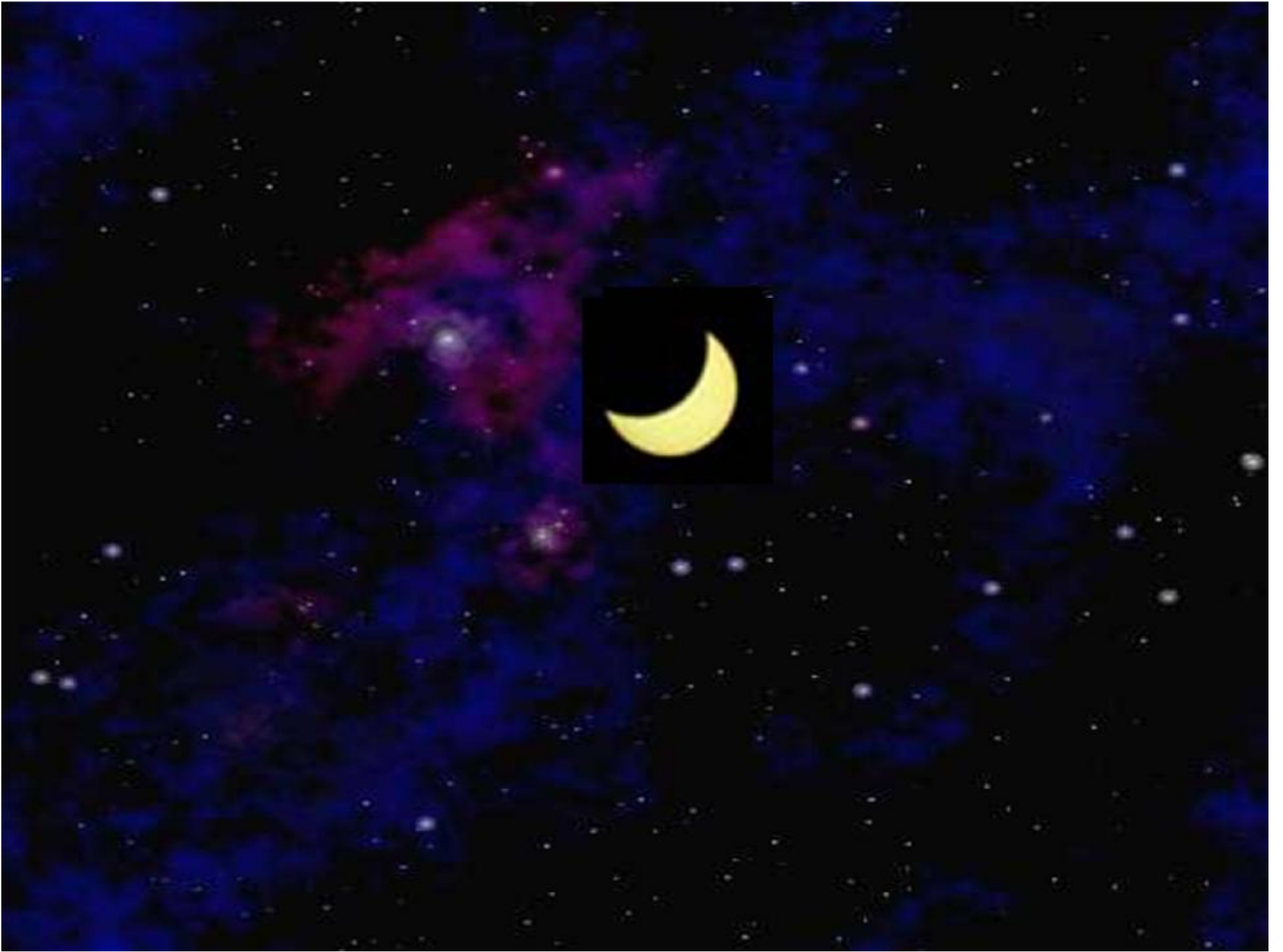


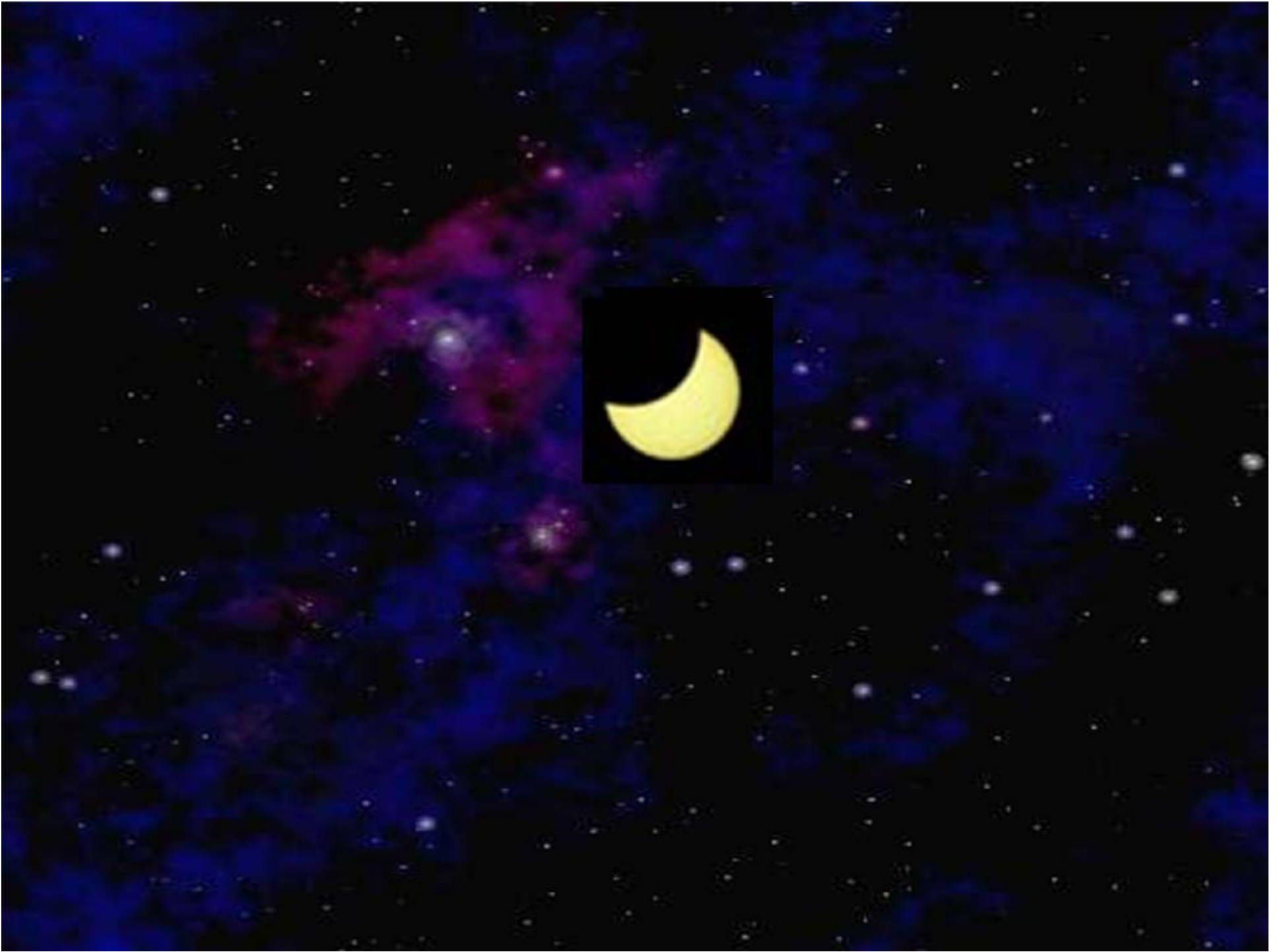


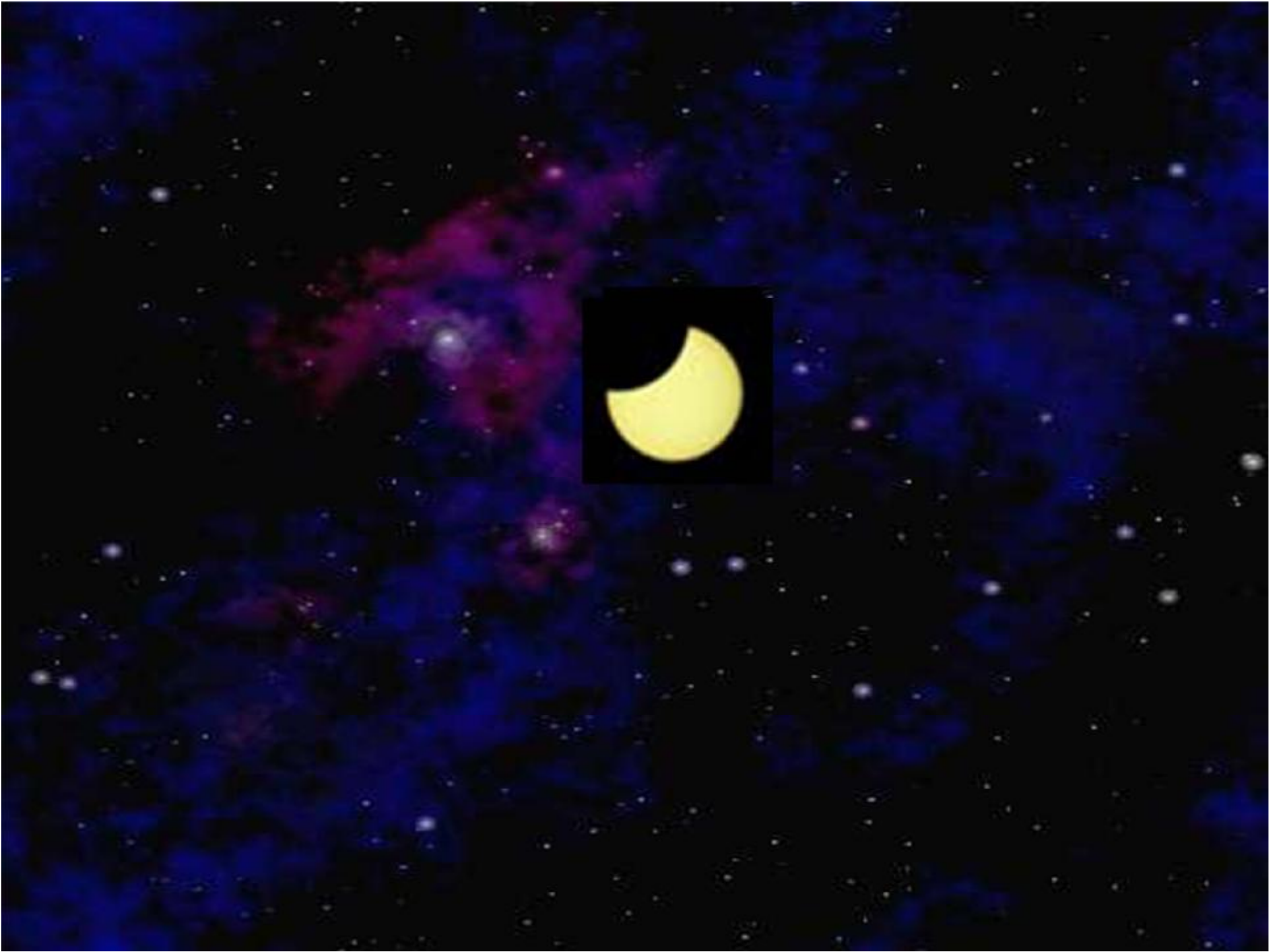


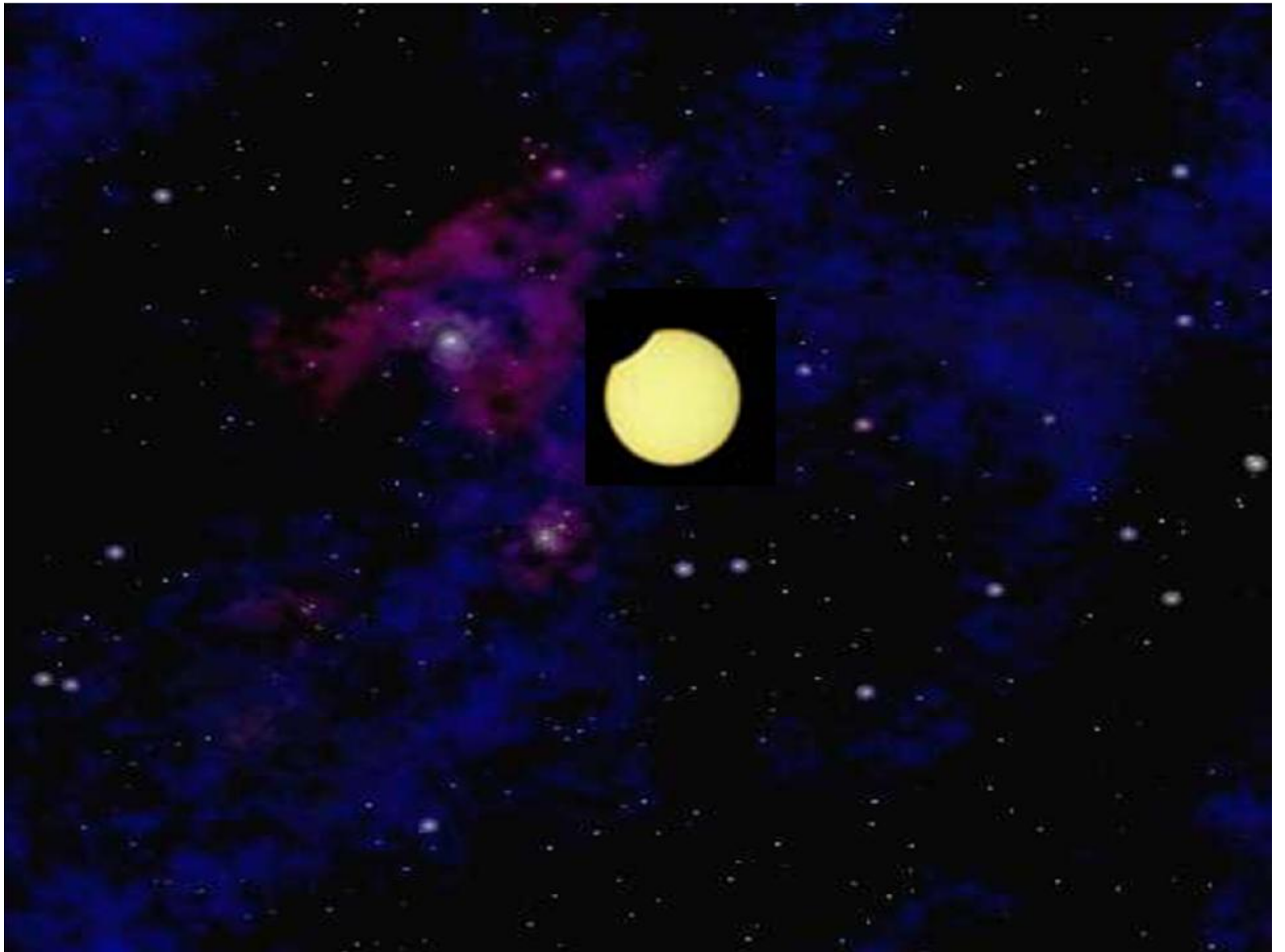








