## THE HISTORY OF ASTRONOMY

"It is therefore impossible that reason not previously instructed should imagine anything other than that the Earth is a kind of vast house with the vault of the sky placed on top of it; it is motionless and within it the Sun being so small passes from one region to another, like a bird wandering through the air."
-J ohannes Kepler
"Our ancestors were eager to understand the world but had not quite ștumbled upon the method.".

- Carl Sagan

Isaac Newton (1642-1727)



The Principia: 1686
The miracle years: 1665-1666

$$
\frac{d z}{d t}=\frac{\partial z}{\partial x} \frac{d x}{d t}+\frac{\partial z}{\partial y} \frac{d y}{d t}
$$

$$
\frac{d y}{d x}=\frac{d y}{d u} \cdot \frac{d u}{d x}
$$

$$
f_{1}\left(u_{1}, \ldots, u_{p}\right)
$$

$$
\int_{a}^{b} f(x) d x=F(b)-F(a)
$$

$$
\int_{\gamma} f(z) d z=F(z(\beta))-F(z(\alpha))
$$

## Newton's First Law

A body at rest remains at rest unless acted upon by an outside force...
$\%$


## Newton's First Law cont'd

... a body in motion remains in motion moving in a straight line at constant speed unless acted upon by an outside force.



## Inertiar

An object's natural tendency to resist changes in motion.

## Mass:

A measure of the amount of material that makes up an object.

## Weight


$\therefore$ A measure of the gravitational force between two bodies.

## $\mathbf{W}=\mathbf{m g}$ <br> W: weight m: mass


g : gravitational accêelération

## Mass

A measure of the amount of an object's inertia

## Speed:

The rate at which something moves.
$\therefore \frac{\text { miles }}{\text { hour }}$


## Velocity

## Speed + Direction

miles
hour
Northbound .

## Acceleration:

change in velocity
change in time


# Is it possible to change your velocity without changing your speed? 



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## Newton's Second Law

The acceleration of an object is directly proportional to the applied force and inversely proportional to its'mass.


## Newton's Third Law

For every force there is an equal but opposite force.






Gravity is a contributing factor in nearly 73 percent of all accidents involving falling objects.

And yet the so-called
'federal govermment' does nothing!

Newton's laws lead to the Universal Law of Gravitation:

## $F=\frac{G m_{1} m_{2}}{r^{2}}$

F = force of gravity
$\mathrm{G}=$ Universal Gravitational Constant
$\because$
$6.67 \times 10^{-11}$ Newton"m²/ $\mathrm{kg}^{2}$
$m_{1}, m_{2}=$ masses of the two bodies
$r=$ distance between the two bodies

## small masses = small force





Distance (r)

Measuring Newton's Constant G



## Q: Do all objects fall at the same rate?

n Ancient Greeks

## NO!

n Galileo
YES!
n A prediction is made by Newtonian Mechanics...

Assume a large mass (M) and a small mass (m)
the acceleration due to the force of gravity $(\mathrm{g})$ :

Recall:

$$
F=m a
$$

where $a=$ acceleration due to gravity $(\mathbf{g})$

So: $\quad F=\mathbf{m} \mathbf{g}$

Recall:

$$
F_{g}=\frac{G m_{1} m_{2}}{r^{2}}
$$

$$
F_{g}=\frac{G m_{\text {object }} M_{\text {Earth }}}{r^{2}}=m_{\text {object }} g
$$

## GM <br> $\frac{\text { Earth }}{2}=g$



The acceleration due to gravity is independent of . small body's mass!

# Orbiting Bodies 



$$
0
$$




## Orbital velocity:

Minimum orbital speed:

## 17,500 miles per hour

5 miles per second
Mach 25!


## Motion if there

 is no gravitys


# Is a "weightless" astronaut really weightless? 







$$
\frac{(6370 \mathrm{~km})^{2}}{(6770 \mathrm{~km})^{2}}
$$

## Lets return to Kepler's $3^{\text {rd }}$ Law

## Kepler's $3^{\text {rd }}$ Law:

"The squares of the sidereal periods of the planets are proportional to the cubes of their semi-major axes."

