THE HISTORY OF ASTRONOMY

"It is therefore impossible that reason not previously instructed should imagine anything other than that the Earth is a kind of vast house with the vault of the sky placed on top of it; it is motionless and within it the Sun being so small passes from one region to another, like a bird wandering through the air."

-Johannes Kepler

"Our ancestors were eager to understand the world but had not quite stumbled upon the method."

- Carl Sagan

Isaac Newton (1642 – 1727)



The miracle years: 1665-1666

The Principia: 1686

$$\begin{aligned} \frac{dz}{dt} &= \frac{\partial z}{\partial x} \frac{dx}{dt} + \frac{\partial z}{\partial y} \frac{dy}{dt}, \\ \frac{dy}{dx} &= \frac{dy}{du} \cdot \frac{du}{dx}, \\ f_1(u_1, \dots, u_p) \end{aligned}$$
$$\int_a^b f(x) \, dx &= F(b) - F(a), \\ \int_g^f f(z) \, dz &= F(z(\beta)) - F(z(\alpha)). \end{aligned}$$

Newton's First Law A body at rest remains at rest unless acted upon by an outside force...



Newton's First Law cont'd

... a body in motion remains in motion moving in a straight line at constant speed unless acted upon by an outside force.



Inertia:

An object's natural tendency to resist changes in motion.

Mass:

A measure of the amount of material that makes up an object.

<u>Weight</u>



A measure of the gravitational force between two bodies.



W=mg W: weight m: mass

g: gravitational acceleration

Mass

A measure of the amount of an object's inertia



The rate at which something moves.



Velocity

Speed + Direction

miles Northbound hour

Acceleration:

in velocity change

change in time



Is it possible to change your velocity without changing your speed?



Is is possible to change your velocity without changing your speed?

Sun

Is is possible to change your velocity without changing your speed?

Sun

Is is possible to change your velocity without changing your speed?

Sun

Newton's Second Law

The acceleration of an object is directly proportional to the applied force and inversely proportional to its mass.

ma



Newton's Third Law

For every force there is an <u>equal</u> but <u>opposite</u> force.











Newton's laws lead to the <u>Universal Law of Gravitation:</u>



F = force of gravity G = Universal Gravitational Constant 6.67×10^{-11} Newton m²/ kg² m₁, m₂ = masses of the two bodies r = distance between the two bodies

small masses = small force

large masses = large force

 Gm_1m_2





Measuring Newton's Constant G



$$F = G \frac{M_a M_b}{r^2}$$



Q: Do all objects fall at the same rate?

n Ancient Greeks

NO!

n Galileo

YES!

n A prediction is made by Newtonian Mechanics...

Assume a large mass (M) and a small mass (m) the acceleration due to the force of gravity (g): **Recall:** F = m awhere a = acceleration due to gravity (g)F = m gSo: $\frac{Gm_1m_2}{2}$ **Recall:** F

Gm object M Earth F m_{object} g 2 g **G**M Earth g 2 The acceleration due to gravity is independent of small body's mass!

Orbiting Bodies








Orbital velocity:

Minimum orbital speed:

17,500 miles per hour

5 miles per second

Mach 25!







Is a "weightless" astronaut <u>really</u> weightless?







 $GM_{earth}m_{astr}$ Gm_1m_2 2 Forbit r^2 r_{orbit} $GM_{earth}m_{astr}$ Gm_1m_2 F ground **r**2 2 *r*_{earth} 2 $GM_{earth}m_{astr} \times$ *r_{earth}* $GM_{earth}m_{astr}$ r_{orbit}²

Forbit *r*_{earth} Fground *r*orbit $(6370 km)^2$ ≈ 89% $(6770 km)^2$

Lets return to Kepler's 3rd Law

Kepler's 3rd Law:

"The squares of the sidereal periods of the planets are proportional to the cubes of their semi-major axes."

$\mathbf{P}^2 = \mathbf{d}^3$

n P = Orbital Period measured in Earth years n d = Orbital distance measured in A.U.'s n Example: Jupiter P = 11.86 years $P^2 = 140.6$ d = 5.2 A.U. $d^3 = 140.6$









Hubble back in business



But how Universal is the Law of Gravity

could heliocentric model plus Kepler's laws predict better than geocentric model

> Solar System travel Venus Transit

Star Cluster orbit





















VENUS Transit

1.04

Twice in a Lifetime!

EARTH







SUN

TRANSIT DATES •December 1631 •December 1639 •June 1761 •June 1769 •December 1874 •December 1882 •June 2004 •June 2012 •December 2117 •December 2125



Pierre Gassendi (1592-1655)



n There are no records to suggest that anyone ever observed this event.
n Kepler predicted it would not be visible in Europe, so he requested that mariners keep a lookout for it.
n Pierre Gassendi (1592-1655) tries and fails to observe it

Joannis Kepler (1571-1630)


















The Nature of Light

 Astronomy is observational not experimental (in general)
All things in nature radiate energy as light.

If we can understand the nature of light, then we can understand the nature of the objects emitting the light.

n What is light?

n How does light behave?

n What can we learn from light?



What is the <u>speed</u> of light?

Ole Roemer (1676)









The Speed of Light n Accurately measured in a vacuum: 186,282 miles per second! 11 million miles per minute 671 million miles per hour 5.9 trillion miles per year

The Speed of Light

n Light's finite speed has important and bizarre consequences.

n It takes time for light to travel a given distance.

Moon: 234,000 miles 1.25 seconds

Sun: 93 million miles 8 minutes 19 seconds

Jupiter: 400 million miles 36 minutes

Betelgeuse: 427 years Rigel: 773 years Orion Nebula: 1600 years

Betelgeuse

Alnilam Mintaka Alnitak

Rigel

M51 galaxy: 23 million years

Light year

The distance a beam of light will travel in one years time.

5.9 **T**rillion miles 5,900,000,000,000 miles Betelgeuse: 427 <u>light years</u> Rigel: 773 <u>light years</u> Orion Nebula: 1600 <u>light years</u>

Betelgeuse

Alnilam Mintaka Alnitak

Rigel

M51 galaxy: 23 million *light years*

Thus, looking into space is to travel in a time machine