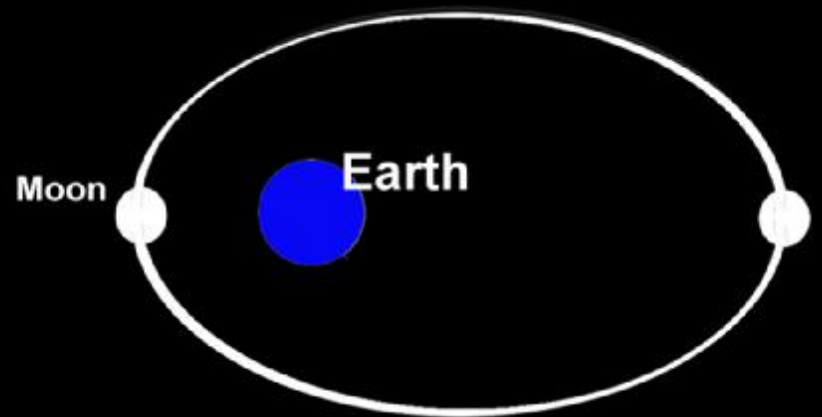




Annular Eclipses



- § The angular size of the Moon can appear smaller than the angular size of the Sun.
- § The Moon's orbit is not a circle, but instead is an *ellipse*.