## $P^{2}=d^{3}$

${ }_{n} \mathrm{P}=$ Orbital Period measured in Earth years
${ }^{n} \mathrm{~d}=$ Orbital distance measured in A.U.'s
n Example: . . Jupiter
$P=11.86$ years
$P^{2}=140.6$
$\mathrm{d}=5.2$ A.U.
$d^{3}=140.6$

S.o we häve....

$$
\mathrm{F}=\frac{\mathrm{GMm}}{\mathrm{r}^{2}} \quad \mathrm{~F}=\frac{\mathrm{mv}^{2}}{\mathrm{r}^{2}} \quad \mathrm{v}=\frac{2 \pi \mathrm{r}}{\mathrm{P}}
$$

$\frac{\mathrm{GMm}}{\mathrm{r}^{2}}=\frac{\mathrm{m} v^{2}}{\mathrm{r}} \longrightarrow \frac{\mathrm{GMm}}{\mathrm{r}^{2} \mathrm{~m}}=\mathrm{v}^{2} \longrightarrow \quad \frac{\mathrm{GM}}{\mathrm{r}}=\mathrm{v}^{2}$

$$
\left(\mathrm{y}=\frac{2 \pi \mathrm{r}}{\mathrm{P}}\right)^{2} \quad \because \mathrm{v}^{2}=\frac{4 \pi^{2} \mathrm{r}^{2}}{\mathrm{P}^{2}} \longrightarrow \frac{\mathrm{GM}}{\mathrm{r}}=\frac{4 \pi^{2} \mathrm{r}^{2}}{\mathrm{P}^{2}} \longrightarrow
$$

$$
\Rightarrow \quad \frac{4 \pi^{2}}{G} \cdot \frac{r^{3}}{P^{2}}=M
$$

Hubble back in business


Hubble back in business
:

## But how Universal is the Law of Gravity

could heliocentric model plus Kepler's laws predict better than geocentric model

- Solar System travel - Venus Transit
- Star Cluster orbit

Voyager $1(118 \mathrm{AU})$ and

- 2(96 AU): (32:50, 28:38) 1977-20011 and going strong












## TRANSIT DATES - December 1631 <br> - December 1639

- June 1761
- June 1769
-December 1874
- December 1882
- June 2004
- June 2012
- December 2117
- December 2125


## December 6, 1631.




J oannis Kepler
(1571-1630)


## It's All in the Geometry






## © Anglo-Australian Observatory

## M31, the Great Andromeda Galaxy

You are here



## The Nature of Light

$\stackrel{*}{*}$
n Astronomy is observational not experimental (in general)
n All things in nature radiate energy as light.
n. If we can understand the nature of light, then we can understand $\therefore$ the nature of the objects emitting the light.

## n What is light?

n How does light behave?
n What can we learn from light?

## What is the speed of light?



## Ole Roemer (1676)




Ganymede


The Speed of Light
n Áccurately measured in a vacuum:

## 186,282 miles per second!

11 million miles per minute
671 million miles per hour
5.9 trillion miles per year

## The Speed of Light

n Light's finite speed has important and. bizarre consequences.
n It takes time for light to travel a given distance.

## Moon:

- 234,000 miles 1.25 seconds


## Sun:

## 93 million miles

8 minutes 19 seconds

## J upiter:

400 million miles
36 minutes

## Betelgeuise: <br> -427 years <br> Betelgeuse

## Rigel:

773 years
Orion Nebula:
1600 years


## M51 galaxy: <br> 23 million years

## Light year

The distance a beam of light will travel in one years time.
-.. 5.9 Irillion miles

## 5,900,000,000,000 miles

## Betelgeuise:

## 427 light years <br> Betelgeuse

## Rigel:

## 773 light years

Orion Nebula:
1600 light years
Alnilam Mintaka
Alnitak

Rigel

M51 galaxy:
23 million light years


Thus, looking into space is to travel in a time machine