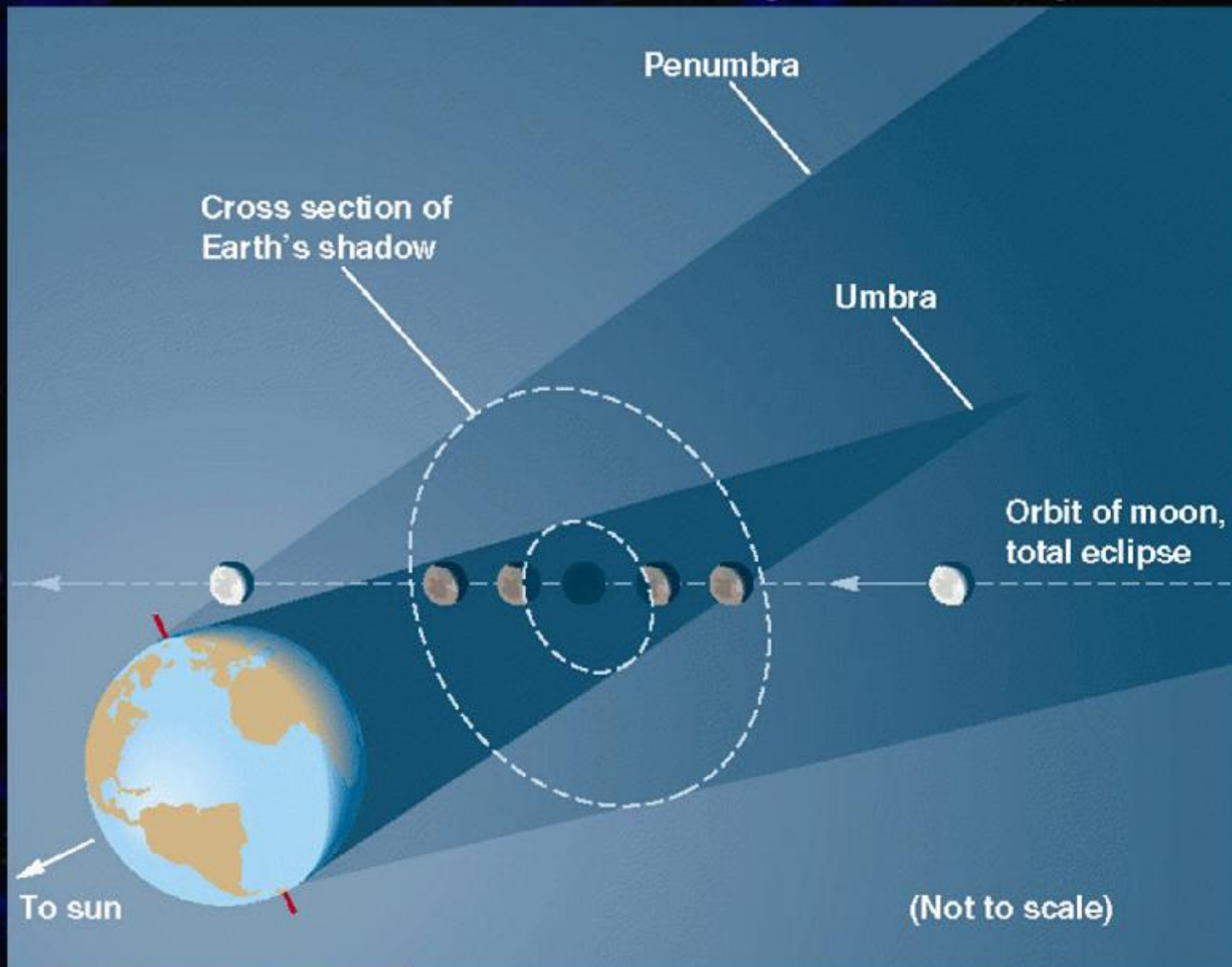


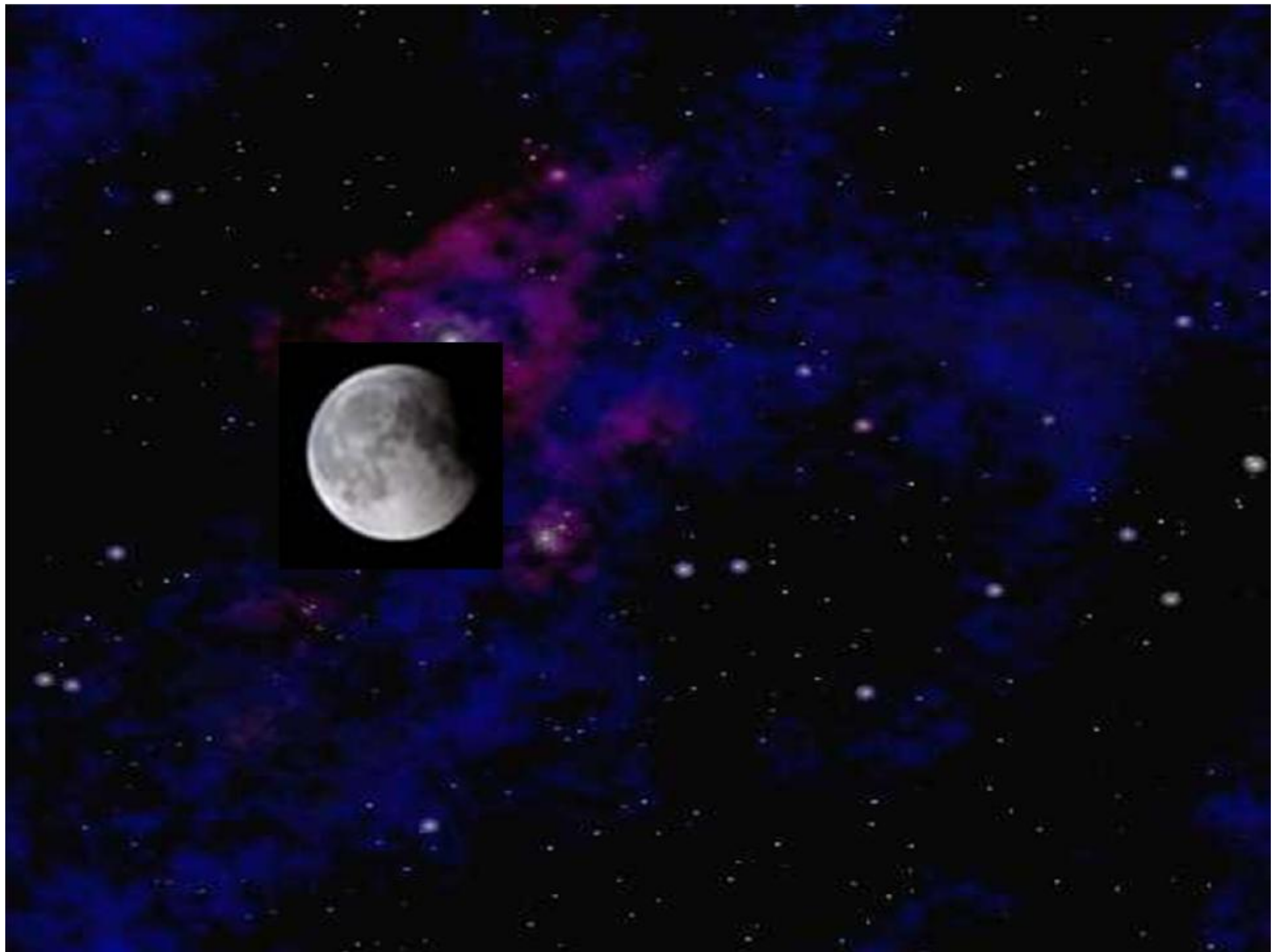


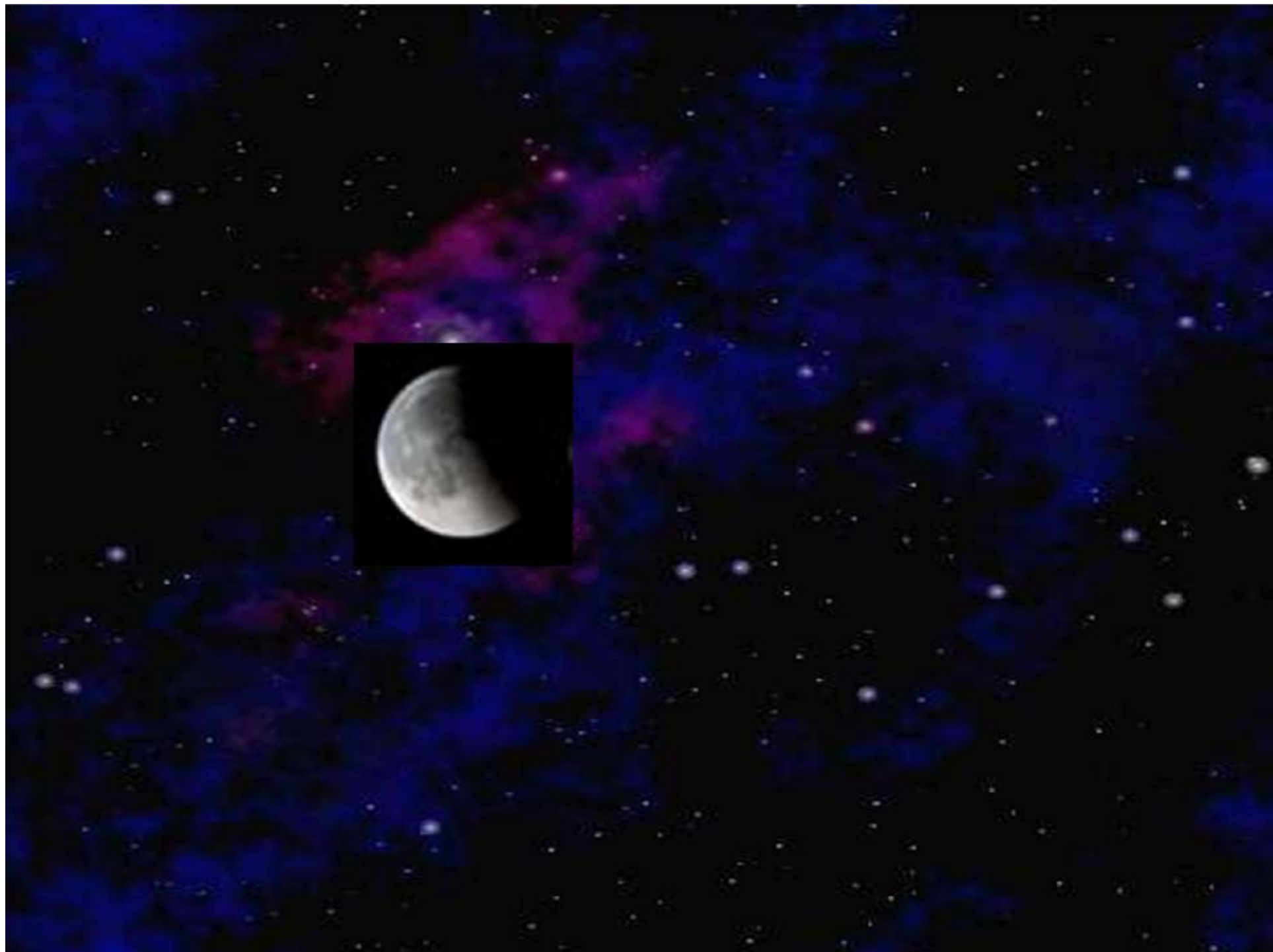
Lunar Eclipses

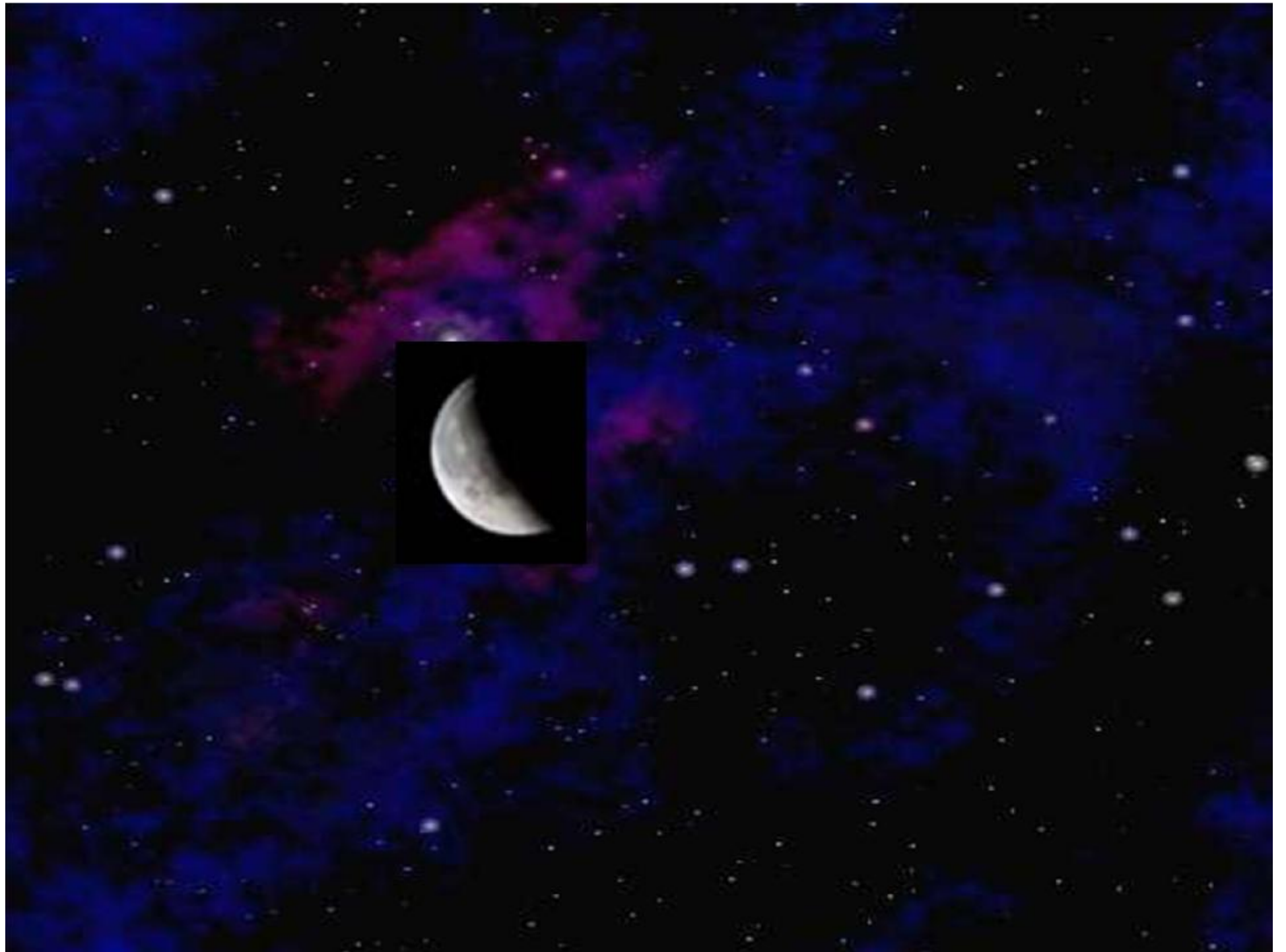
Occurs when the Moon passes through the shadow cast by the Earth.

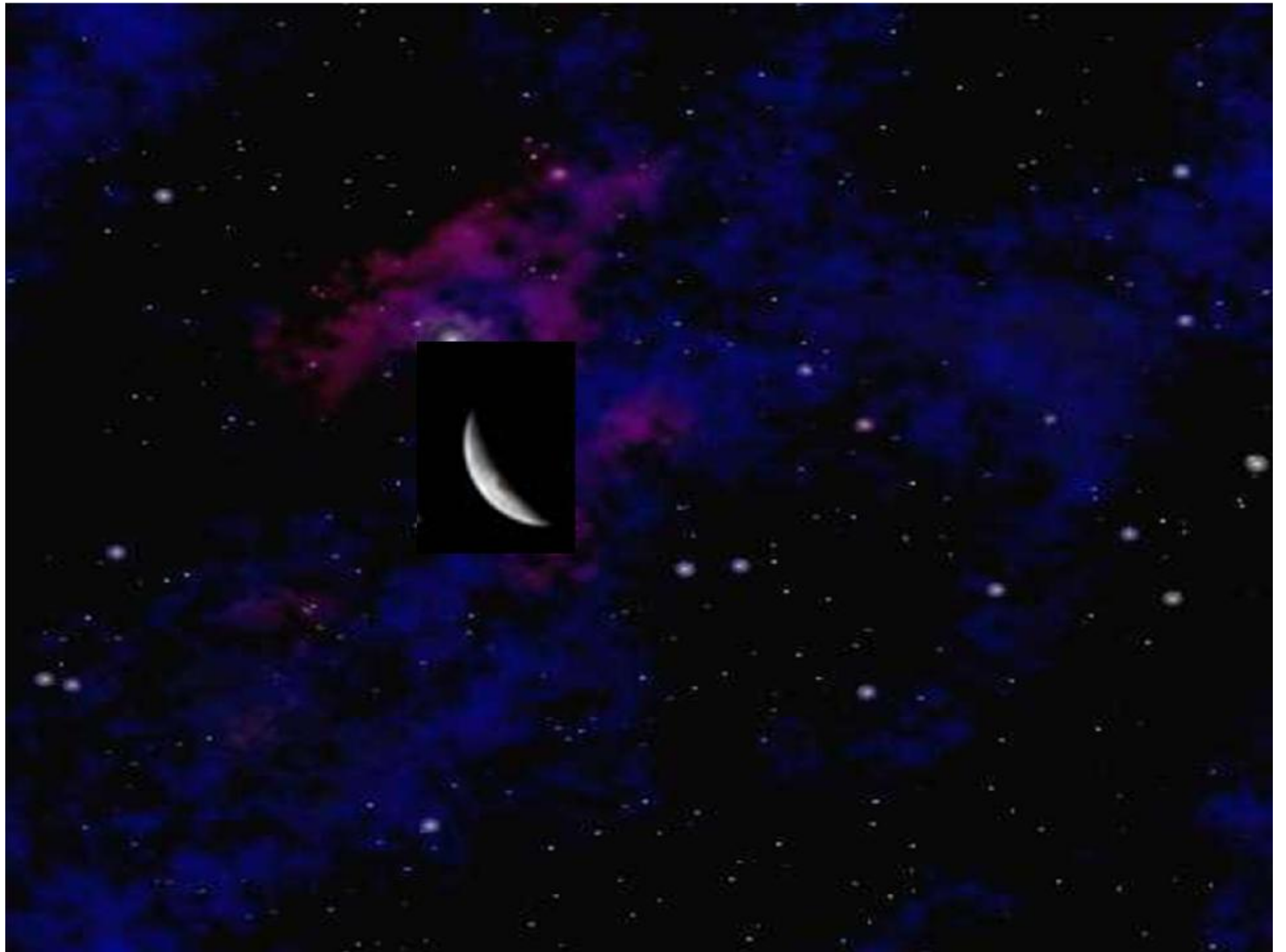
Can only occur at FULL MOON.



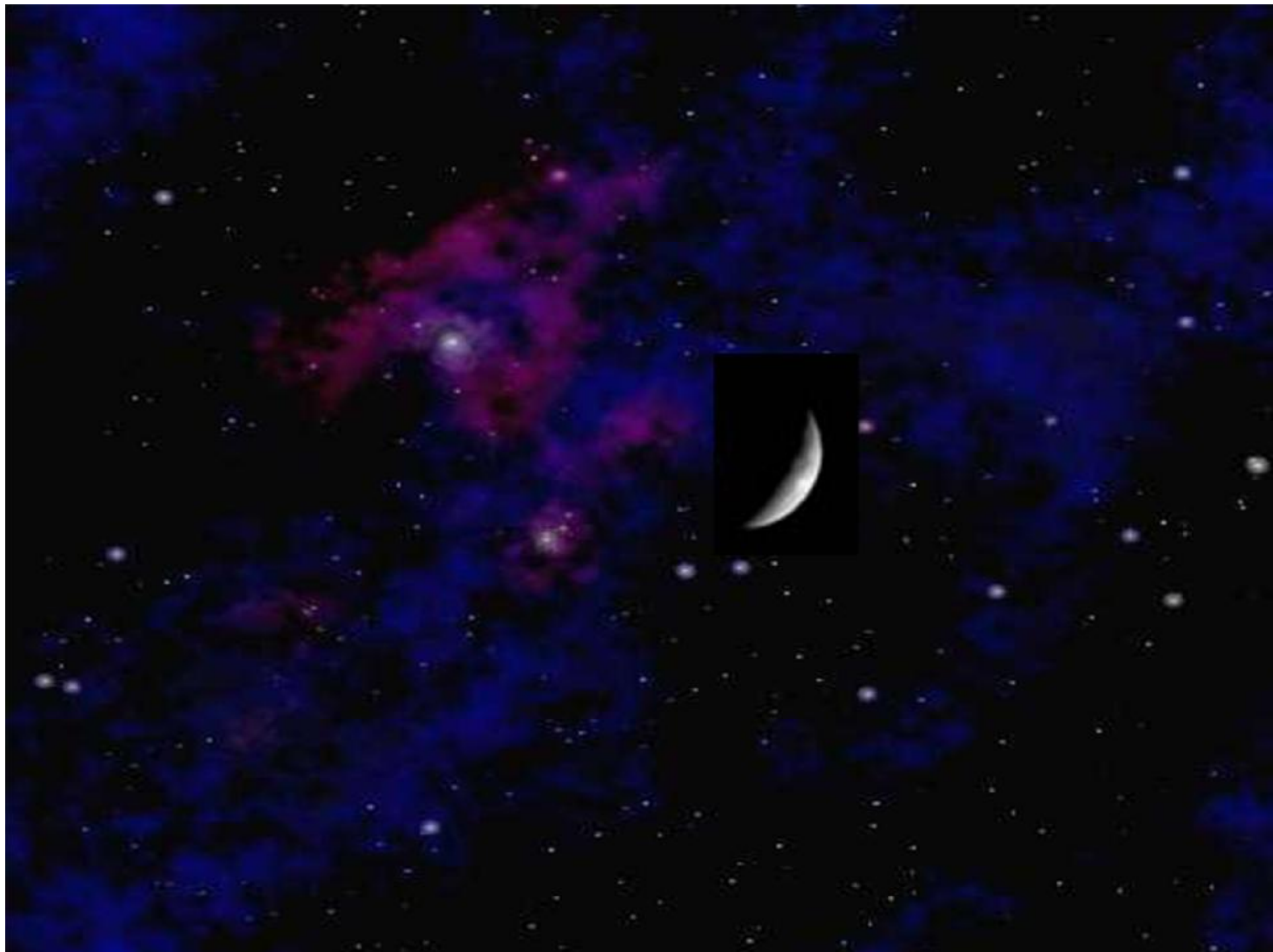


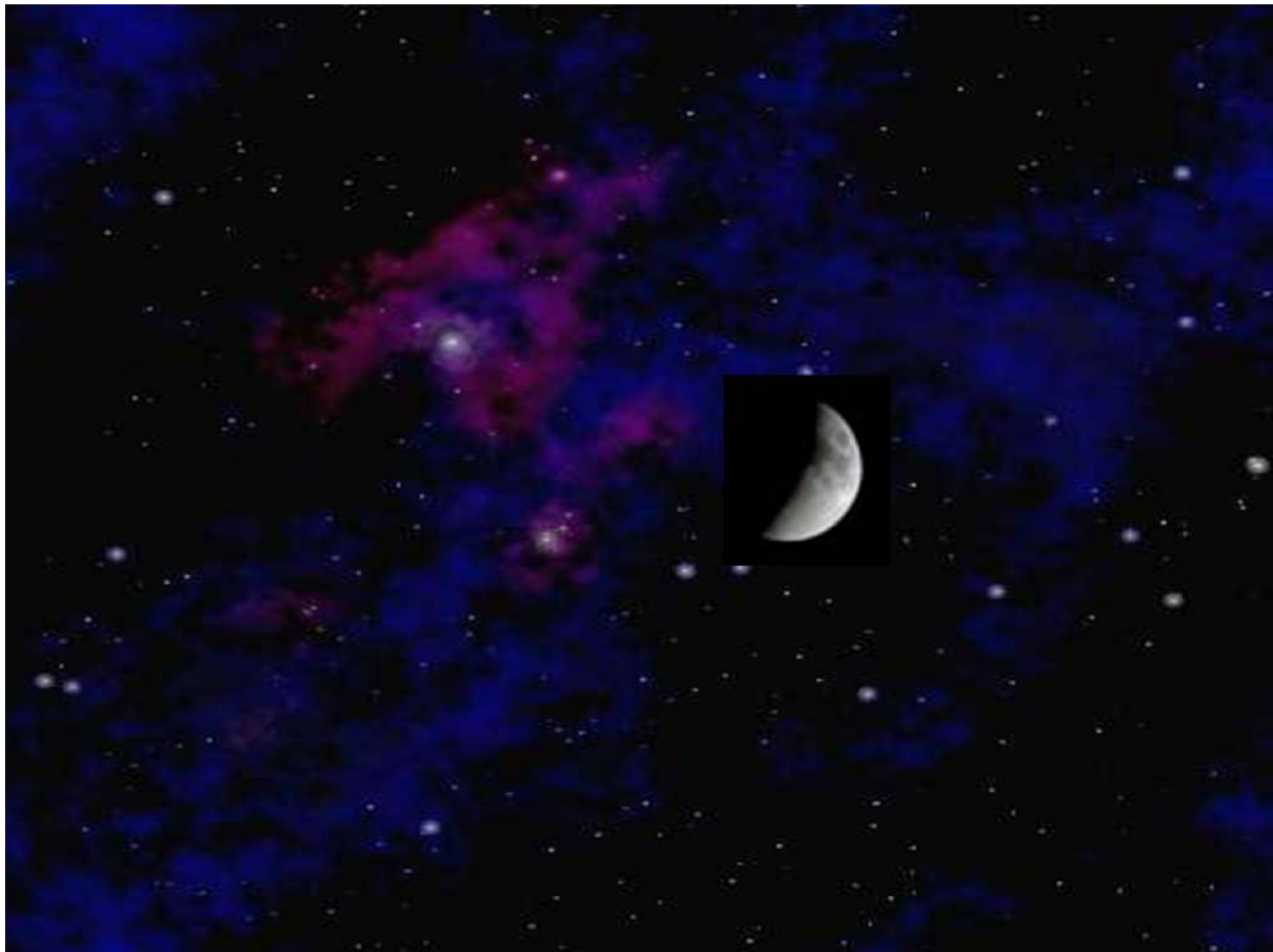


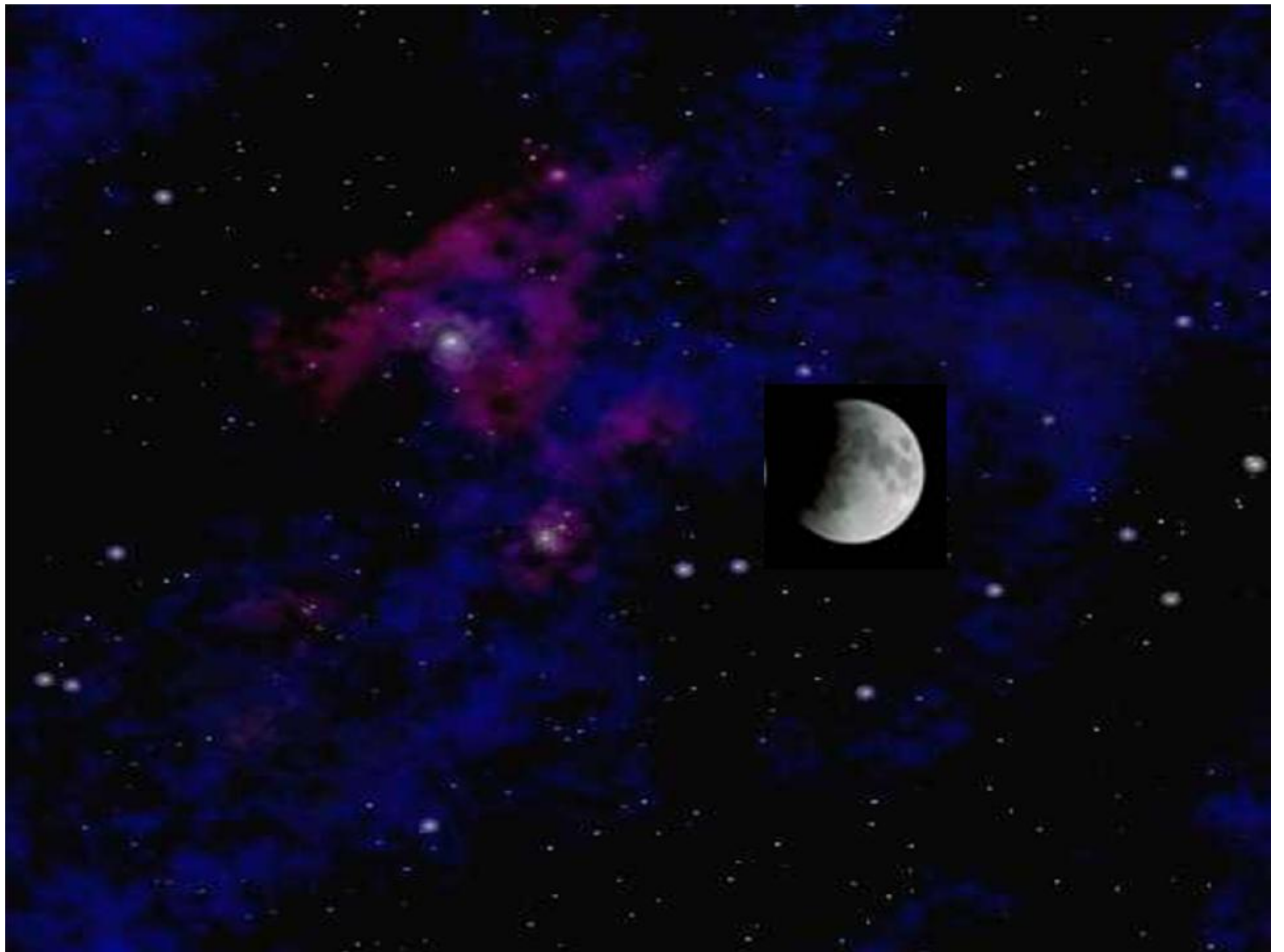


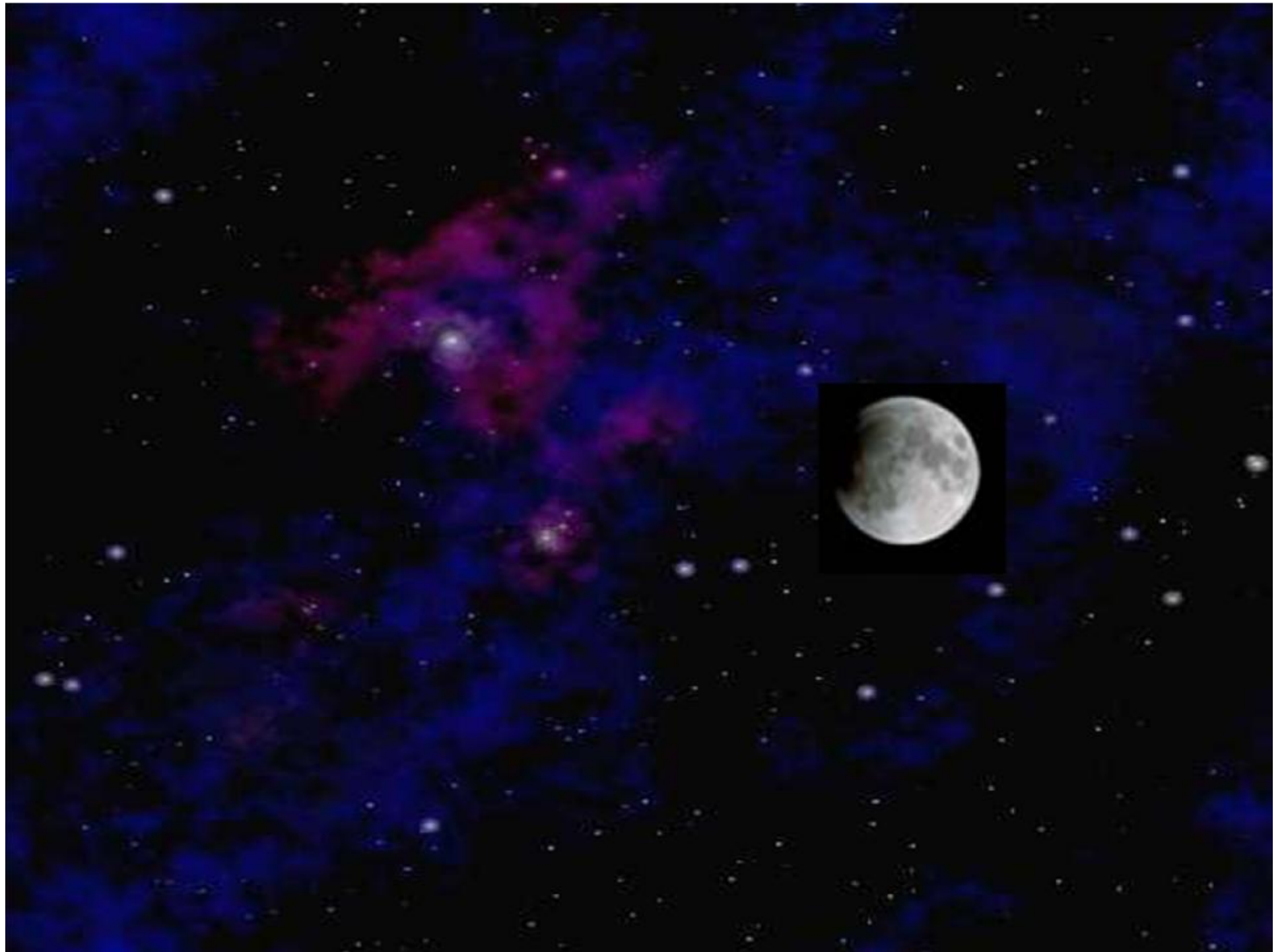














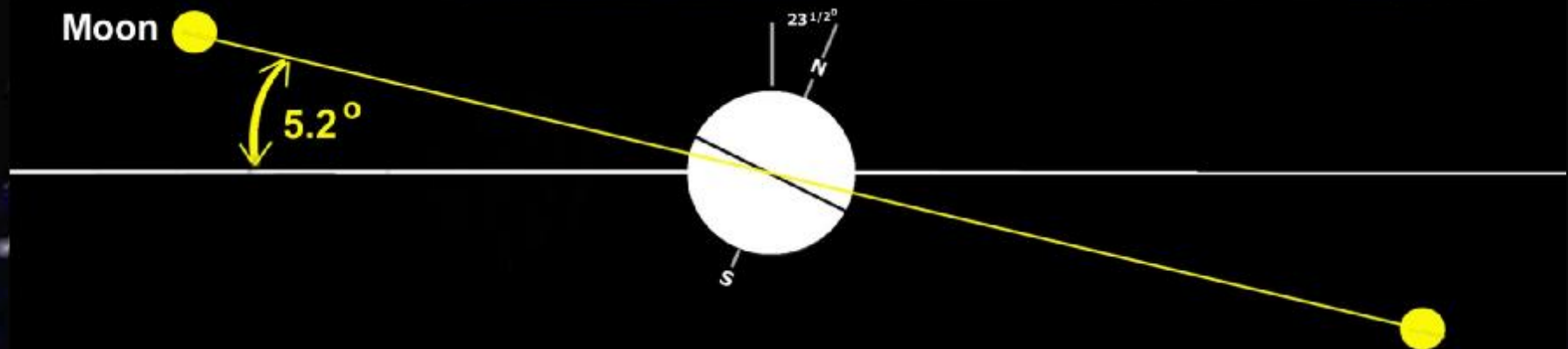
**If a LUNAR ECLIPSE occurs during FULL
MOON...**

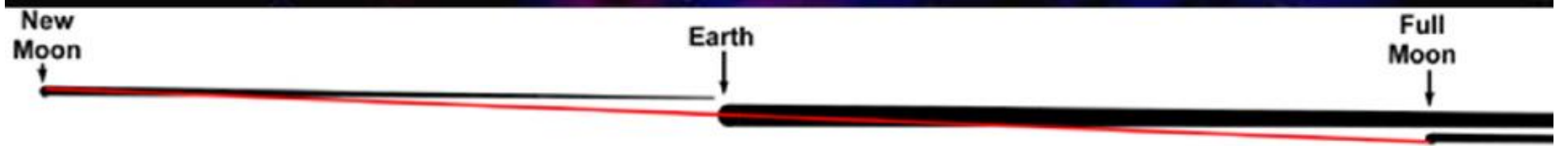
**And a SOLAR ECLIPSE occurs during
NEW MOON...**

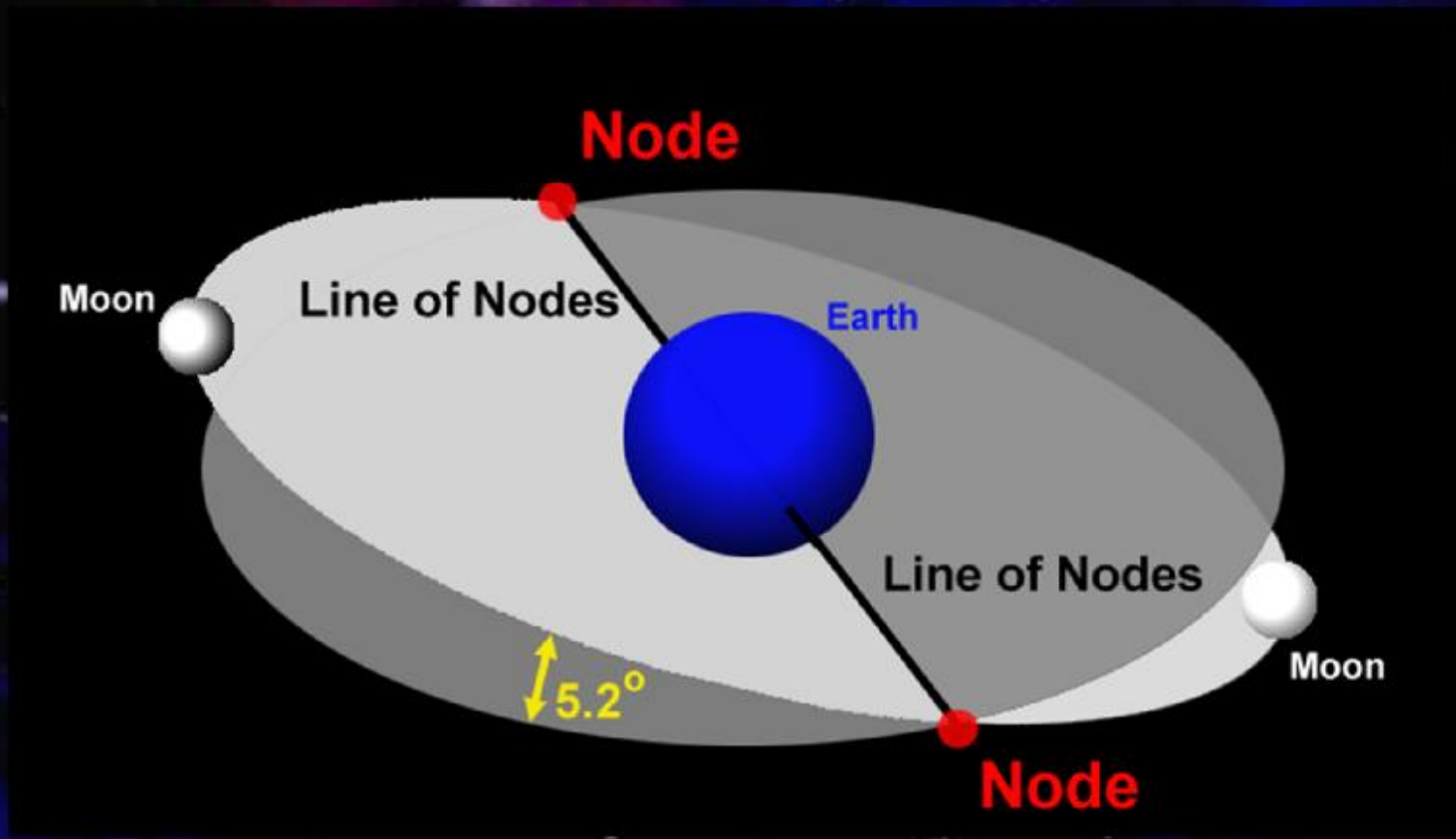
**If there is a FULL MOON and a NEW
MOON every month...**

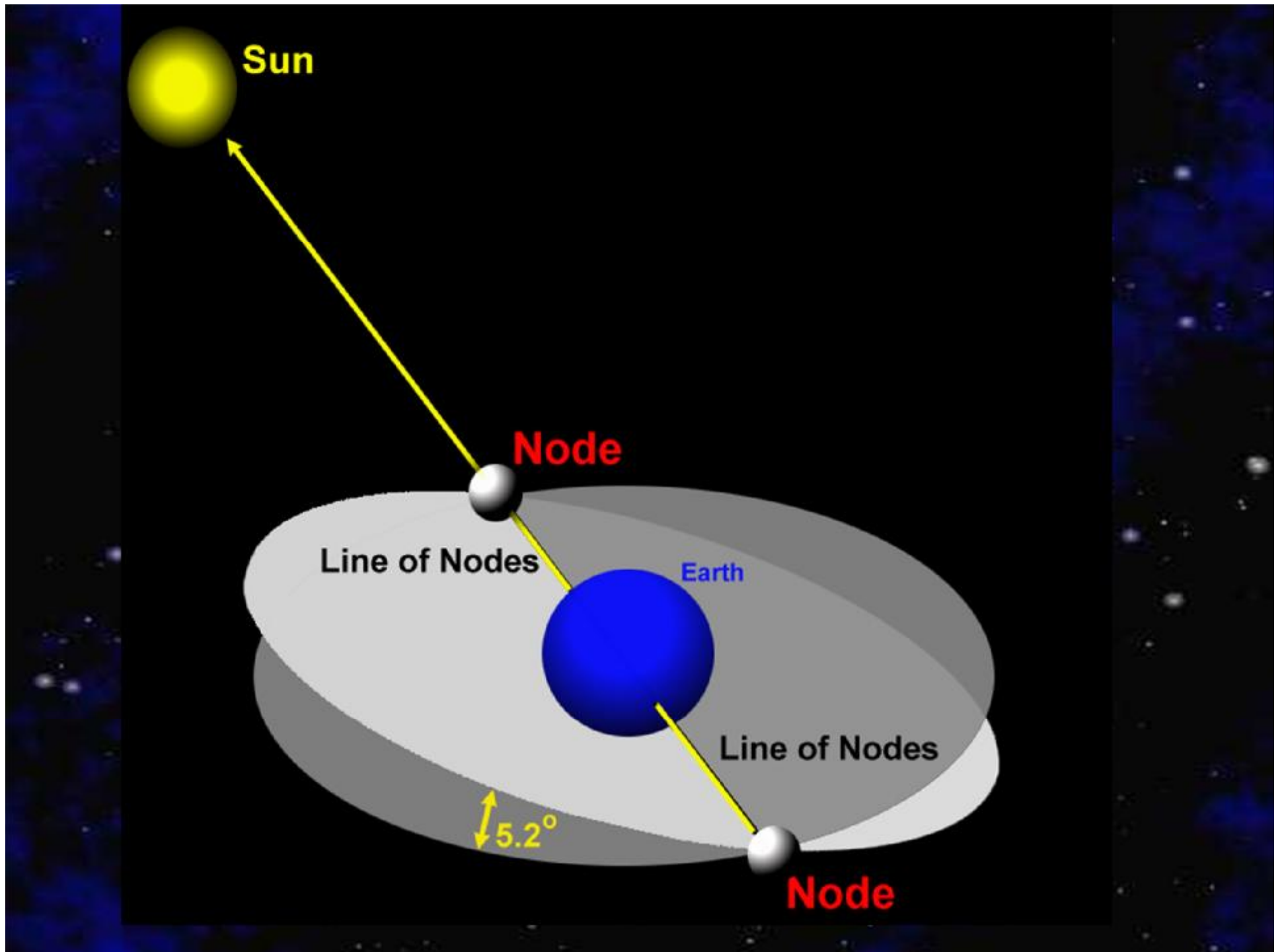
**Then why don't we experience eclipses
every month??**

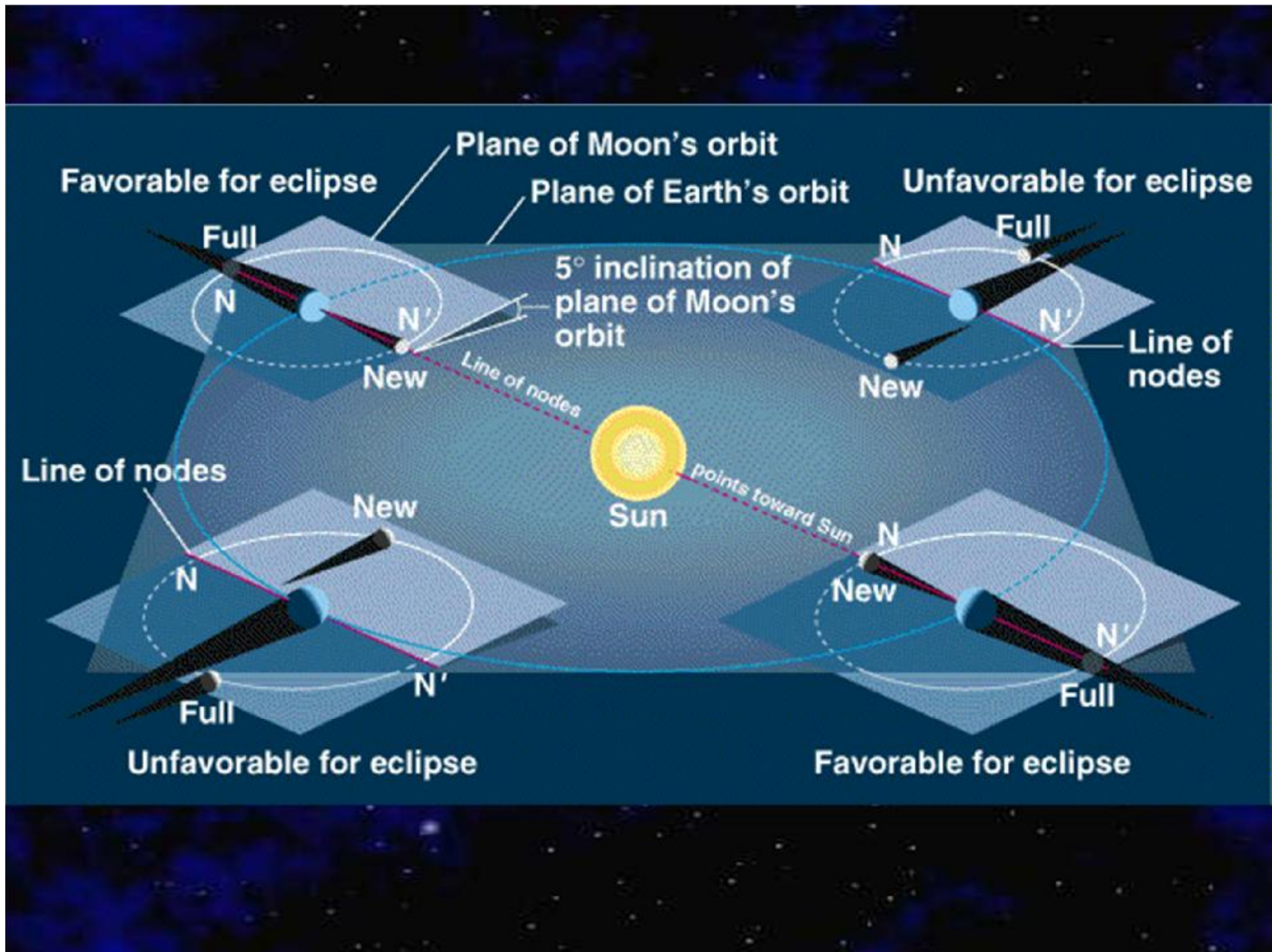
The Moon's orbit is inclined to the Earth's orbit by 5.2°











Eclipse Predictions

How do we know how long an eclipse will last?

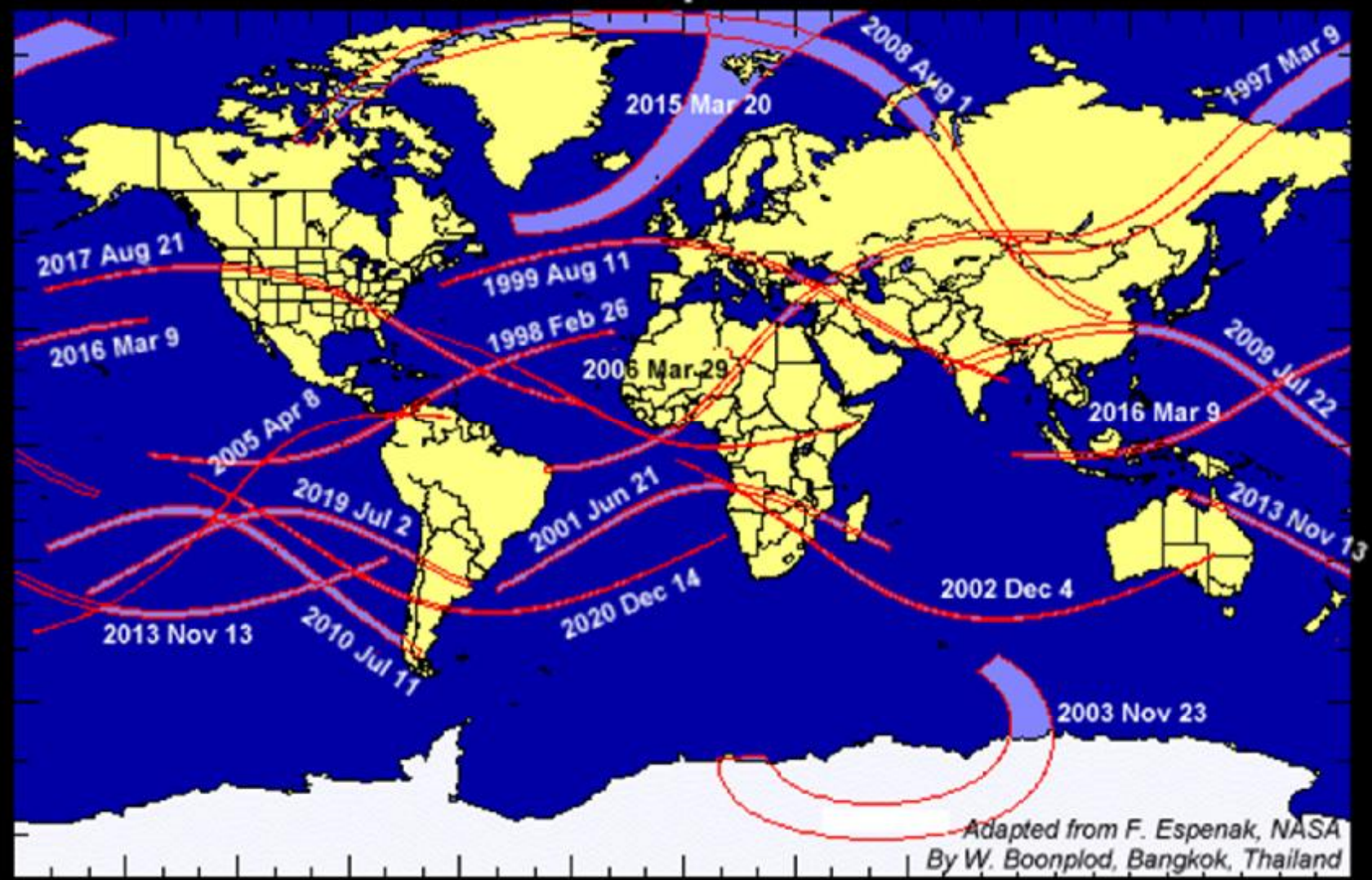
How do we know where eclipses will occur?

Length of Totality

The shortest time totality can last is
a brief moment.

The longest time totality can last is
7½ minutes

Total Solar Eclipse : 1997-2020



Adapted from F. Espenak, NASA
By W. Boonplod, Bangkok, Thailand



Eclipses generally come in twos:

July 11, 1991

LONG total solar eclipse

Jan 4, 1992

Annular solar eclipse

